



## Dark Matter @ Colliders

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**BSM lectures @ NPAC** 

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Dark Matter @ Colliders

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### Outline



## New physics and dark matter



### Dark matter in cosmology and at colliders

Dark matter is searched for directly, indirectly and at colliders
 This huge experimental effort offers a strategy to constrain models



#### From dark matter to missing transverse energy



- - $\star$  Larger  $\sqrt{s} >$  heavier new particles
- Unknown longitudinal momenta
  - $\star$  Use of quantities invariant under longitudinal boosts ( $p_T$ , etc.)



#### Summary

## From dark matter to missing transverse energy



#### Summary

## From dark matter to missing transverse energy



- \*Unknown *partonic* centre-of-mass energy  $\sqrt{s}$ 
  - $\star$  Larger  $\sqrt{s} >$  heavier new particles
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  - $\star$  Use of quantities invariant under longitudinal boosts ( $p_T$ , etc.)



#### Energy-momentum conservation (in the transverse plane)

- The initial-state total transverse momentum is zero
  - $\succ$  the final-state total pT is zero
- Invisible particles (DM in particular) = missing momentum
  - $\star$  Weakly interacting and neutral  $\succ$  detector is transparent
  - ★ Presence inferred from momentum imbalance

$$\mathbf{E}_T = ||\mathbf{p}_T|| = \left|\left|-\sum_{\text{visible}}\mathbf{p}_T\right|\right|$$

**\***Beware: MET  $\neq$  DM

- **★** MET could originate from neutral long-lived states, or even neutrinos
- **★** DM may not yield large MET (if light or from a compressed spectrum)

How to detect missing energy?

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# Detecting missing energy at colliders (I)



# Detecting missing energy at colliders (2)



# Detecting missing energy at colliders (3)



## Dark matter signatures at the LHC



## Dark matter signatures at the LHC



# Dark matter signatures at the LHC



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#### Almost two decades of mono-X searches...

| The mono-X (DM) story is almost 20 years old                    | · · · · · · · · · · · · · · · · · · ·     |
|---|---|
| The problem was to trigger on DM signals $\rightarrow$ need for | a visible object                          |
| Introduced first as mono-photons in lepton collisions           | [Birkedal, Matchev & Perelstein (PRD`04)] |
| Extension to mono-jets in hadron collisions                     | [Feng, Su & Takayama (PRL`06)]            |

### Almost two decades of mono-X searches...

![](_page_14_Figure_4.jpeg)

## Almost two decades of mono-X searches...

![](_page_15_Figure_4.jpeg)

#### Towards the modern epoch

New dark signals: mono-top, mono-Z, mono-lepton & mono-Higgs
[Andrea, BF & Maltoni (PRD`II); Bell, Dent, Galea, Jacques, Krauss & Weiler (PRD`I2); Bai & Tait (PLB`I3); Petrov & Shepherd (PLB`I4)]

First experimental studies: CDF, and then ATLAS/CMS

## A dark matter search strategy at the LHC

A typical LHC dark matter search strategy
 Requirement of a significant amount of missing transverse energy
 Requirement of a significantly hard visible object (jet, di-lepton pair, photon, etc.)
 Extra constraints (angular correlations, vetoes, etc.) to reduce the backgrounds
 Cut and count and looking for excesses over the background

## A dark matter search strategy at the LHC

#### A typical LHC dark matter search strategy

- Requirement of a significant amount of missing transverse energy
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- Cut and count and looking for excesses over the background

#### Backgrounds - the mono-jet case

Invisible Z decays

 → irreducible backgrounds

 W decays with a lost lepton

 → not very frequent but large (total) rate
 → Mis-measurements in multi-jet production
 → rare, but huge QCD total rate
 → steeply falling with the MET value

![](_page_17_Figure_11.jpeg)

### Outline

![](_page_18_Figure_4.jpeg)

#### An EFT interpretation

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## **Connecting direct detection and colliders**

![](_page_20_Figure_4.jpeg)

![](_page_20_Picture_5.jpeg)

![](_page_20_Picture_6.jpeg)

![](_page_20_Figure_7.jpeg)

## **Connecting direct detection and colliders**

![](_page_21_Figure_4.jpeg)

### **DM direct detection in a nutshell**

[ Drees & Nojiri (PRD`93); Hisano, Nagai & Nagata (JHEP`I5) ]

![](_page_22_Figure_5.jpeg)

## DM direct detection in a nutshell

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![](_page_23_Figure_5.jpeg)

## DM direct detection in a nutshell

[ Drees & Nojiri (PRD`93); Hisano, Nagai & Nagata (JHEP`I5) ]

![](_page_24_Figure_5.jpeg)

#### **Complementary constraints**

![](_page_25_Figure_4.jpeg)

