## Exercises (1)

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## 1 The normalisation of the cosmic ray spectrum

The *Voyager 1* probe measured the intensity of local cosmic ray protons j(E) which can be (very) roughly represented as:

$$j(E) = A\left(\frac{E}{E_0}\right)^{0.1} \qquad E < E_0$$
$$= A\left(\frac{E}{E_0}\right)^{-2.7} \qquad E > E_0$$

where  $E_0 \sim 1$  GeV and A a normalisation constant in units of eV<sup>-1</sup> cm<sup>-2</sup> s<sup>-1</sup> sr<sup>-1</sup>.

Find the value of the normalisation constant A by imposing that the local energy density of cosmic ray protons is  $w_{CR} \sim 1 \text{ eV/cm}^3$ . Show that most of the energy is carried by particles with energy  $E \gtrsim E_0$ .

Finally, estimate the normalisation B (in units of  $eV^{-1} s^{-1}$ ) and the spectral slope  $\alpha$  of the injection spectrum of cosmic rays in the Galactic disk (for  $E > E_0$ ):

$$Q(E) = B\left(\frac{E}{E_0}\right)^{-\alpha} \tag{1}$$

knowing that the residence time of cosmic rays in the Galactic disk scales as  $\tau_c \propto E^{-0.3}$  and that the total cosmic ray power of the galaxy is  $P_{CR} \sim 10^{41}$  erg/s.