

# Lectures outline

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- 1- end of gaseous detectors (MPGD part)**
- 2- trigger example with gaseous detector**
- 3- Calorimeter detectors** (*one word about neutron interaction*)
- 4- trigger with calorimeter det. (and MM gaseous detector)**
- 5- magnets (briefly)**

Thanks to Laurent Serin (LAL) for the calorimetry part.

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# Calorimeter detectors

## Energy measurements



# Outline

**1- Calorimeter definition & history, illustration with some major physics results**

**2- Electromagnetic interaction and shower development**

**3- Electromagnetic calorimeters technologies**

**4-  $e/\gamma$  reconstruction, calibration and performance**

**4- Hadronic shower development**

**5- Hadronic calorimeters technologies**

**6- Jets reconstruction calibration and performance**

**7- Missing transverse energy measurement**

**8- Conclusion**

*Slides categories*

- \* For information*
- \*\* Useful to know*
- \*\*\* Needs to know*

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**Calorimetry definition**

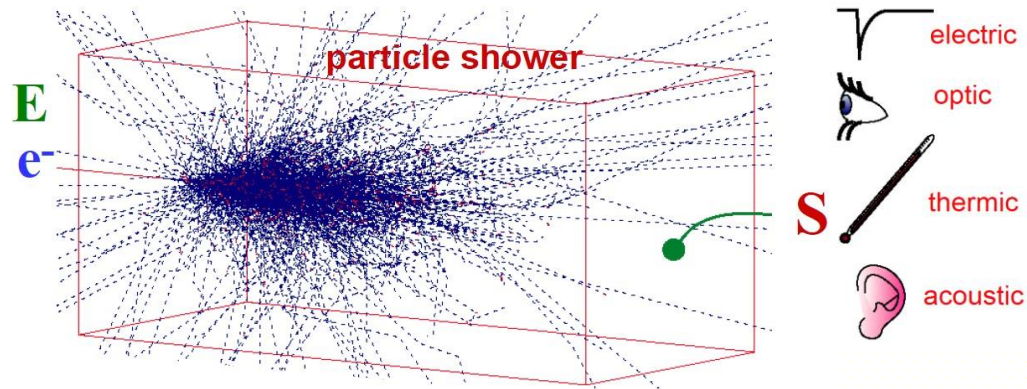
**History of calorimeters**

**Illustration with some major  
physics results**

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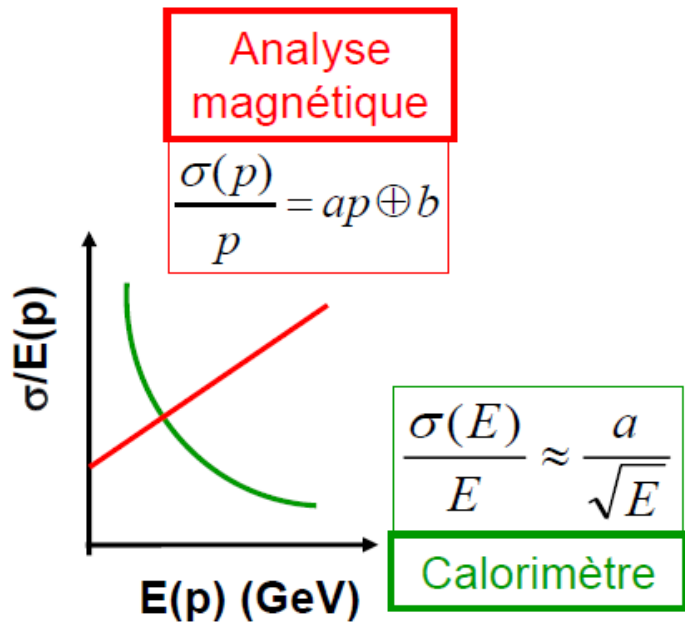
# Calorimetry definition

- Experimental technique used in Nuclear Physics, Particle Physics and Astroparticle to detect a particle and measure some of its properties based on total or partial absorption of the particle in a fiducial volume
- Destructive process :  
Particle is absorbed in the medium or exit it quite modified
- Particle energy is converted in a detectable signal.
- Key element of any High Energy Physics (HEP) experiment



# Calorimeters needed for HEP

Sensitive to **all charged and neutrals** particles in final state  
**Good resolution at high energy**, and “**sizeable**” detectors



Calorimeter shower depth  $\sim \ln E/E_c$   
 almost energy independent  
 → Calorimeter can be compact detector

Magnetic spectrometer :

$$\sigma(p) / p \sim p / (BL^2)$$

→ Detector size has to grow quadratically to maintain resolution

Calorimeter can also provide:

- Position/angular measurement
- Time measurement
- **Trigger**
- Particle identification (e,  $\gamma$ ,  $\pi$ ,  $\mu$ , h...)

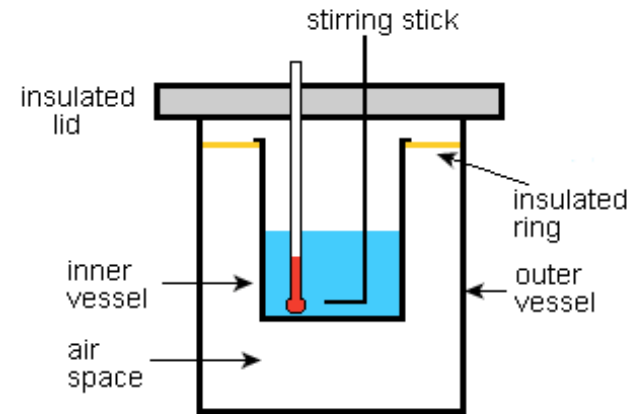
ATLAS : calo  $\sigma(E)/E \sim \frac{10\%}{\sqrt{E}} \oplus 0.7\%$   
 tracking  $\sigma(p) / p \sim 5 \cdot 10^{-4} p \oplus 1\%$

At 40 GeV for electrons similar energy resolution

# Calorimeter History (1)

Calorimetry (calor = heat in latin) is originally a concept used in thermodynamics/chemistry :

- Isolated box with a substance to study
- Exchange of heat measured by temperature variation
- 1 calorie = 4.185 Joule =  $2.6 \cdot 10^7$  TeV
- increases by 1 °C in normal condition 1g of water
- 1 GeV induces a  $\Delta T \sim 4 \cdot 10^{-14}$  K in 1 liter of water



First use in 1878 (Langley) to measure electromagnetic radiation from sun :

- 2 platinum strips, one isolated from radiation, and the second receiving the radiation connected to a Wheatson bridge
- measure Energy/Temperature through resistance change
- 30 % accuracy measurement :  $1.77 \text{ kW/m}^2$  instead of  $1.38 \text{ kW/m}^2$

Orthmann & Meitner (1930) : differential calorimeter used to measure mean energy of electrons in  $^{210}\text{B}$  beta decay :  $E=0.33 \text{ MeV}$  @ 6 %

→ Such calorimeters still used in the field, named “ Bolometers”, used in dark matter experiments (Edelweiss, CDMS....) or Cosmic Microwave Background (Planck) (see M. Charles' lesson)

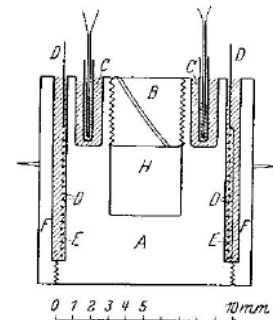
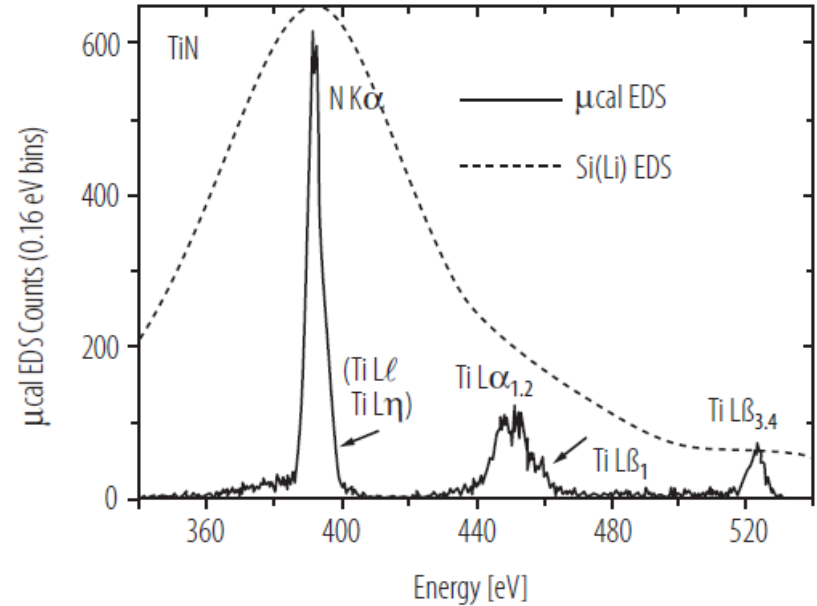
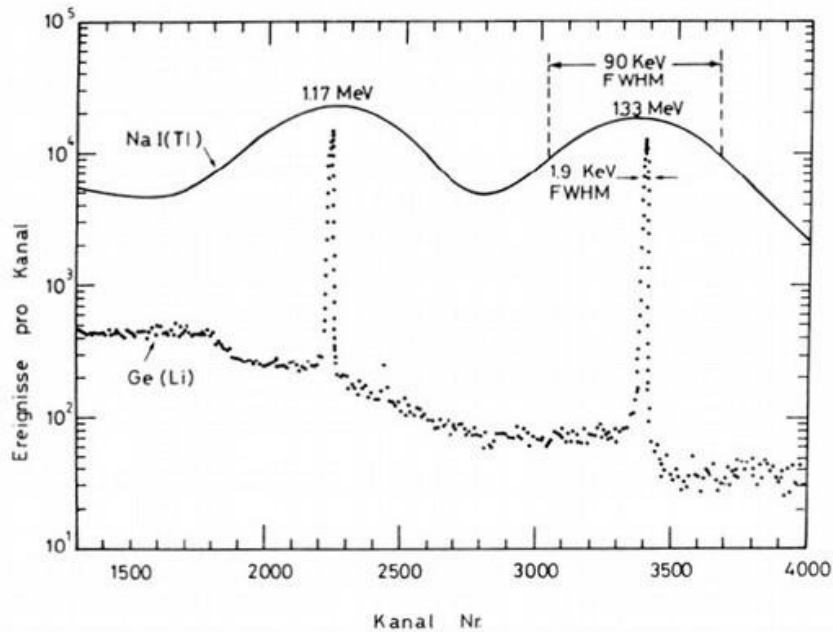


Fig. 1.  
Längsschnitt durch ein  
Kalorimetergefäß.  
Wood'sches Metall.

# Interlude : cryogenic $\mu$ calorimeter

Definitely best energy resolution for very low energy but not the subject of this lesson



At **100 MeV**, solid state (Si, Ge) detectors have  $\sim 25/30$  better resolution than scintillators

At few **hundred eV**, cryogenic bolometer can have 50 better resolution  
Ok for event energy measurement but not individual particle energy measurement



# Calorimeter History (2)

First HEP sampling calorimeter in 1953  
for High Energy cosmic ray particle  $E > 10^{14}$  eV

Sandwich of ionization chambers and scintillation  
counters interleaved with iron :

-visible energy extracted from numbers of secondary  
particle ( $n(x)$ ) and energy loss ionisation and scintillation  
Counters ( $E_{\text{visible}} = dE/dx \int n(x).dx$ )

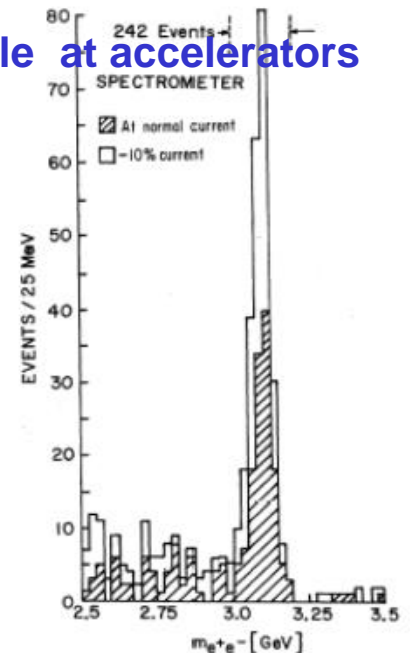
- Fraction of the visible energy lost in absorbers plate

→ **Need to be calibrated with particle of known energy, not yet available at accelerators**

➤ 60-70' accelerators became main facilities for particle physics :

- Need to measure also neutral particles ( $\pi^0$ ,  $\gamma$ , neutrons...)
- Charged particle accurately measured with  
large spectrometers detectors : 10m arms  
for electrons from  $J/\Psi$

→ **Plenty of calorimeters development/technologies**



# Classification of calorimeters

Per particle type

**Electromagnetic calorimeters :**

$e^{+/-}$ ,  $\gamma$  and  $\pi^0$

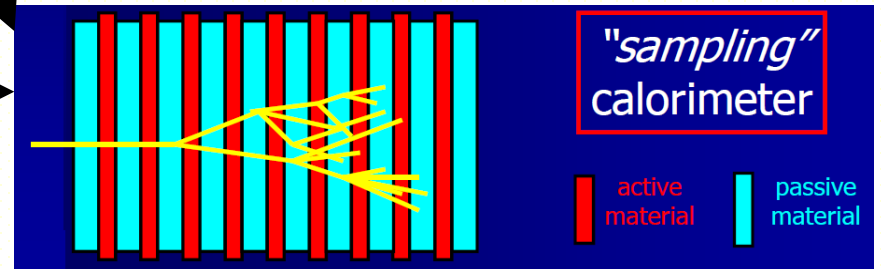
Per construction technique



Full absorption detector, active medium for energy degradation and signal generation

**Hadron calorimeters :**

Charged and neutral hadrons, jets



Alternate layers of absorbers to degrade particle energy and active medium to provide detectable signal

# Classification of calorimeters

## By signal detection technology

Homogeneous Calorimeters	Scintillation/ Crystal
	Semiconductor
	Cherenkov
	Ionization (Noble Liquids)

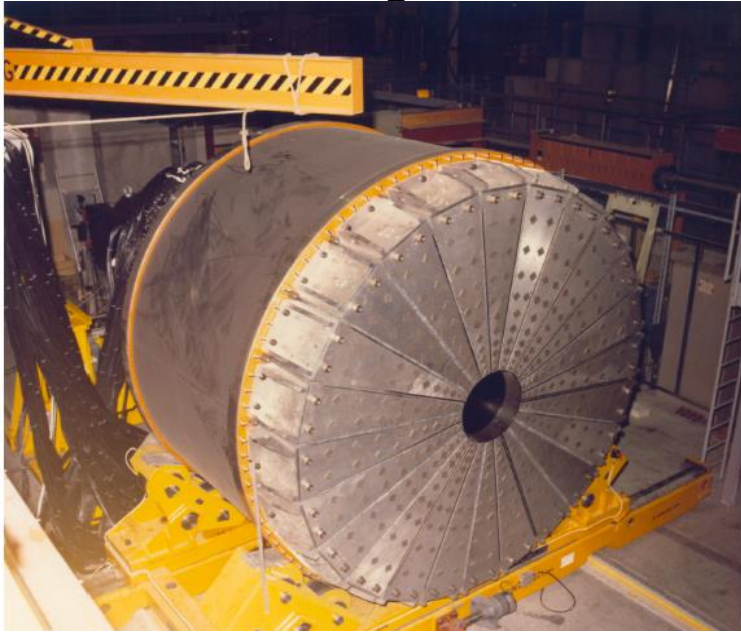
Sampling Calorimeters	Scintillation
	Gas
	Solid State
	Liquids
	<i>Common Absorbers: Pb, Fe, Cu, U, W</i>