#### **Complementary constraints**

![](_page_26_Figure_4.jpeg)

## The failure of the EFT interpretation

![](_page_27_Figure_4.jpeg)

# The failure of the EFT interpretation

![](_page_28_Figure_4.jpeg)

## The failure of the EFT interpretation

![](_page_29_Figure_4.jpeg)

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Dark Matter simplified models

The s-channel case

# From EFT to simplified model interpretations

![](_page_31_Figure_4.jpeg)

## From EFT to simplified model interpretations

![](_page_32_Figure_4.jpeg)

## Simplified dark matter models @ LHC

![](_page_33_Figure_4.jpeg)

# Simplified dark matter models @ LHC

![](_page_34_Figure_4.jpeg)

#### s-channel models at colliders

![](_page_35_Figure_4.jpeg)

#### s-channel models at colliders

![](_page_36_Figure_4.jpeg)

## Example: top-philic fermion DM / scalar mediator

![](_page_37_Figure_4.jpeg)

## Example: top-philic fermion DM / scalar mediator

![](_page_38_Figure_4.jpeg)

[Arina, Backovic, Conte, BF, Guo, Heisig, Hespel, Krämer, Maltoni, Martini, Mawatari, Pellen & Vryonidou (JHEP'16)]

![](_page_39_Figure_5.jpeg)

[Arina, Backovic, Conte, BF, Guo, Heisig, Hespel, Krämer, Maltoni, Martini, Mawatari, Pellen & Vryonidou (JHEP'16)]

![](_page_40_Figure_5.jpeg)

[Arina, Backovic, Conte, BF, Guo, Heisig, Hespel, Krämer, Maltoni, Martini, Mawatari, Pellen & Vryonidou (JHEP'16)]

![](_page_41_Figure_5.jpeg)

[ Arina, Backovic, Conte, BF, Guo, Heisig, Hespel, Krämer, Maltoni, Martini, Mawatari, Pellen & Vryonidou (JHEP'16) ]

![](_page_42_Figure_5.jpeg)

1.00

0.75

0.50

0.25

0.00

-0.25

-0.50

-0.75

-1.00

 $\log_{10}(g_t)$ 

[ Arina, Backovic, Conte, BF, Guo, Heisig, Hespel, Krämer, Maltoni, Martini, Mawatari, Pellen & Vryonidou (JHEP'16) ]

![](_page_43_Figure_5.jpeg)

## **Details on the simulations**

◆ Several key aspects behind the previous results
 ◆ Simulations at the NLO-QCD accuracy → precision predictions
 ◆ Recasting with public tools
 ★ Many ATLAS and CMS searches for new physics
 ★ Interpretation within popular frameworks and simplified models
 ★ Need for interpretations in all kind of models

Dark matter simplified models

The t-channel case

#### t-channel models at colliders

![](_page_46_Figure_4.jpeg)

![](_page_46_Figure_5.jpeg)

SM

### t-channel models at colliders

![](_page_47_Figure_4.jpeg)

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## Recasting ATLAS mono-jet search (36/fb)

#### CLs exclusion from the best region (I TeV mediator; I 50 GeV DM)

| Process                         | $CL_s$ [LO]              | $E_T^{\text{miss}}$ constraint | $CL_s$ [NLO]           | $E_T^{\text{miss}}$ constrtaint |
|---------------------------------|--------------------------|--------------------------------|------------------------|---------------------------------|
| Total                           | $75.6^{+10.1}_{-10.5}$ % | $\in [700,800]~{\rm GeV}$      | $97.8^{+0.9}_{-1.4}$ % | $\geq 700 {\rm ~GeV}$           |
| XX                              | $0.7^{+0.6}_{-0.6}$ %    | $\in [250, 300]~{\rm GeV}$     | $3.6^{+0.3}_{-0.6}$ %  | $\geq 900~{\rm GeV}$            |
| XY                              | $62.7^{+12.3}_{-10.4}$ % | $\in [500,600]~{\rm GeV}$      | $83.9^{+2.9}_{-4.3}$ % | $\in [700,800]~{\rm GeV}$       |
| YY [total]                      | $24.0^{+3.1}_{-3.1}$ %   | $\geq 900~{\rm GeV}$           | $58.1^{+2.2}_{-3.1}$ % | $\geq 900 {\rm ~GeV}$           |
| YY [QCD]                        | $10.7^{+4.4}_{-2.6}$ %   | $\geq 900~{\rm GeV}$           | $17.0^{+2.1}_{-2.1}$ % | $\geq 900~{\rm GeV}$            |
| YY [t-channel]                  | $29.6^{+3.3}_{-2.6}$ %   | $\geq 900~{\rm GeV}$           | $38.9^{+1.2}_{-1.8}$ % | $\geq 900~{\rm GeV}$            |
| [ Arina, BF & Mantani (EPJC`20) |                          |                                |                        |                                 |