## Existing Electromagnetic Calorimeters

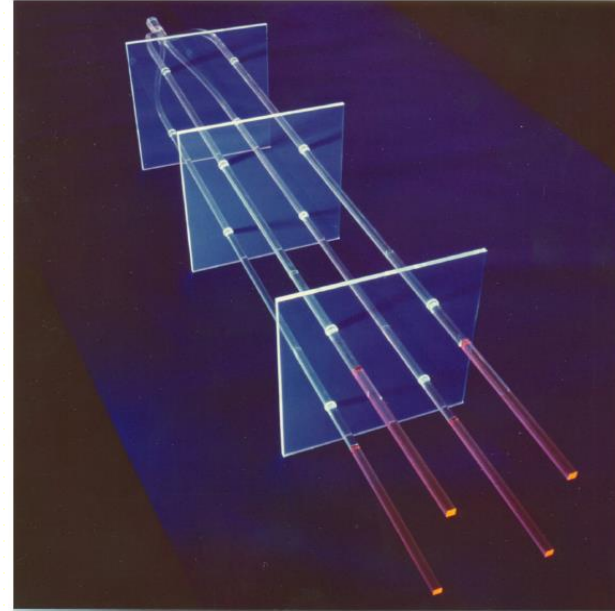
Technology/Experiment	Depth	Resolution	Year
NaI(Tl) (Crystal Ball)	20 $X_0$	2.7%/ $E^{1/4}$	1983
Bi <sub>4</sub> Ge <sub>3</sub> O <sub>12</sub> (BGO) (L3)	22 $X_0$	2%/ $\sqrt{E} \oplus 0.7\%$	1993
CsI (KTeV)	27 $X_0$	2%/ $\sqrt{E} \oplus 0.45\%$	1996
CsI(Tl) (BaBar)	16–18 $X_0$	2.3%/ $E^{1/4} \oplus 1.4\%$	1999
CsI(Tl) (BELLE)	16 $X_0$	1.7% for $E_\gamma > 3.5$ GeV	1998
PbWO <sub>4</sub> (PWO) (CMS)	25 $X_0$	3%/ $\sqrt{E} \oplus 0.5\% \oplus 0.2/E$	1997
Lead glass (OPAL)	20.5 $X_0$	5%/ $\sqrt{E}$	1990
Liquid Kr (NA48)	27 $X_0$	3.2%/ $\sqrt{E} \oplus 0.42\% \oplus 0.09/E$	1998
Scintillator/depleted U (ZEUS)	20–30 $X_0$	18%/ $\sqrt{E}$	1988
Scintillator/Pb (CDF)	18 $X_0$	13.5%/ $\sqrt{E}$	1988
Scintillator fiber/Pb spaghetti (KLOE)	15 $X_0$	5.7%/ $\sqrt{E} \oplus 0.6\%$	1995
Liquid Ar/Pb (NA31)	27 $X_0$	7.5%/ $\sqrt{E} \oplus 0.5\% \oplus 0.1/E$	1988
Liquid Ar/Pb (SLD)	21 $X_0$	8%/ $\sqrt{E}$	1993
Liquid Ar/Pb (H1)	20–30 $X_0$	12%/ $\sqrt{E} \oplus 1\%$	1998
Liquid Ar/depl. U (DØ)	20.5 $X_0$	16%/ $\sqrt{E} \oplus 0.3\% \oplus 0.3/E$	1993
Liquid Ar/Pb accordion (ATLAS)	25 $X_0$	10%/ $\sqrt{E} \oplus 0.4\% \oplus 0.3/E$	1996

# Example of calorimeters

Fixed target calorimeters : NA5 at CERN (1978) QCD measurements  
One of the first segmented calorimeter



24 ( $\varphi$ ) x 10 ( $\theta$ ) cells  
EM section : Scintillator/Pb  
Had section : Scintillator/Fe  
using two different Wave Length  
Shifter (WLS)



Main idea : guide the light of both  
section in single rod read by two  
PM behind yellow (EM) and  
green (Had) filters

# Example of calorimeters

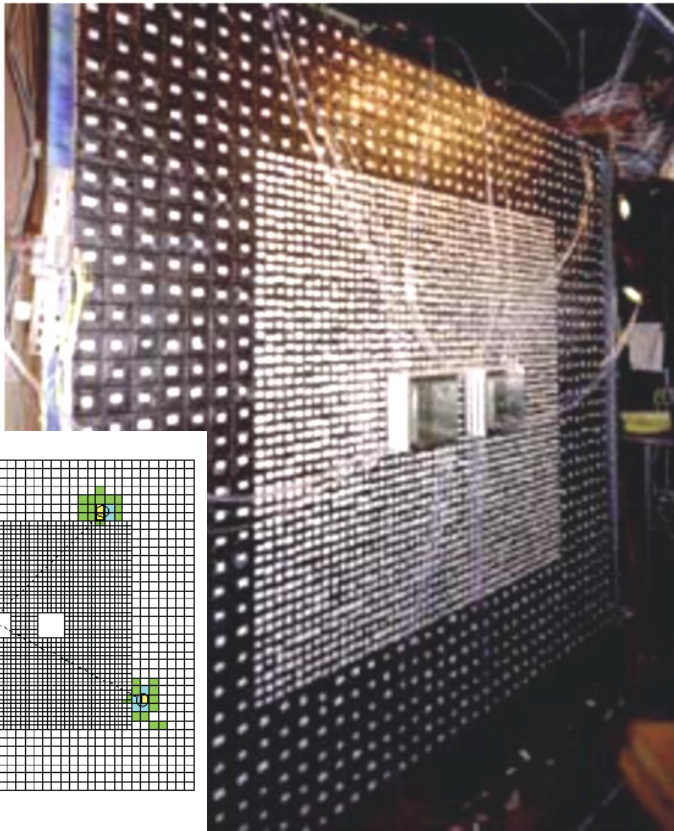
**Fixed target calorimeters** : CP violation in K decays experiment : NA31 / NA48 / KTeV

→ Need to measure accurately  $K_L \rightarrow \pi^0 \pi^0 \rightarrow 4\gamma$  ( $\text{Br}(K_L \rightarrow 3\pi^0)/(\text{Br}(K_L \rightarrow 2\pi^0)) \sim 300$ )

→ Shower separation + invariant mass : fine granularity and energy resolution

→ Homogeneous calorimeters

KTeV 3100 pure CsI crystals



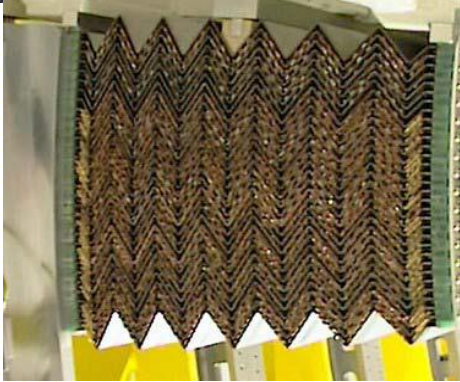
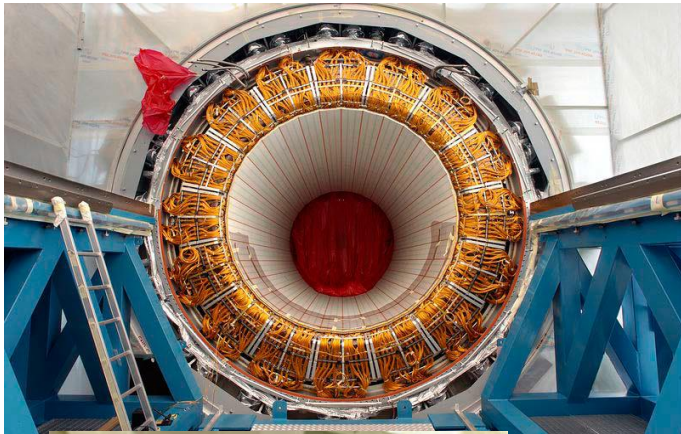
Liquid Krypton calorimeters,  
still in used in K experiments at CERN



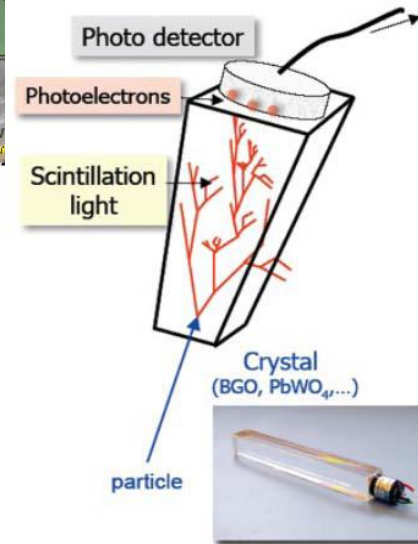
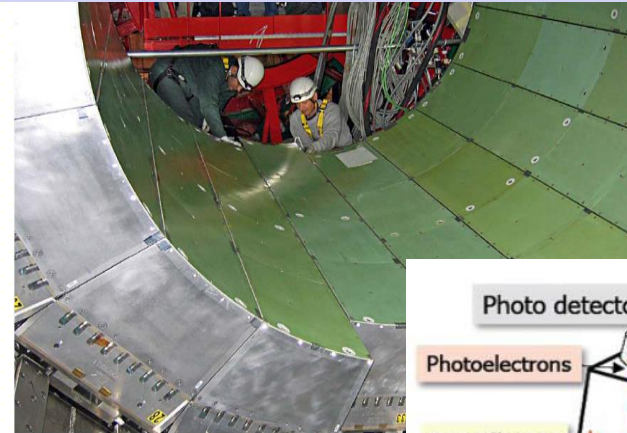
# LHC calorimeters

LHC electromagnetic calorimeters, two different approaches

ATLAS : Liquid Argon / Lead sampling electromagnetic calorimeter



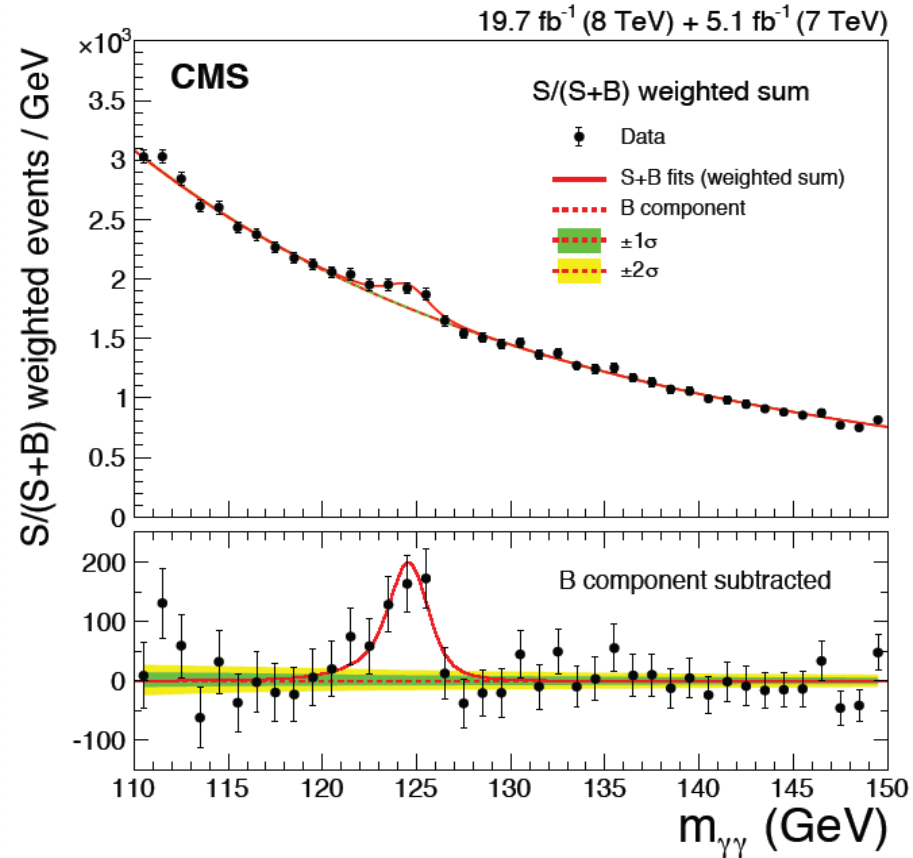
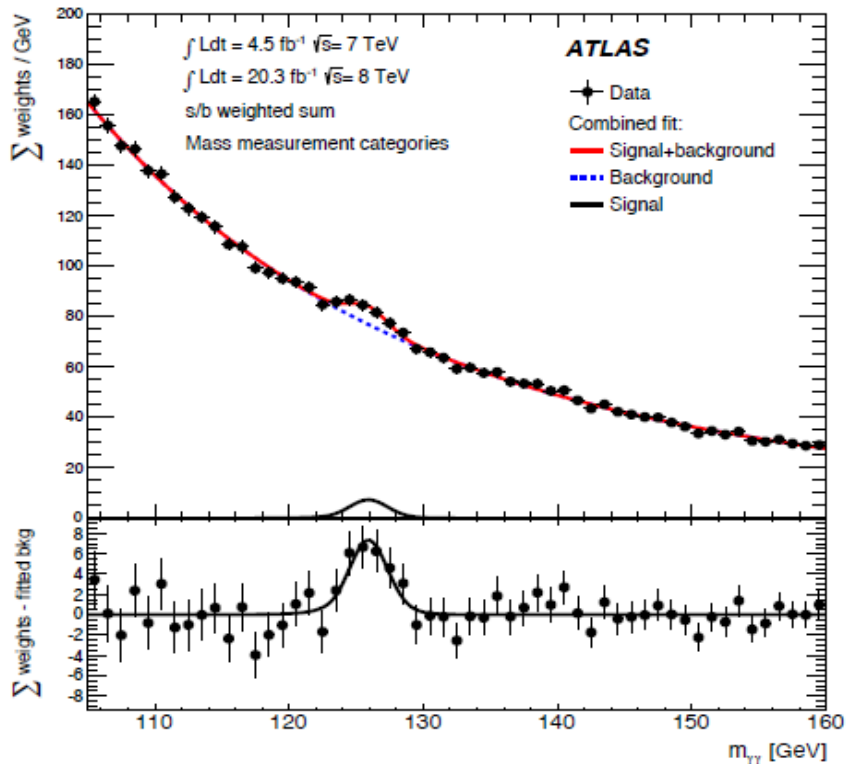
CMS : Homogeneous calorimeter PbWO<sub>4</sub> crystals



Key detector of Higgs discovery

ATLAS calorimeter better than CMS or CMS better than ATLAS ?

# H → γγ in ATLAS and CMS



Quite a similar result with different detectors (similar  $S/\sqrt{B}$ , scales as  $1/\sigma_{(masse)}$ )

CMS : Energy resolution

ATLAS : Granularity (jet rejection) and angular resolution

$$m_{\gamma\gamma} = 2 E_1 E_2 (1 - \cos(\theta))$$

$$\frac{\Delta m_{\gamma\gamma}}{m_{\gamma\gamma}} = \frac{1}{2} \left( \frac{\Delta E_1}{E_1} \oplus \frac{\Delta E_2}{E_2} \oplus \frac{\Delta \theta_{\gamma\gamma}}{\tan(\theta_{\gamma\gamma}/2)} \right)$$

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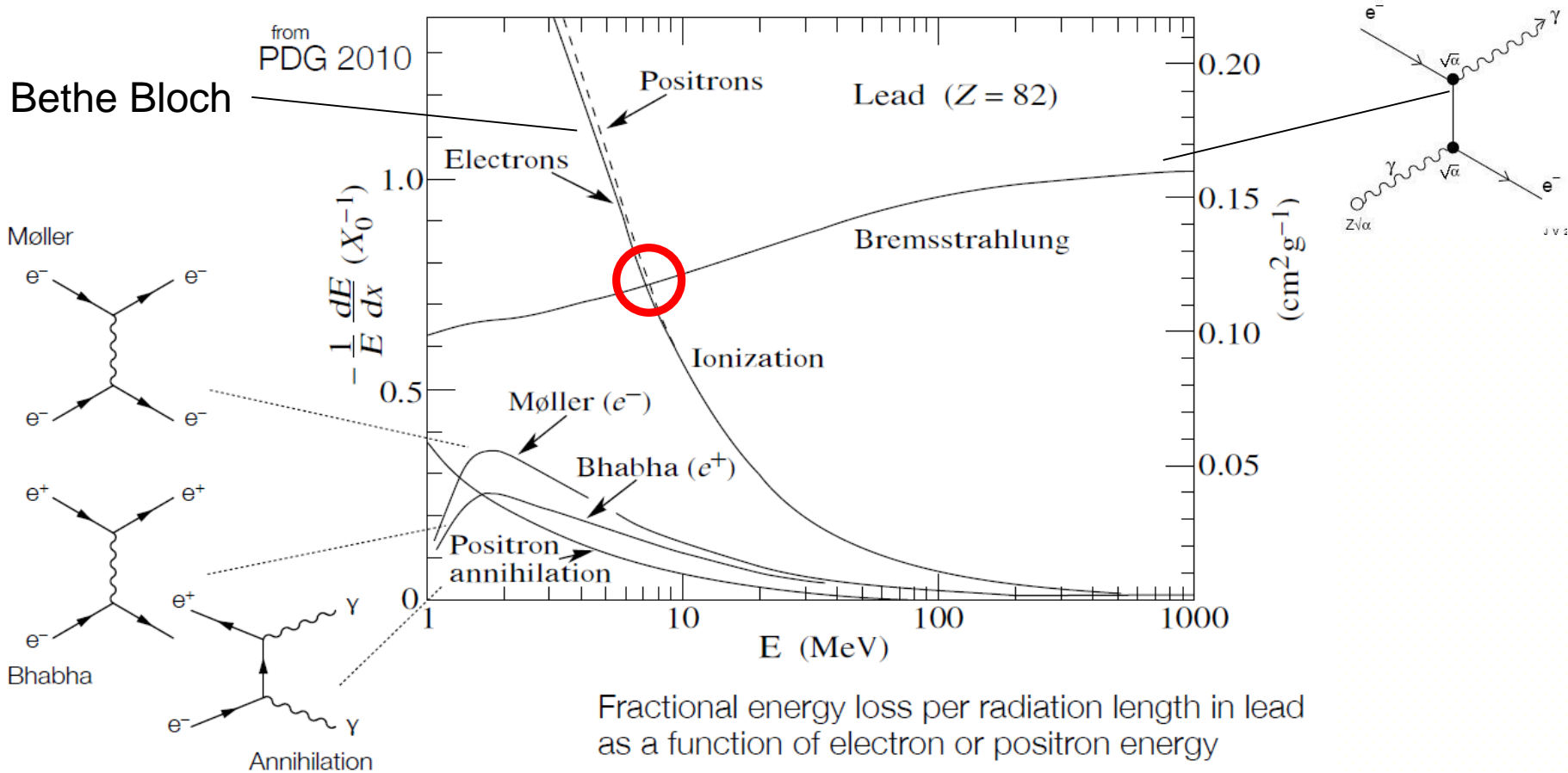
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# **Electromagnetic interaction and shower development**

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# e<sup>+</sup>/e<sup>-</sup> interaction in matter



**Critical energy  $E_c$**  : defined by  $(dE/dX)_{\text{ion}} = (dE/dx)_{\text{brem}}$

**Radiation length** : mean distance after which an electron has lost by radiation all but a fraction  $1/e$  of its initial energy  $X_0$  ( $E(\text{after } 1 X_0) = E(\text{initial})/e$ )

# Summary for e+/e-

1) Above critical energy  $E_c$  (~a few MeV) fractional energy loss dominated by bremsstrahlung, below dominated by ionization/excitation

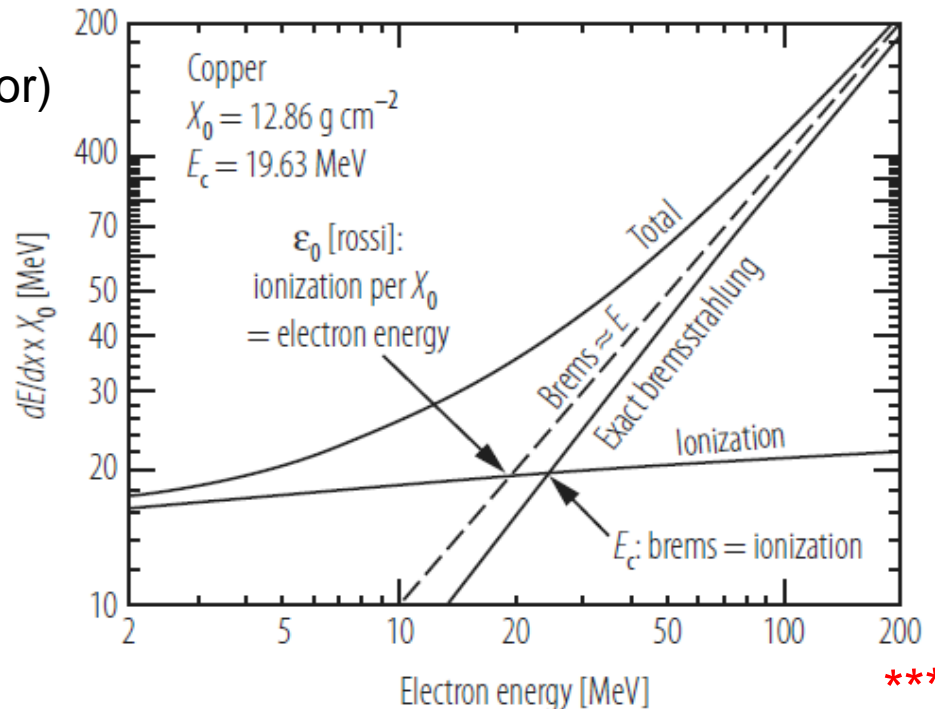
2) Energy loss by ionisation almost independent of incident energy, by radiation linear with energy

3)  $\epsilon_0 = 610 \text{ MeV} / (Z + 1.24)$

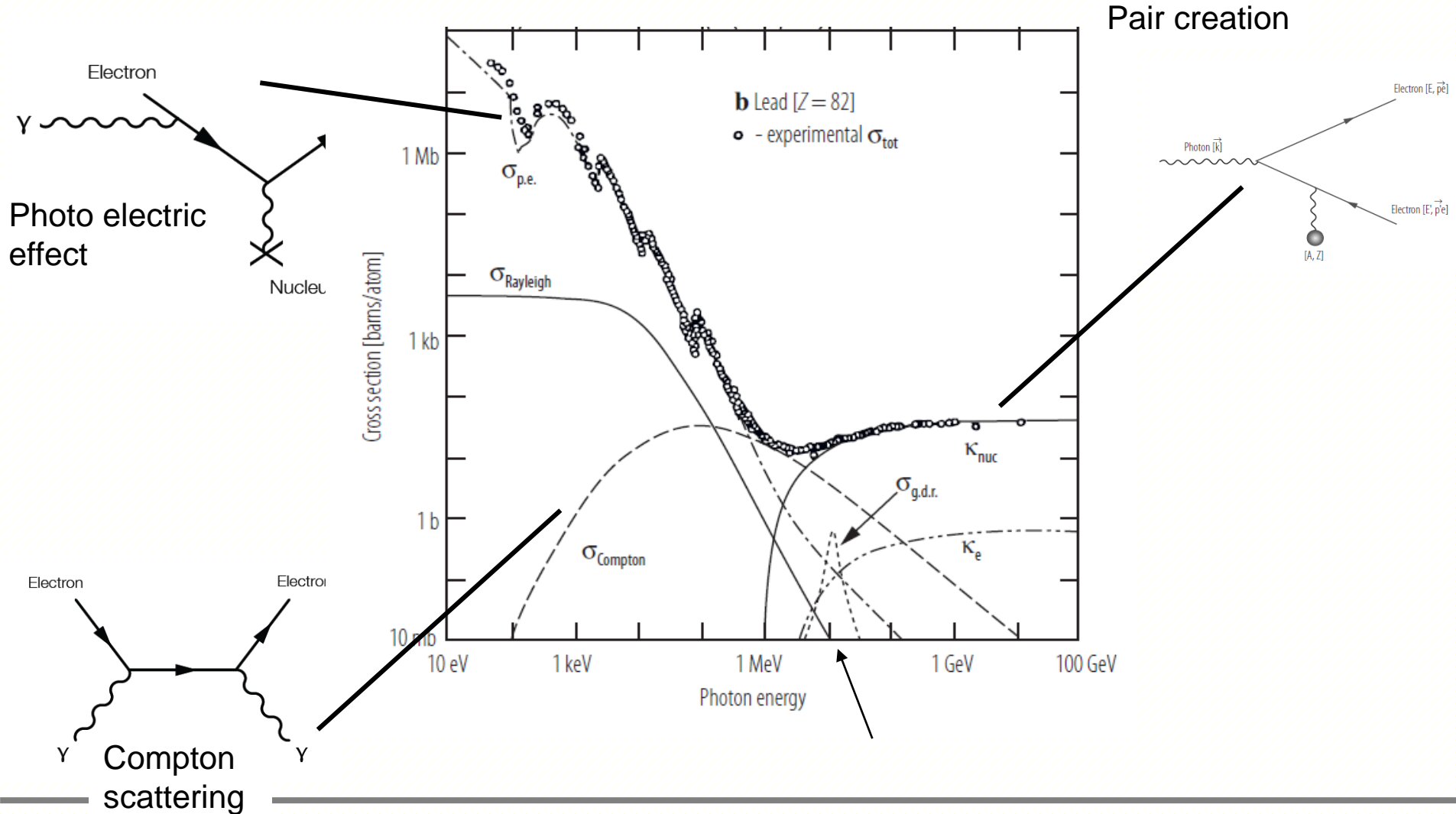
$$X_0 = \frac{716.4 \text{ g.cm}^{-2} A}{Z(Z+1) \ln(187 / \sqrt{Z})}$$

High Z material provide low critical energy and small radiation length (compact detector)

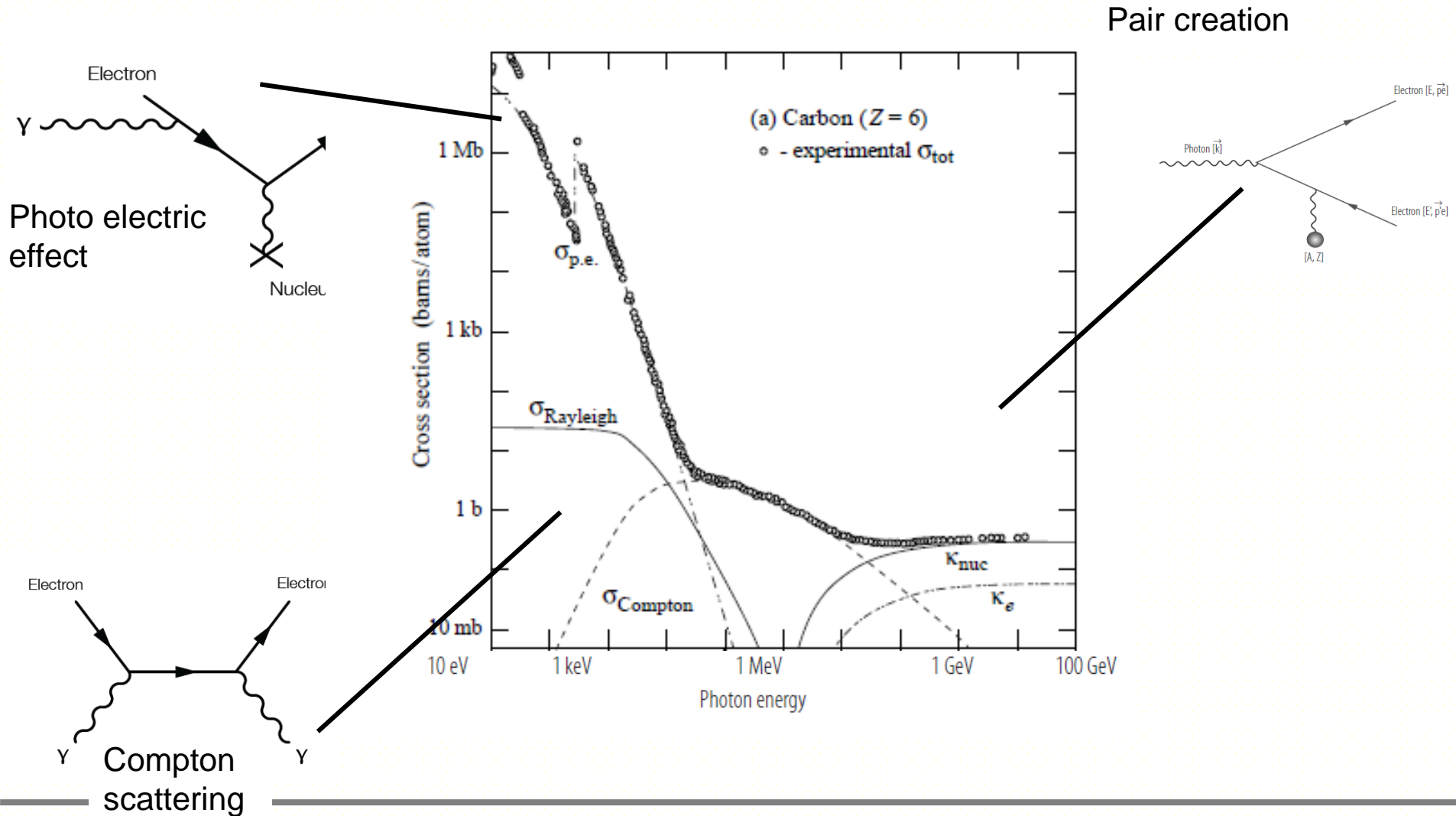
(You should divide by the density  $\rho$  to have  $X_0$  in cm)



# Photon interaction in matter



# Photon interaction in matter



# Summary for $\gamma$

- 1) Above a few  $m_e c^2$ , photon interaction is dominated by pair production.  
Cross section is constant with energy (similar fractional energy loss in brem)  
Probability that a high energy photon is **not** converted into  $e+e^-$  pair **after  $9/7 X_0$  is  $1/e$**  (37 %)  $\sigma = 7/9 A/(X_0 N_A)$ .
- 2) At intermediate energy (keV  $\rightarrow$  GeV), Compton scattering contribution  
For high Z, max of cross section  $\sim$  pair creation cross-section  
For small Z, max of cross section  $>$  pair creation cross-section
- 3) Low energy photon ( $<$  MeV) is dominated by photo-electric effect  
 **$Z^5$**  dependence of cross-section.  
In low Z material, photon can show large mean free path length .... and escape detection.