#### NLO simulations are crucial

- \* Modification of the rates (larger yields) and shapes (different best region)
- $\star$  Better control of the theory errors
- Considering all signal components is crucial
  - $\star$  One component alone is not sufficient to exclude the scenario

#### Ist gen. mediator & Majorana DM

![](_page_49_Figure_4.jpeg)

#### Ist gen. mediator & Majorana DM

![](_page_50_Figure_4.jpeg)

### More strongly coupled dark matter

![](_page_51_Figure_4.jpeg)

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#### More strongly coupled dark matter

![](_page_52_Figure_4.jpeg)

#### $\blacklozenge \lambda = 5$

- \*All channels contribute (larger rates)  $\star XX \sim \lambda^4$   $\star XY \sim \lambda^2$  $\star YY \sim \lambda^4 + \lambda^2 + \lambda^0$
- Simulations unreliable
  - ★The NWA breaks down
  - $\star \Gamma_Y/M_Y > 10\%$  or compressed spectrum
  - ★Most 'excluded' points inconclusive
- \*  $\Gamma_Y$  plays a role for large  $\lambda$  values

#### More strongly coupled dark matter

![](_page_53_Figure_4.jpeg)

#### $\lambda = 5$

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  - $\star \Gamma_Y/M_Y > 10\%$  or compressed spectrum
  - ★Most 'excluded' points inconclusive
- $\Gamma_Y$  plays a role for large  $\lambda$  values
- Sensitivity to all channels
  - Different jet properties
    - $\rightarrow$  XX: small N<sub>j</sub>, mostly soft jets
    - $\rightarrow$  XY: medium N<sub>j</sub>, hard and softer jets
    - $\rightarrow$  YY: large N<sub>j</sub>, hard jets
  - Dedicated regions for all cases

#### Fixed coupling vs fixed width

[Arina, BF, Mantani, Mies, Panizzi & Salko (PLB`20)]

![](_page_54_Figure_5.jpeg)

#### $\lambda = 2 \text{ vs } \Gamma_Y/M_Y = 5\%$

Signal = XX + XY + YY

Regions with 2 very hard jets (SR2j) ~YY production and decay

✤Regions with more not so hard jets (SR4j, SR5j, SR6j) ~ compressed regime

#### Reliability of the simulations

 $\star$  Fixed  $\Gamma_Y/M_Y$ : compressed spectrum = non-perturbative regime

 $\star$  Fixed  $\lambda$ : split spectrum = broad mediator

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★ Fixed  $\Gamma_Y/M_Y$ : compressed spectrum = non-perturbativ Care to be taken with barogue setups

**★** Fixed  $\lambda$ : split spectrum = broad mediator

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#### Beyond simplified models

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#### **Benchmark models**

#### More complex than simplified models

\*A large set of models exist  $\rightarrow$  focus on few benchmarks (e.g. supersymmetry)

- Dedicated searches for those specific benchmarks
  - $\star$  Simplified models encapsulate characteristics of varied theories

#### ✦ 3 (subjective) examples

- The Higgs portal model (very few parameters and one new state)
- Dilaton-induced DM (very few parameters and two new states)
- Supersymmetry (lots of parameters and new states)

There are many more: dark photons, axions, etc. (not covered here)

# I. The Higgs portal

![](_page_58_Figure_4.jpeg)

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![](_page_59_Figure_4.jpeg)

## 2. Dilaton induced DM

![](_page_60_Figure_4.jpeg)

## 2. Dilaton induced DM

![](_page_61_Figure_4.jpeg)

# 3. Supersymmetry (I)

![](_page_62_Figure_4.jpeg)

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![](_page_63_Figure_4.jpeg)

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![](_page_64_Figure_4.jpeg)

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## 3. Supersymmetry (2)

![](_page_65_Figure_4.jpeg)

Summary

## 3. Supersymmetry (2)

![](_page_66_Figure_4.jpeg)

### Outline

![](_page_67_Figure_4.jpeg)

## New physics and dark matter

Dark matter is an important motivation for new physics \*Galaxy rotation curves, gravitational lensing, cosmic microwave background, ... DM SM Searched for in a complementary way Dark matter relic abundance must be reproduced Dark matter direct/indirect detection constraints Production at (hadron) colliders SM DM Many signatures are considered at the LHC From various benchmarks: simplified models, EFTs, UV-complete models Accurate predictions are necessary for the best conclusions NLO-QCD computations for BSM are automated

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  - NLO-QCD computations for BSM are automated

A lot of fun is planned for the next decades