# A simplified EM shower model

EM shower model :

- After 1  $X_0$ ,  $e^{+/-} \rightarrow e^{+/-} \gamma$  and  $\gamma \rightarrow e^+ e^-$  with proba 100 %
- Equal energy split
- Cascade stops when electron/positron reaches  $\varepsilon$  (~critical energy)

1) What is the number of particles after  $n X_0$   $N(n) =$

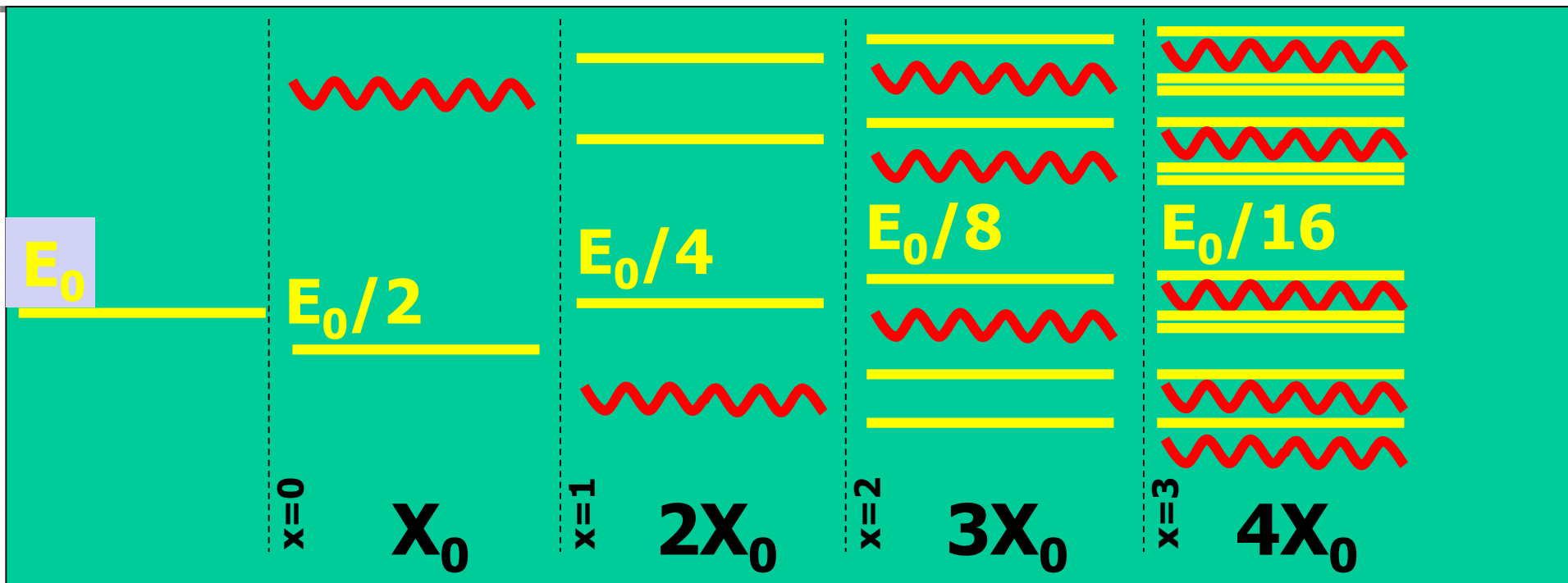
2) What is the charged particle energy after  $n X_0$   $E(n) =$

3) The cascade process stops when  $E = \varepsilon$  ( $E_c$ ) at  $n_{\max} =$   
What is the total number of particles at  $n_{\max}$   $N_{\max} =$

4) By defining  $s_0$  as the track length of electrons below the critical energy, compute the total track length  $T$  of all charged particles  $T =$   
(neglect 1 wrt  $2^{n_{\max}}$ )

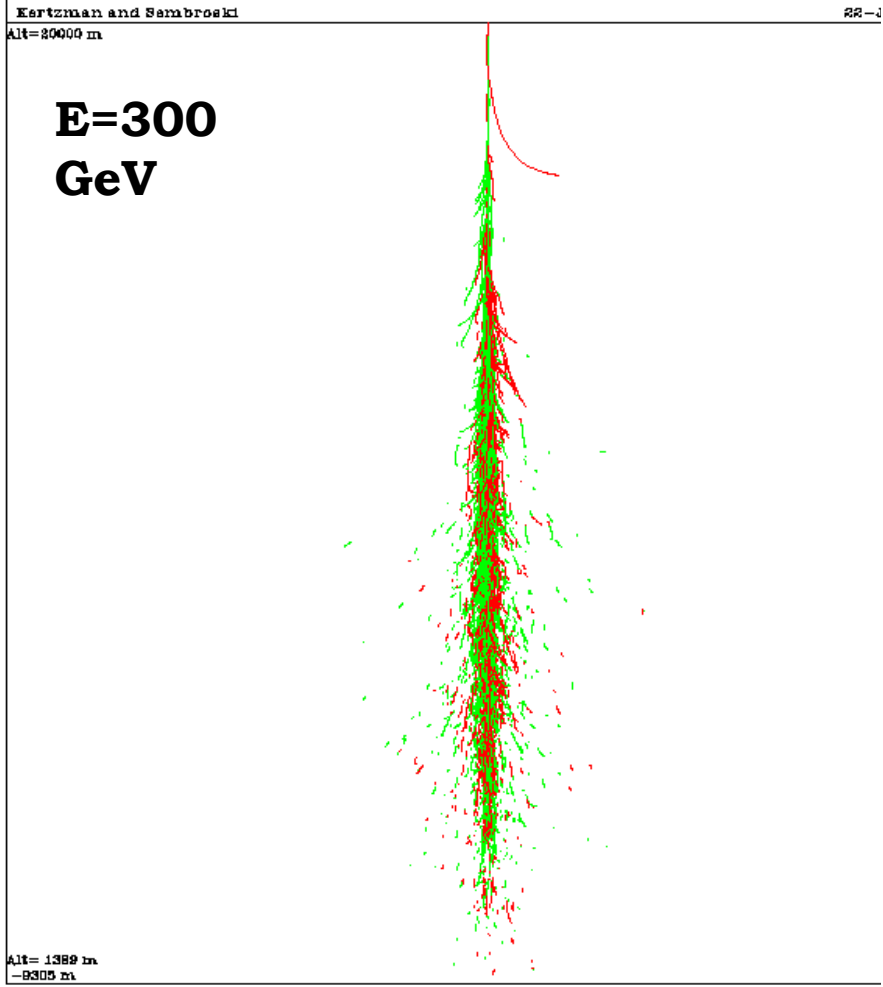
Conclude about the energy resolution if you measure  $T$  ?

# Simplified EM shower model

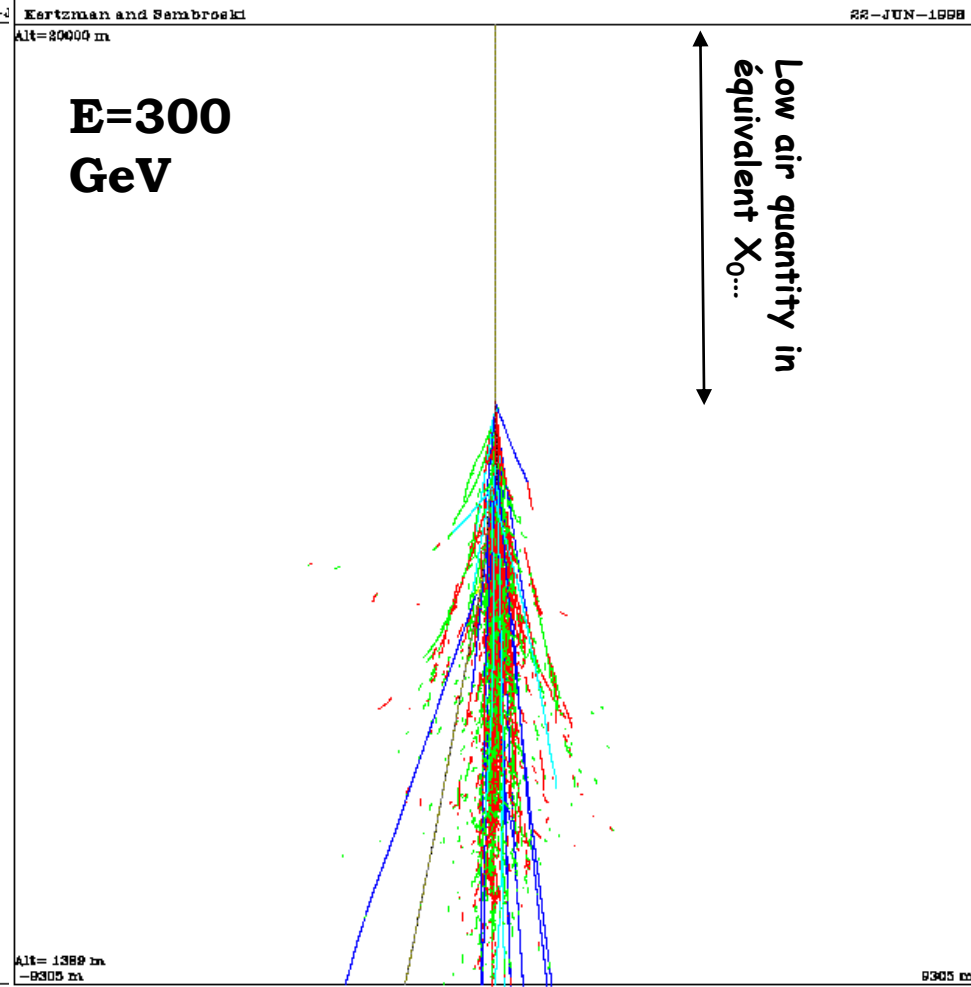


# EM shower versus Had. shower in air

## Electromagnetic shower



## Hadronic shower

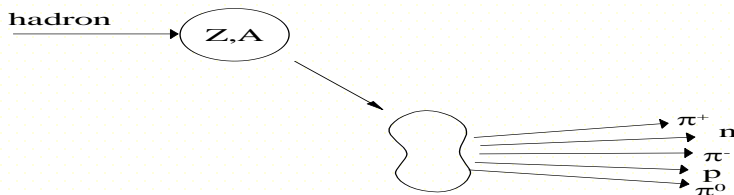


Total amount of air at sea level  $\sim 23.X_0$



# Implication : the hadronic showers

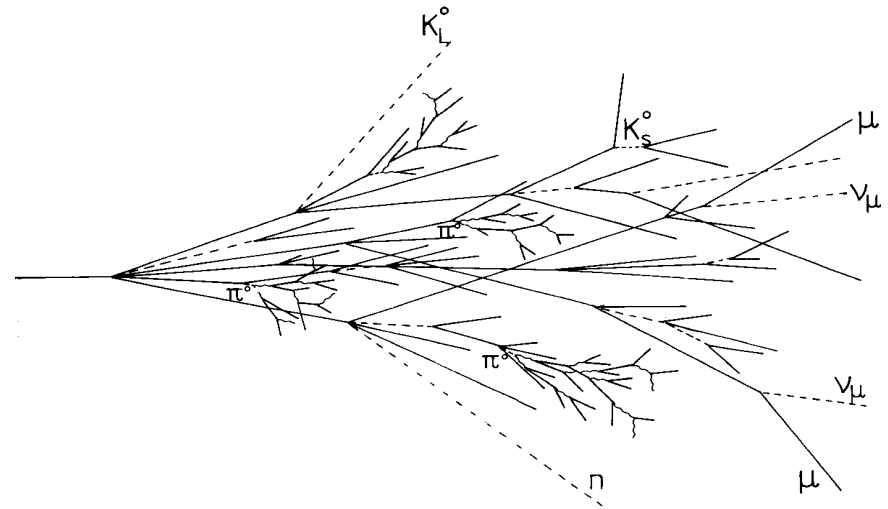
In an hadronic shower, there will be production of many  $\pi$ , K and neutrons.  $\pi^0$  will give an EM component (from 15 to 20% of initial E), some of the  $\pi$  et K at low energies will give – by decay -  $\mu$ ,  $\nu$ . Neutrons are difficult to detect (neutral, heavy part.) and will escape. This gives with neutrino the invisible energy of the shower.



Multiplicity varies with  $E \propto \ln(E)$   
 $\Rightarrow$  Quick development of the shower

$$\sigma_{inel} \approx \sigma_0 A^{0.7} \quad \sigma_0 \approx 35 \text{ mb}$$

$\sim$  independant of the energie  
 above 1 GeV for p,  $\pi$ , K...



$$n(\pi^0) \approx \ln E(\text{GeV}) - 4.6$$

example 100 GeV:  $n(\pi^0) \approx 18$

Remarq : energy profil deposition are different between EM and Had.  
 showers : higher multiplicity for hadronic interaction at the beginning of the shower development.

secondaries :  $p_t \approx 0.35 \text{ GeV}/c$

# How the hadronic shower is produced ?

Secondary particles production in hadronic showers are coming from "spallation" :

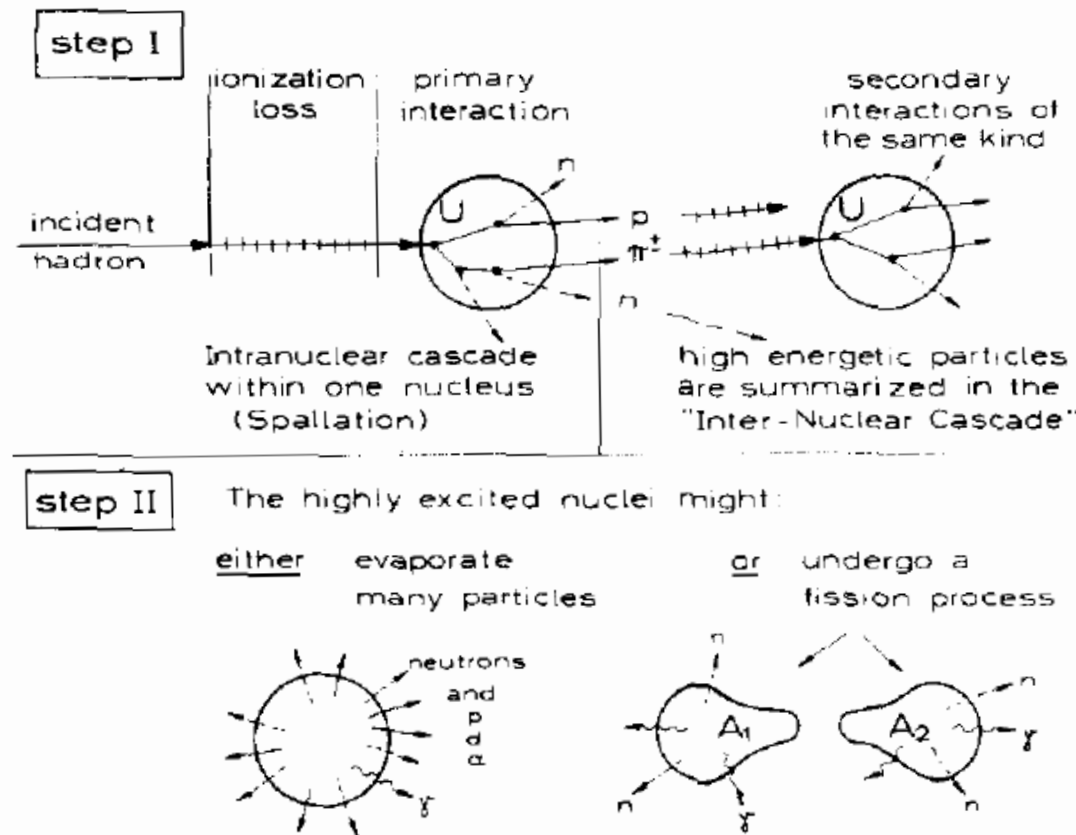
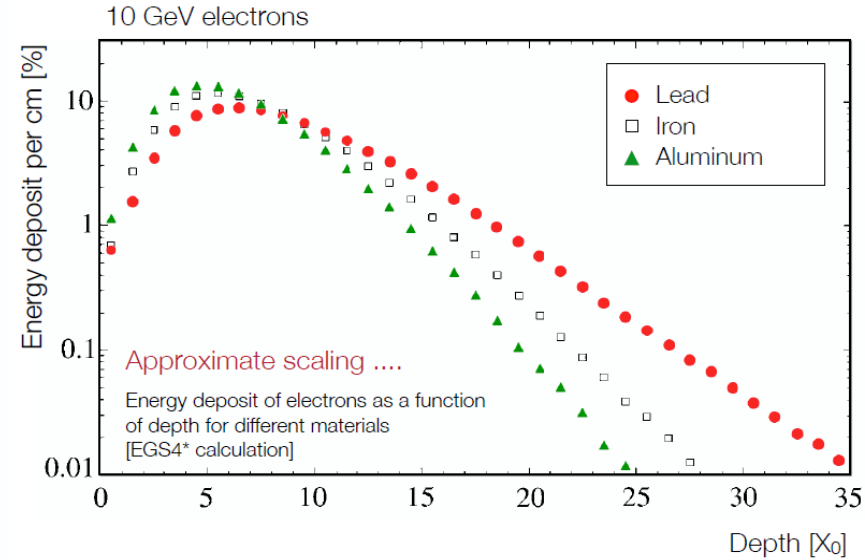
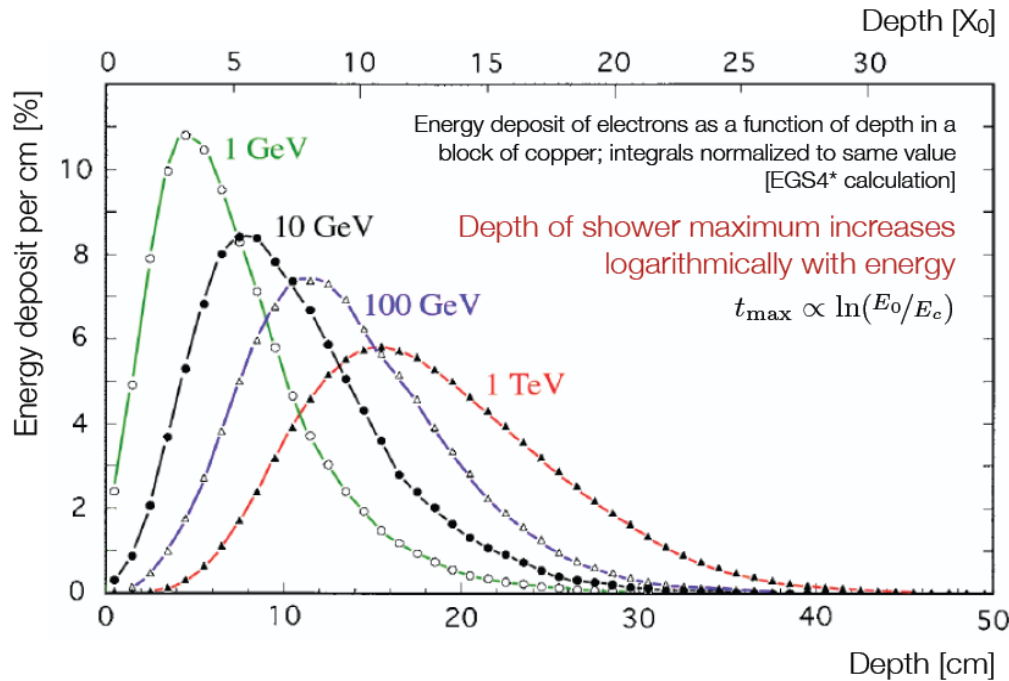


Fig. 6. Step I: Development of an "internuclear cascade". From one nucleus an intranuclear cascade releases a few high energetic spallation products, which are able to initiate further intranuclear cascade processes. Step II: The highly excited nuclei remaining from each intranuclear cascade deexcite.

# Longitudinal shower development (1)



Longitudinal profile well described by

$$\frac{dE}{dt} = E_0 b \frac{(bt)^{a-1} e^{-bt}}{\Gamma(a)},$$

$$t_{\max} = \frac{a-1}{b} = \ln\left(\frac{E_0}{E_c}\right) + C_{\gamma e}$$

-0.5 for e+/-  
+0.5 for  $\gamma$

$$\varepsilon(\text{Pb}) = 7 \text{ MeV}, \quad \varepsilon(\text{Al}) = 39 \text{ MeV}$$

Shower starts early in low Z material  
( $t_{\max}$  dependence with  $E_c$ )

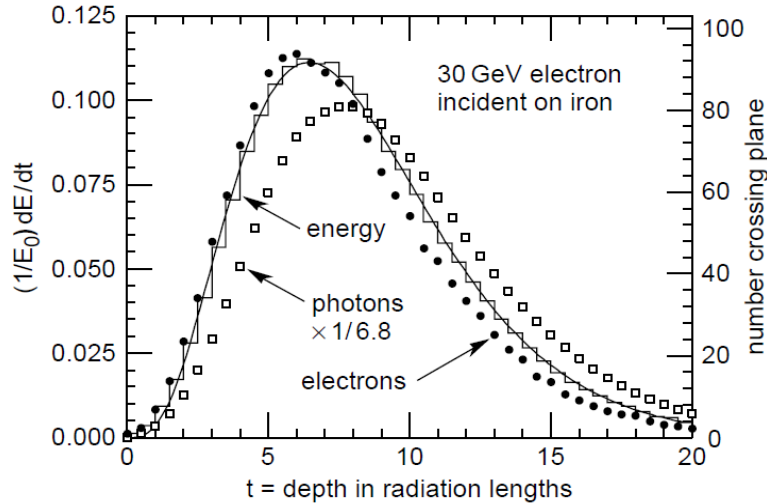
$$[dE/dx * X_0 \sim Z \times 1/Z^2 = 1/Z]$$

Not true when looking at depth in cm !

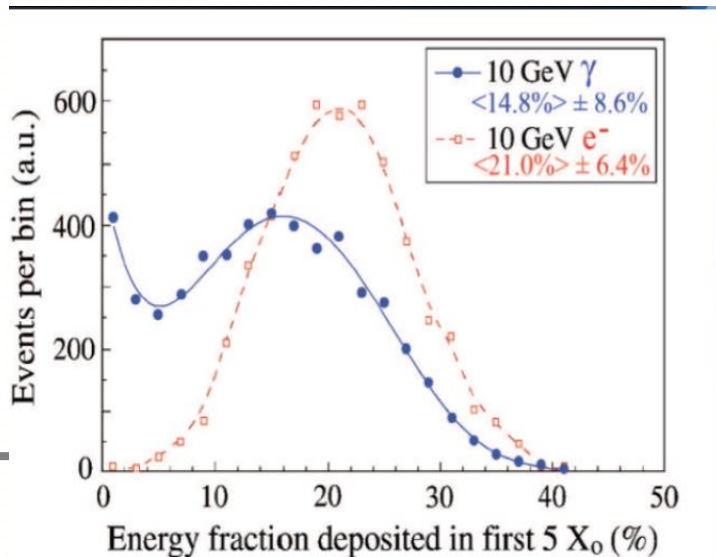
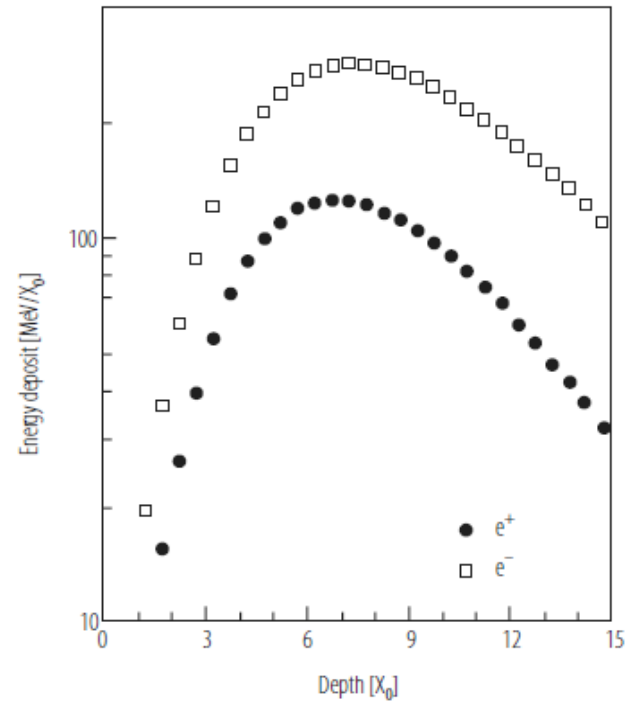
Radiation length Al : 89 mm, Pb 5.6 mm

# Longitudinal shower development (2)

Higher penetration power of photons

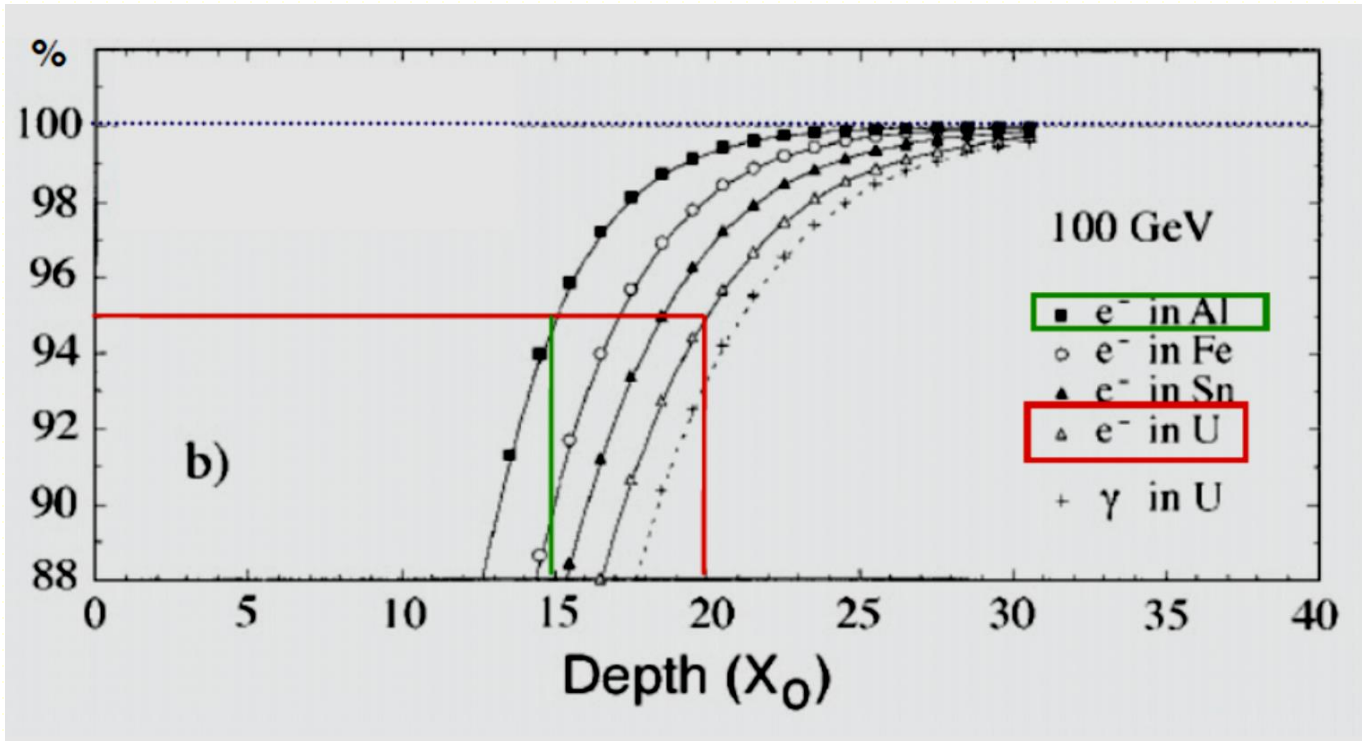


Fraction of energy deposited by  $e^-$  and  $e^+$



75 % deposited by  $e^-$ , 25 % by  $e^+$   
 Why ?

# Shower containment

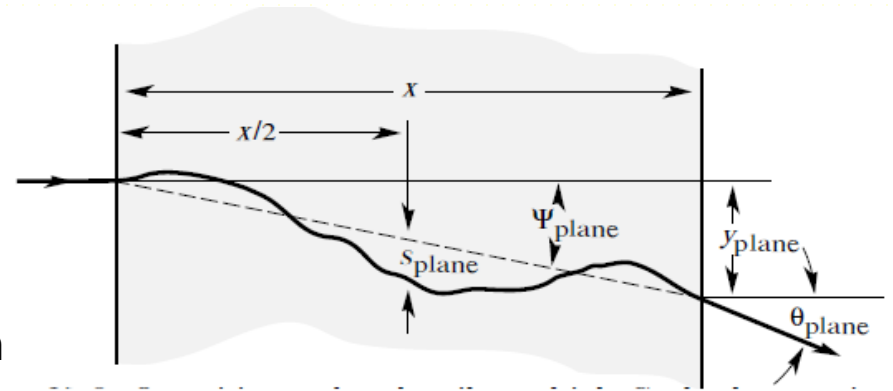


100 GeV electron contained in **15**  $X_0$  of Al and **20**  $X_0$  of U  
but remind that  $X_0(\text{Al}) = 89$  mm and  $X_0(\text{U}) = 3.2$  mm, i.e. **130** vs **6** cm !

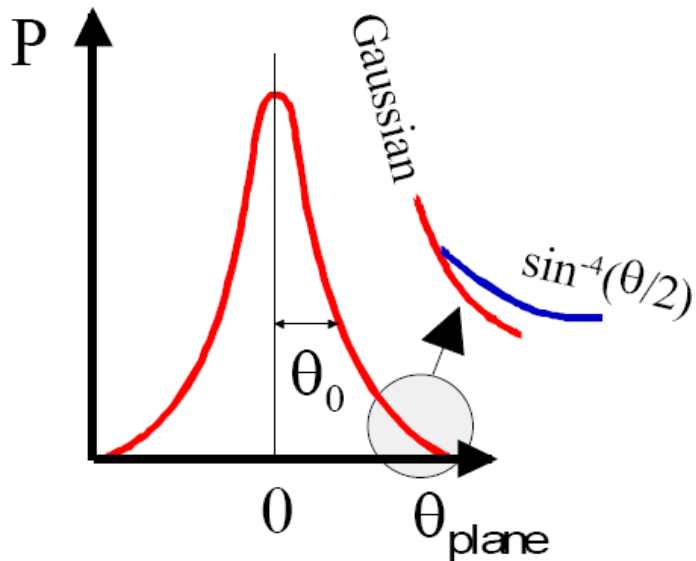
Useful parameterisation of containment  $L(95\%) = t_{\max} + 0.08Z + 9.6 [X_0]$

# Multiple scattering

Charge particle interaction with nuclei  
 → Momentum transfer (p)  
 → Particle deflexion ( $\theta$ ) ( Rutherford scattering formula  $1/\sin^4(\theta)$  )  
 → If thick material (absorber) multiple scattering. On average null effect on position but see as a fluctuation ( $\theta_0$ ) / r.m.s.



$$\theta_0 = \frac{13.6 \text{ MeV}}{\beta c p} z \sqrt{\frac{L}{X_0}} \left\{ 1 + 0.038 \ln \left( \frac{L}{X_0} \right) \right\}$$



$\theta_0$   
 -smaller for high energy (p)  
 -smaller if small material thickness (L)  
 - smaller if large radiation length

# Lateral shower development (1)

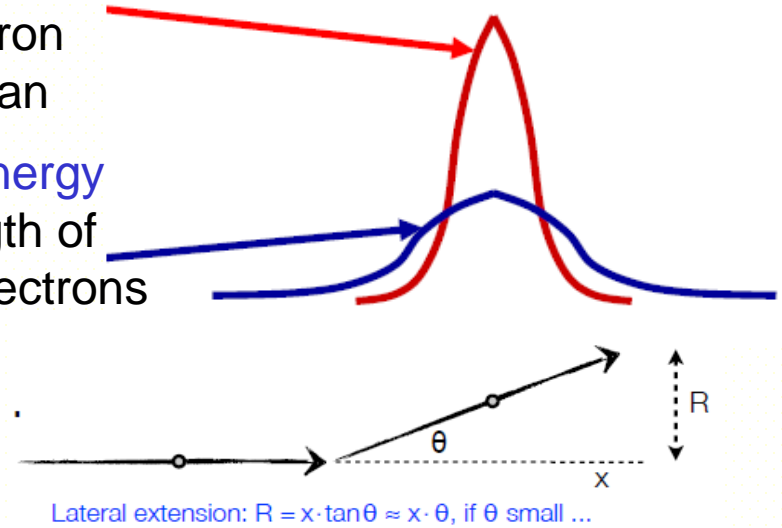
## Pair creation and multiple scattering :

At shower start, dominated by electron/positron scattering along shower axis. Mostly Gaussian

## Compton and photo-electric effect at small energy

Process are isotropic. Large penetration length of low energy photon Compton and by photo-electrons

$$\langle \theta \rangle = \frac{21.2 \text{ MeV}}{E_e} \sqrt{\frac{x}{X_0}}$$



Main contribution comes from low energy electron,  $E_c$   
If one assume that the approximate range of electrons is about  $1 X_0$ ,  
 $\rightarrow \langle \theta \rangle = 21 \text{ MeV} / E_c$  and lateral extension  $R = X_0 \langle \theta \rangle$

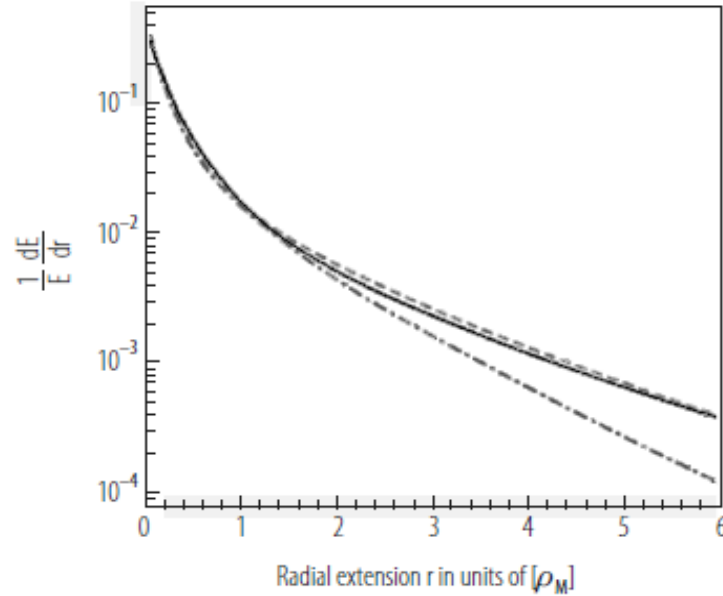
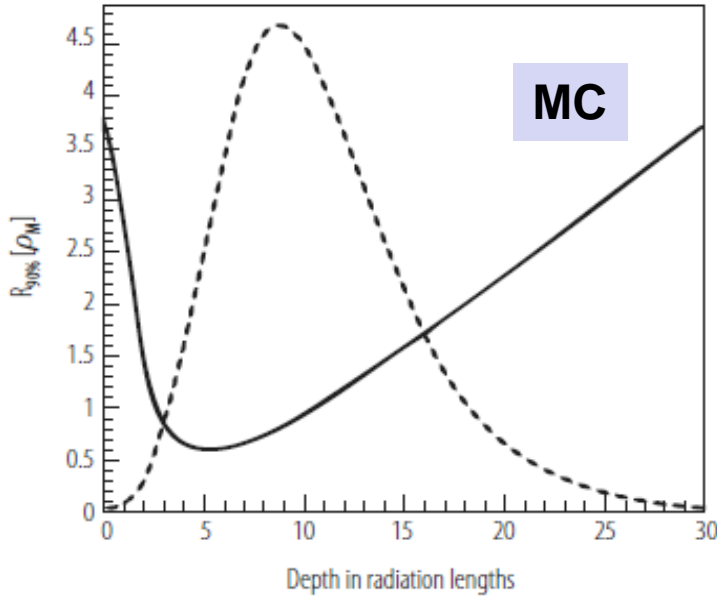
**Molière Radius** is defined as  $R_m = 21 \text{ MeV} / E_c \cdot X_0$

Convenient parameter to estimate lateral shower containment

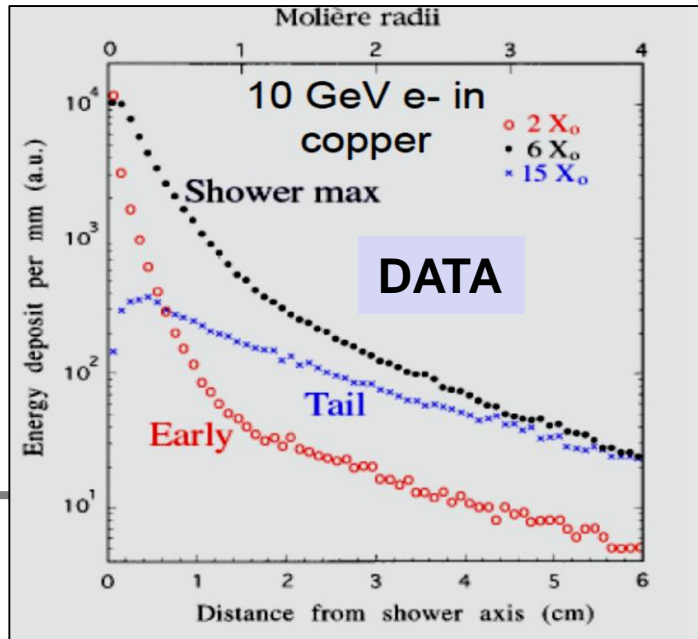
87 % (96%) of the energy of a electron shower are contained in 1 (2)  $R_m$

**Molière Radius** governed by material density

# Lateral shower development (2)



	$R_m$
Cu	15.2mm
Al	44 mm
Pb	16 mm



- Core of shower well described with  $R_m$  independently of the material (as  $X_0$  in depth)
- Same lateral shower in Cu and Pb while longer depth by 2.5
- Tails at large distance.  $Z^5$  dependence of photo electric cross section



# Material properties

Material	Z	Density [g cm <sup>-3</sup> ]	$\epsilon_c$ [MeV]	$X_0$ [mm]	$\rho_M$ [mm]	dE/dx mip [MeV cm <sup>-1</sup> ]	$\lambda_{int}$ [mm]
C	6	2.27	8.3	188	48	3.95	381
Al	13	2.70	43	89	44	4.36	390
Fe	26	7.87	22	17.6	16.9	11.4	168
Cu	29	8.96	20	14.3	15.2	12.6	151
Sn	50	7.31	12	12.1	21.6	9.24	223
W	74	19.30	8.0	3.5	9.3	22.1	96
Pb	82	11.30	7.4	5.6	16	12.7	170
U 238	92	18.95	6.8	3.2	10	20.5	105
Concrete		2.50	55	107	41	4.28	400
Glass		2.23	51	127	53	3.78	438
Marble		2.93	56	96	36	4.77	362
Si	14	2.33	41	93.6	48	3.88	455
Ar (liquid)	18	1.40	37	140	80	2.13	837
Kr (liquid)	36	2.41	18	47	55	3.23	607
Xe (liquid)	54	2.95	12	24	42	3.71	572
Polystyrene		1.032	94	424	96	2.00	795
Plexiglas		1.18	86	344	85	2.28	708
Quarz		2.32	51	117	49	3.94	428
Pb glass		4.06	15	25.1	35	5.45	330
Air (2C,1atm)		0.0012	87	304m	74m	0.0022	747m
H <sub>2</sub> O		1.00	83	361	92	1.99	849
PbWO <sub>4</sub>		8.3		8.9	20	10.2	207
CeF <sub>3</sub>		6.16		16.8	26	7.9	259
LYSO		7.40		11.4	20.7	9.6	209

Formulae for compound material :  $1/X_0 = \sum w_j / X_j$

# Summary of useful definition

Energy loss by radiation :

$$\langle E(x) \rangle = E_0 e^{-\frac{x}{X_0}} \quad \gamma \text{ absorption (e+e-)} \quad \langle I(x) \rangle = I_0 e^{-\frac{7}{9} \frac{x}{X_0}}$$

Radiation length:

$$X_0 = \frac{180A}{Z^2} \frac{\text{g}}{\text{cm}^2}$$

Critical energy:

[Attention: Definition of Rossi used]

$$E_c = \frac{550 \text{ MeV}}{Z}$$

Shower maximum:

$$t_{\max} = \ln \frac{E}{E_c} - \begin{cases} 1.0 & e^- \text{ induced shower} \\ 0.5 & \gamma \text{ induced shower} \end{cases}$$

Longitudinal energy containment:

$$L(95\%) = t_{\max} + 0.08Z + 9.6 [X_0]$$

Transverse Energy containment:

$$R(90\%) = R_M$$

$$R(95\%) = 2R_M$$

$$R_m = (21 \text{ MeV} / E_c) \cdot X_0$$

$$\sim 1/Z * Z (Z/A)$$

Small dependence with Z

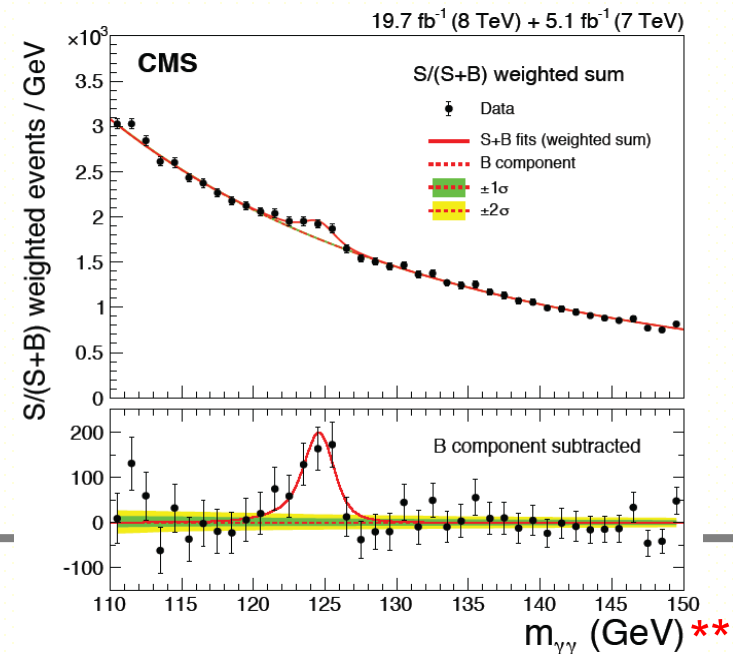
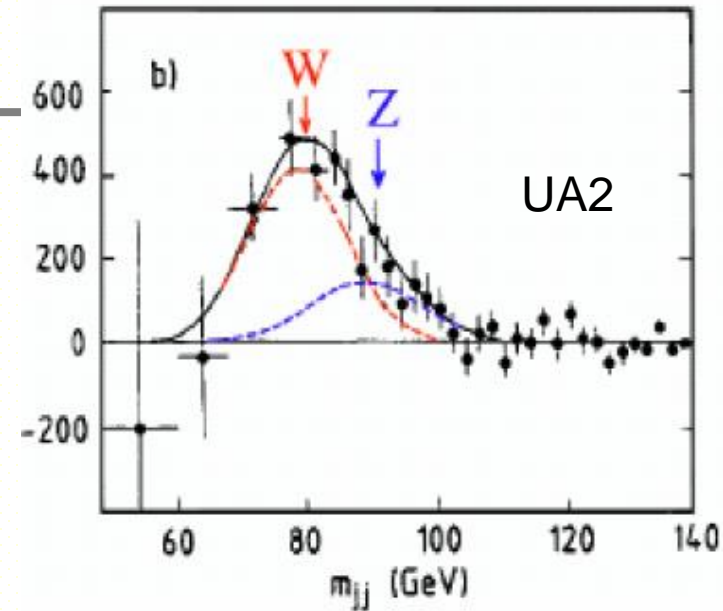
# Energy resolution

Detectable visible energy subject to fluctuation  
 → Finite energy resolution

$$\sigma(E) = \sqrt{\sigma_1^2 + \sigma_2^2 + \sigma_3^2} = \sigma_1 \oplus \sigma_2 \oplus \sigma_3$$

Most of the sources of fluctuation can be considered uncorrelated :

- Shower fluctuations
- Sampling fluctuations in sampling calorimeter
- Signal quantum fluctuations (photo detectors..)
- Leakage
- Noise in the readout
- Specific technology effects (recombination, light attenuation, gas saturation....)
- Specific to detector construction (mechanics tolerance, electronics response...)



# Energy resolution

Usually parameterized as

$$\frac{\sigma}{E} = \frac{a}{\sqrt{E}} \oplus \frac{b}{E} \oplus c$$

▶ **a** stochastic term:

- ◆ intrinsic statistical shower fluctuations
- ◆ sampling fluctuations
- ◆ signal quantum fluctuations (e.g. photo-electron statistics)

▶ **b** is the noise term:

- ◆ electronic readout noise
- ◆ pileup noise in high luminosity environment: fluctuations of energy from sources other than the primary particle (e.g. particles from other collisions in the same or in previous bunch crossings)

▶ **c** is the constant term that incorporates all other systematics:

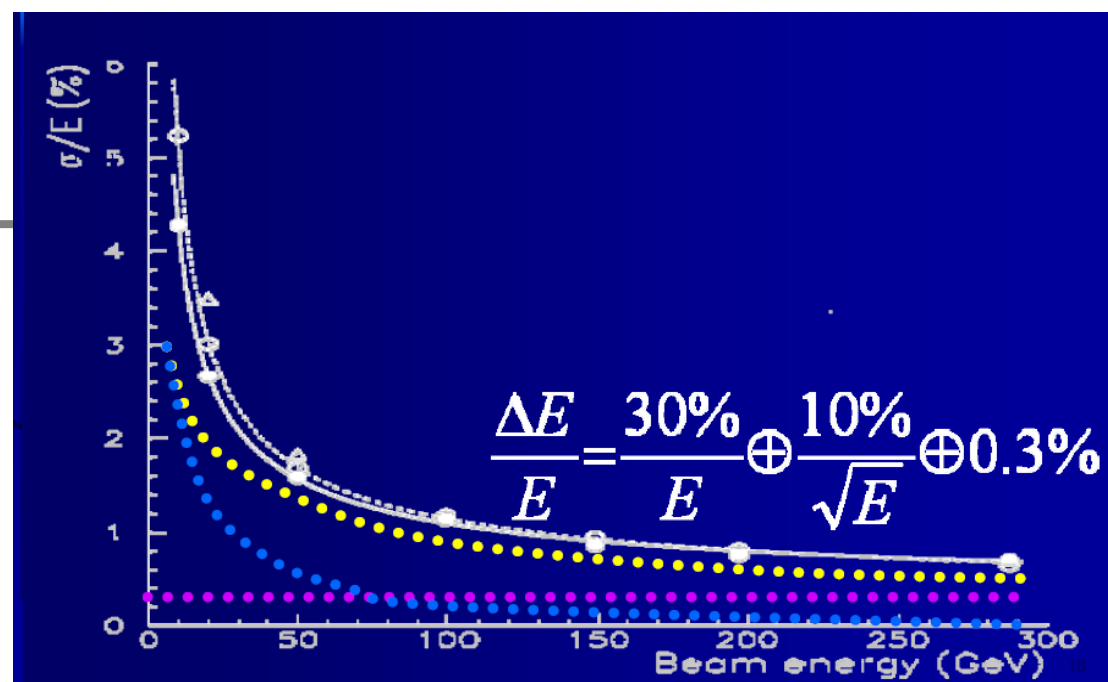
- ◆ detector response non-uniformity (hardware or calibration)
- ◆ imperfections in calorimeter construction and geometry (e.g. not fully hermetic, cracks...)
- ◆ longitudinal leakage
- ◆ energy lost in dead material upfront...
- ◆ dominant contribution at high energies

# Energy resolution

Usually parameterized as

$$\frac{\sigma}{E} = \frac{a}{\sqrt{E}} \oplus \frac{b}{E} \oplus c$$

**Quadratic sum !**



Questions: which term is affected by:

- fluctuations in the # of particles in the shower?  **$1/\sqrt{E}$  (a)**
- global scale (gain) shift? **1 (c)**
- electronics noise?  **$1/E$  (b)**
- global offset (pedestal) shift?  **$1/E$  (b)**
- shower particles escaping the calorimeter?  **$1^*$  also  $1/\sqrt{E}$  if upstream**
- fluctuations in the # of photo-electrons detected ?  **$1/\sqrt{E}$  (a)**
- pile-up (remnants of earlier events)?  **$1/E$  (b)**
- radioactivity ?  **$1/E$  (b)**
- presence of dead material? **1 (c)**
- statistical uncertainty of scale (gain) constants? **1 (c)**
- statistical uncertainty on offset (pedestal) constants?  **$1/E$  (b)**

# **Electromagnetic calorimeter technologies :**

**Homogeneous calorimeters**

**Sampling calorimeters**

# Homogeneous calorimeters

- Combine both the role of absorber and signal generation.
- Total volume sensitive to the deposited energy. (Large fraction of visible energy)



## Advantage

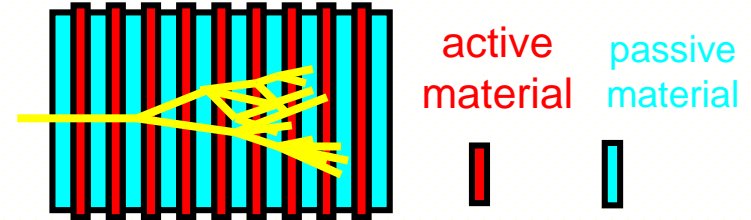
- Best energy resolution ( $1/\sqrt{E}$  term) limited by physical factors as number of photo-electrons if ideal calo (no leakage)
- Intrinsically linear in principle
- Well suited for low energy application (nuclear spectroscopy, medical application...)

## Disadvantage

- Limited segmentation (especially in depth)
- Can not be used for hadronic shower to keep "reasonable" detector size.
- Cost (Pb or Cu less expensive than crystals, Silicon or noble liquid)

# Sampling calorimeters

- Shower is sampled in active layers interleaved with absorbers

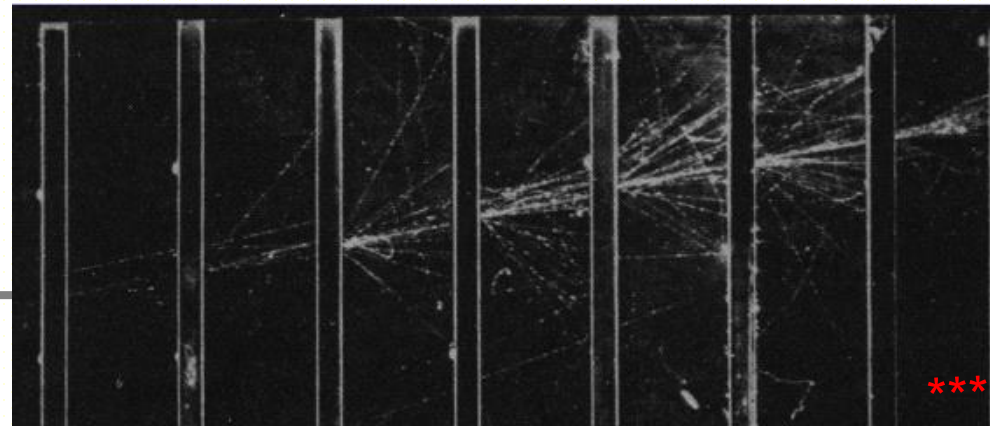


## Advantage

- Can achieve easily lateral and longitudinal segmentation  
→ Angular measurement and particle Identification
- cheaper calorimeter (in principle !) as absorber not too expensive
- Only possibility for Hadron calorimeters

## Disadvantage

- Small fraction of energy seen  
→ Stochastic term degraded



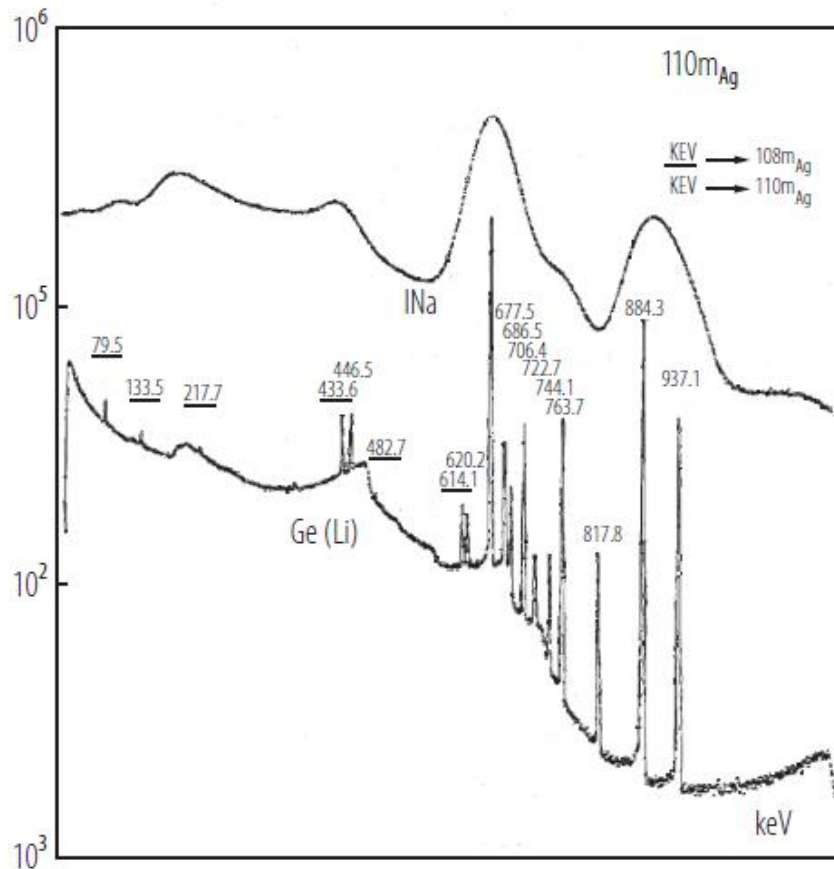


# Homogeneous calorimeter technology

Should be dense enough to contain EM shower, give enough signal

- **Semiconductor Si, Ge** : very low threshold to create electron-hole pair (2.9 eV in Ge )  
→ *Use in nuclear spectroscopy, medical application*
- **Cerenkov** : high refractive index induces cerenkov light with relativistic charged particle  
→ *Lead glass, OPAL @LEP*
- **Scintillators** : ionisation tracks converted in light in crystals (fluorescence)  
→ *NaI(Tl) (Crystal Ball), L3 (BGO), Babar, Belle, KTeV (CsI), PbWO<sub>4</sub> (CMS)*
- **Noble Liquids** : cryogenics detectors. Ionisation produces charge and light (scintillation)  
→ *Kr (NA48, KEDR)*

# Si/Ge low energy homogeneous calorimeters



Ge : energy to create an electron-hole pair  
at 77 K : 2.9 eV

1 MeV  $\rightarrow$   $N = 3.4 \cdot 10^5$  pairs

$\sigma_E/E \sim 1/\sqrt{N} = 0.17\%$

Even better due to Fano factor (pairs created not statistically independent, constrained by total energy of incident particle, similar to binomial variance )  $F = 0.13$  in Ge

$\sigma_E/E \sim \sqrt{(F/N)} = 0.06 \%$

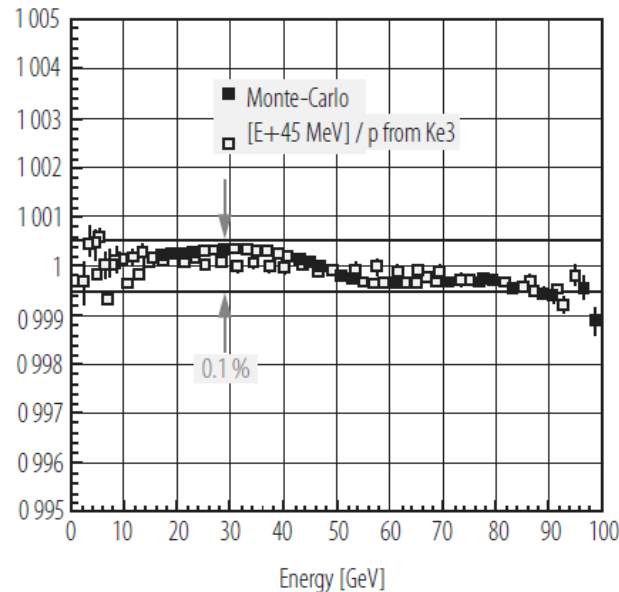
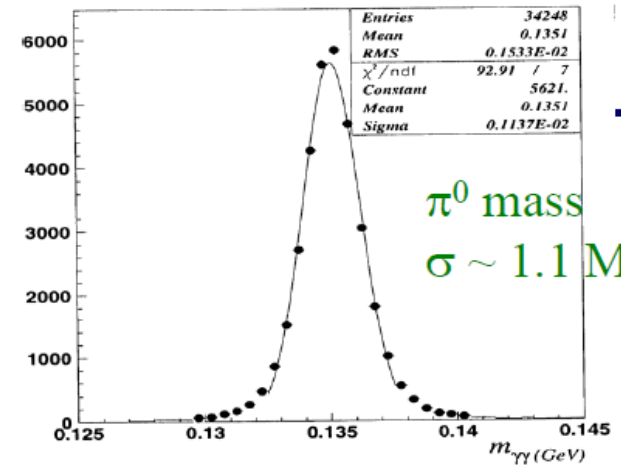
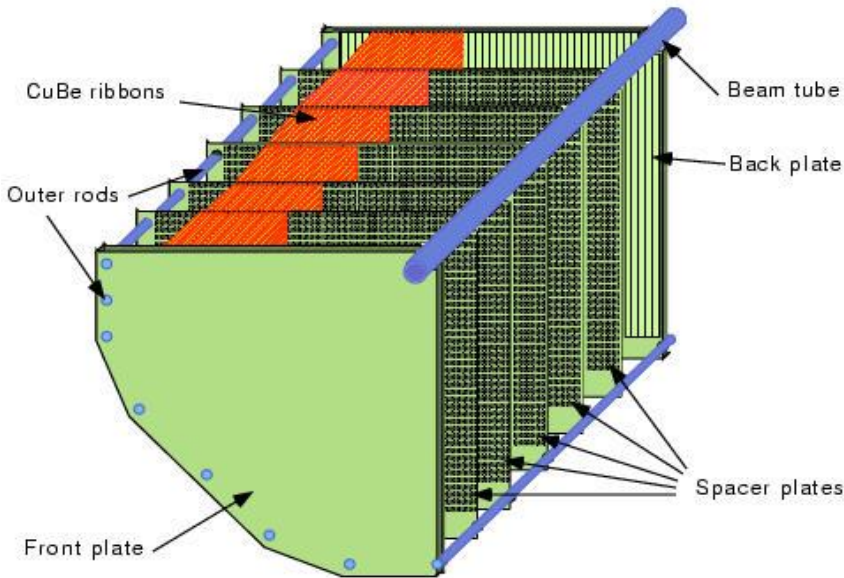
Popular detectors in Nuclear physics (AGATA for instance)

# Noble liquid homogeneous calorimeters

NA48 : Fixed target experiment for CP violation, and now NA62 rare Kaon decays  
 1 MeV resolution needed on  $\pi^0$  mass

Liquid Krypton @ 120 K

LKr CALORIMETER ELECTRODE STRUCTURE



Linearity as important as resolution

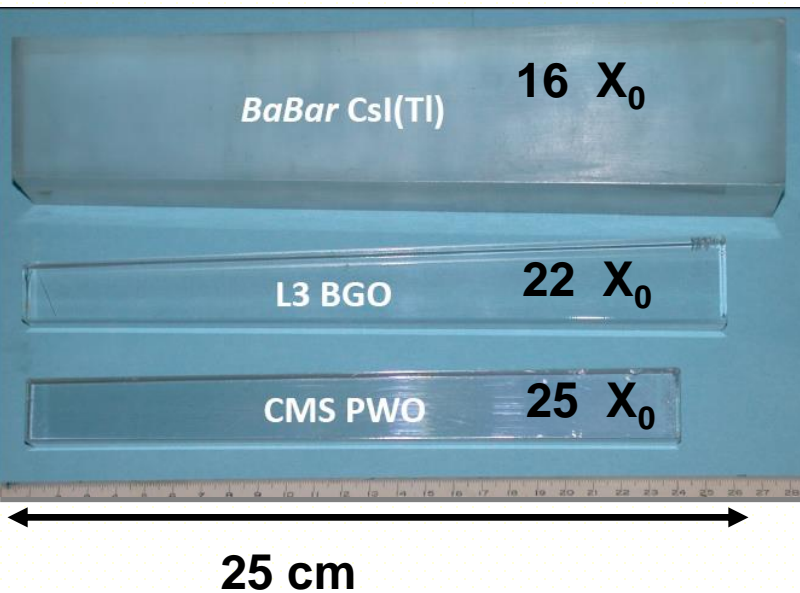
Need a careful control of upstream dead material

# Homogeneous crystals calorimeters

Crystal	light	$\rho$ [g/cm <sup>3</sup> ]	$X_0$ [cm]	$\tau$ [ns]	$\lambda$ [nm]	Output	Damage (Gy)
NaI	Scint	3.67	2.59	250	410	1 (40000 ph/MeV)	10
BGO	Scint	7.13	1.12	300	410	0.15	10 *
BaF <sub>2</sub>	Scint	4.89	2.05	600	310	0.20	10 <sup>5</sup> *
CsI (TI)	Scint	4.53	1.85	35 (1000)	420	0.05 (0.45)	10 <sup>3</sup>
PbWO <sub>4</sub>	Scint	8.28	0.89	5-15	430	0.01	10 <sup>4</sup> *
CeF <sub>3</sub>	Scint	6.16	1.68	10-30	325	0.10	
Pbglas5	Cer	4.08	2.54	fast	< 350	0.00015	
Pbglas6	Cer	5.20	1.69	fast	< 350	0.00023	

\* Hygroscopy ~2%/°C \*

# Homogeneous crystals calorimeters



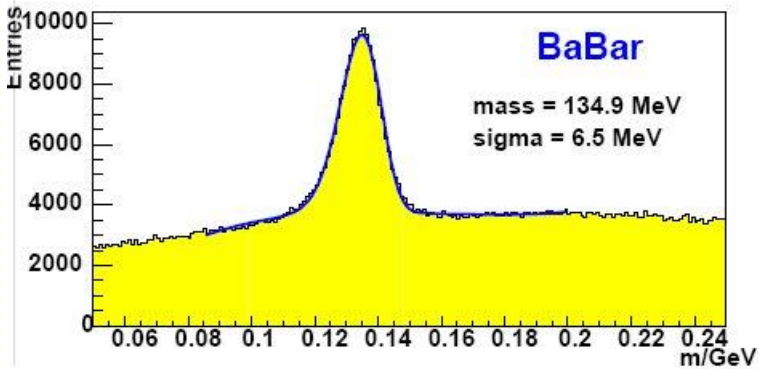
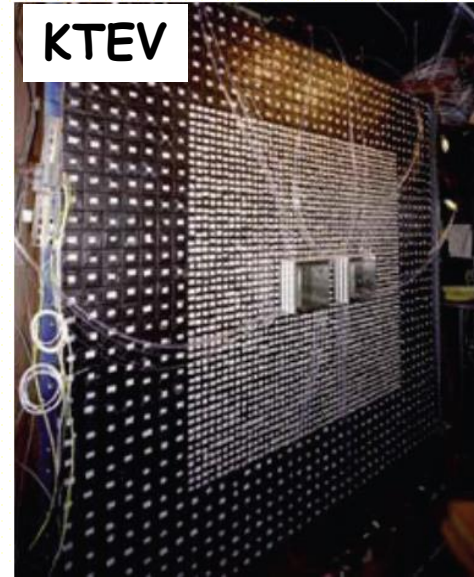
Growing crystals not always easy task  
By construction non uniform response from one crystal to another (up to 10-20%), different transparency

# Homogeneous crystals calorimeters

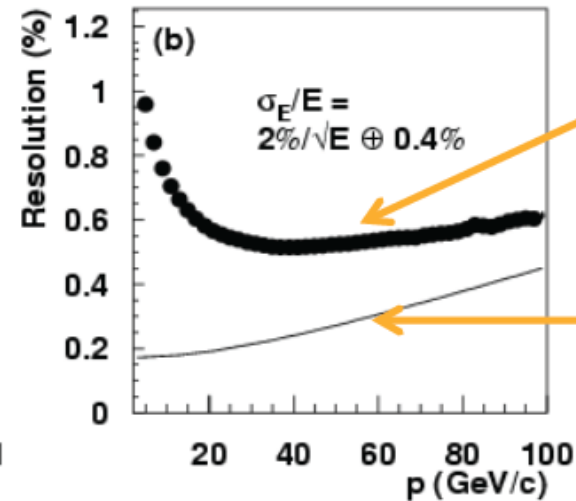
Babar



KTEV



$$\frac{\sigma_E}{E} = \sqrt{\left(\frac{0.066\%}{E_n}\right)^2 + \left(\frac{0.81\%}{\sqrt[3]{E_n}}\right)^2 + (1.34\%)^2}, \quad E_n = E/\text{GeV},$$



E/P resolution

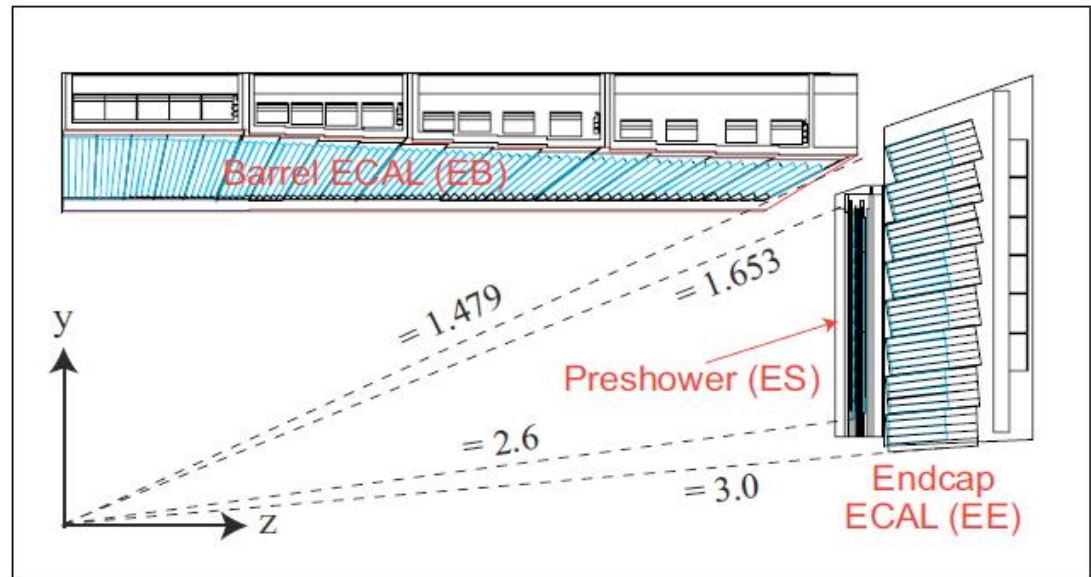
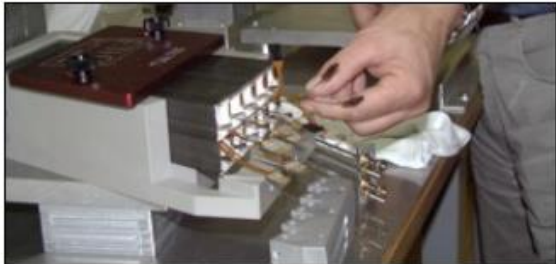
Estimated momentum resolution

# CMS calorimeter

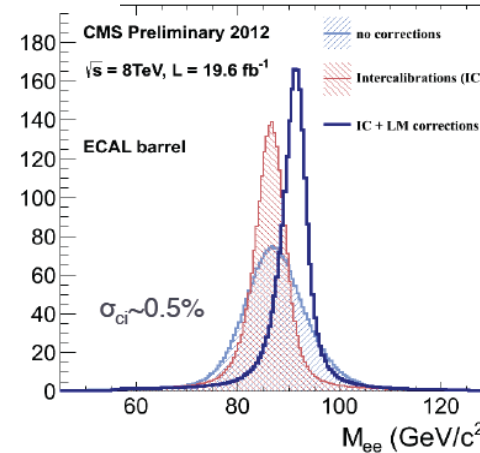
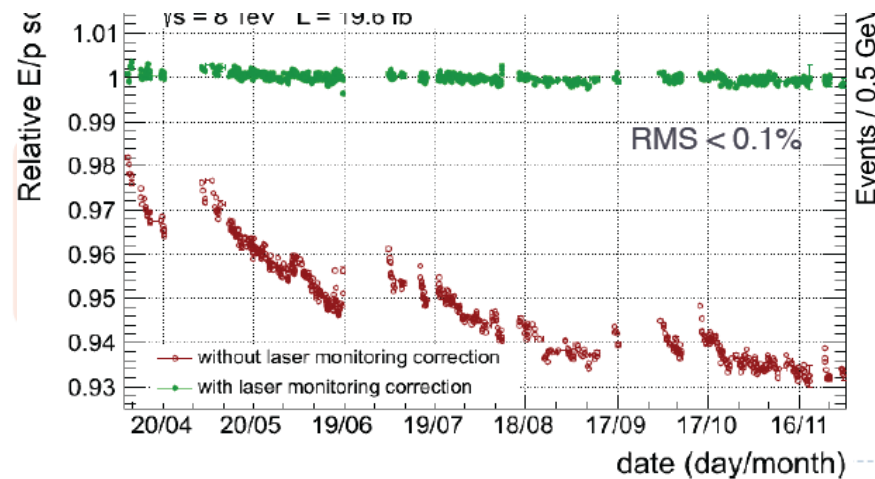
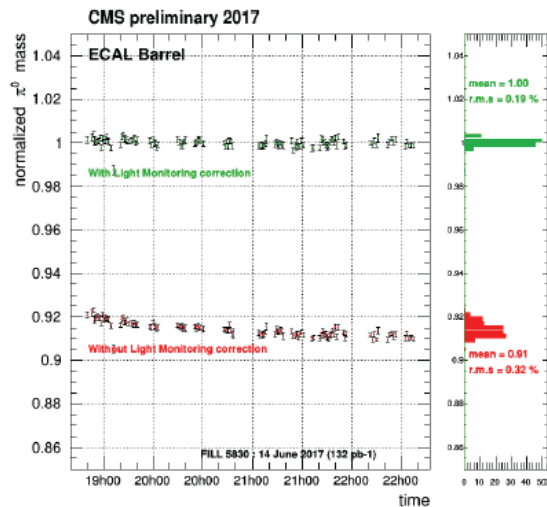
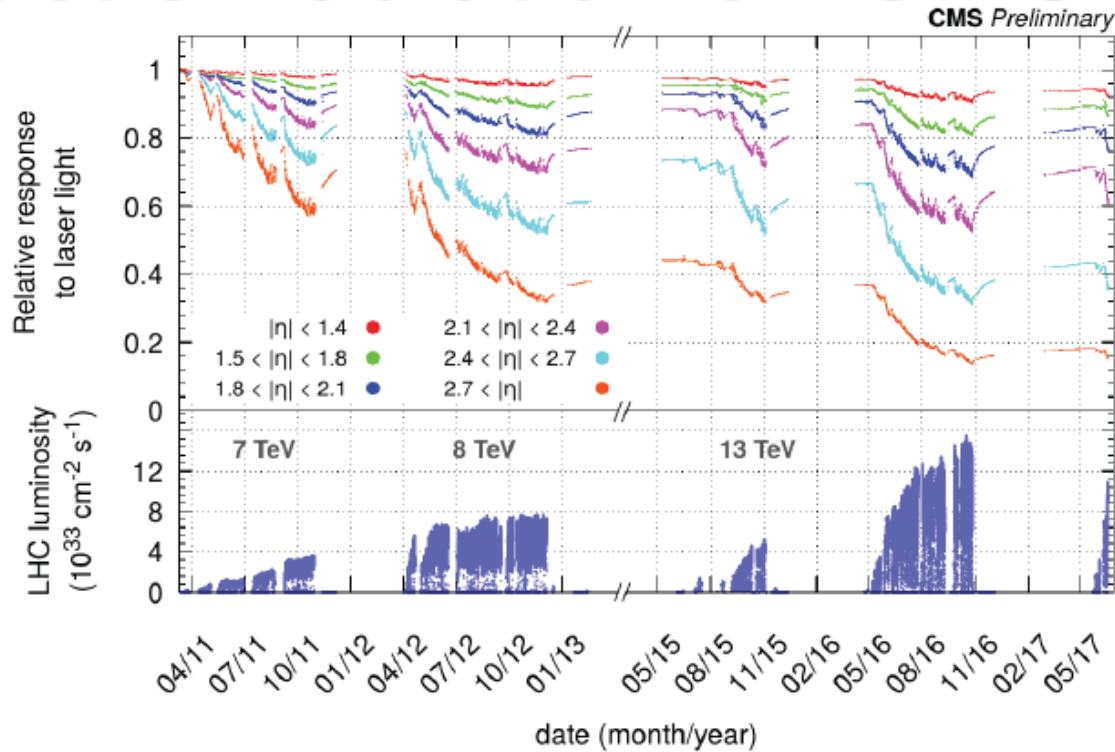


Scintillator : PBWO<sub>4</sub> [Lead Tungsten]  
Photosensor : APDs [Avalanche Photodiodes]

Number of crystals: ~ 70000  
Light output: 4.5 photons/MeV



# Impact of radiation on CMS calo





# Performance of homogeneous calorimeters

Technology (Experiment)	Depth	Energy resolution	Date
NaI(Tl) (Crystal Ball)	$20X_0$	$2.7\%/E^{1/4}$	1983
Bi <sub>4</sub> Ge <sub>3</sub> O <sub>12</sub> (BGO) (L3)	$22X_0$	$2\%/\sqrt{E} \oplus 0.7\%$	1993
CsI (KTeV)	$27X_0$	$2\%/\sqrt{E} \oplus 0.45\%$	1996
CsI(Tl) (BaBar)	$16-18X_0$	$2.3\%/E^{1/4} \oplus 1.4\%$	1999
CsI(Tl) (BELLE)	$16X_0$	1.7% for $E_\gamma > 3.5$ GeV	1998
PbWO <sub>4</sub> (PWO) (CMS)	$25X_0$	$3\%/\sqrt{E} \oplus 0.5\% \oplus 0.2/E$	1997
Lead glass (OPAL)	$20.5X_0$	$5\%/\sqrt{E}$	1990
Liquid Kr (NA48)	$27X_0$	$3.2\%/\sqrt{E} \oplus 0.42\% \oplus 0.09/E$	1998

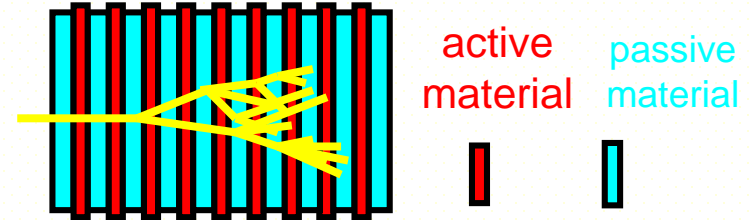
Homogeneous

For crystals/Cerenkov stochastic term contains both shower fluctuation and photo-electron statistics (converting photons in electrical signal).

Example : Lead Glass : Cerenkov only if  $e^\pm$  with  $E > 0.7$  MeV and photon-detector provides 1000 photo-electrons/GeV  
 Expected resolution for 1 GeV ?  $\rightarrow$  stochastic term ?

# Sampling calorimeters

- Shower is sampled in active layers interleaved with absorbers

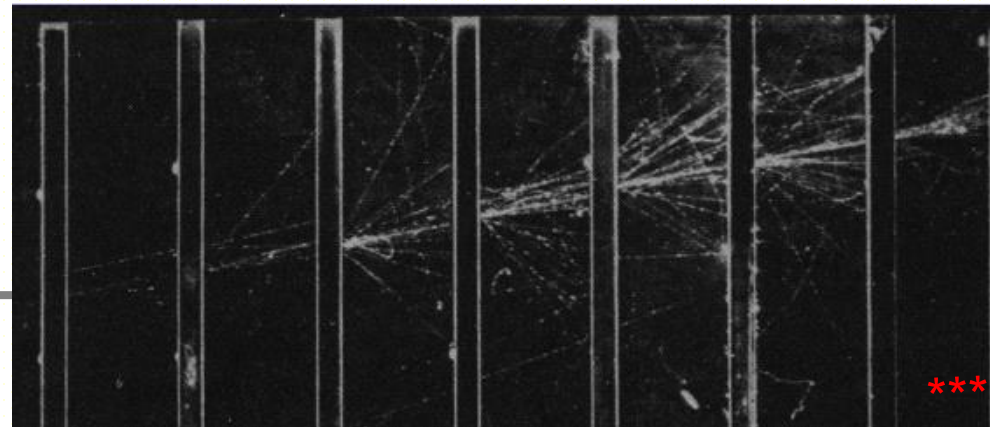


## Advantage

- Can achieve easily lateral and longitudinal segmentation  
→ Angular measurement and particle Identification
- cheaper calorimeter (in principle !) as absorber not too expensive
- Only possibility for Hadron calorimeters

## Disadvantage

- Small fraction of energy seen  
→ Stochastic term degraded



# EM sampling calorimeter technology

Absorber with dense material with low critical energy (high Z) for shower development (U, Pb, W...). All technologies possible for active layers :

- **Scintillators**  
→ U + scint (Zeus@Hera) , Pb + scint (CDF@Tevatron)
- **Gaseous detectors**  
→ Pb + wire chambers (ALEPH@LEP)
- **Liquid Argon :**  
→ LAr + Pb (Cello , NA31, SLD, H1@Hera, ATLAS@LHC)  
→ LAr + U (D0@Tevatron)  
*Kr considered as option at SSC & LHC*
- **Semiconductors**  
→ Si+W (Pamela, Calice@ILC, CMS HGCal@HL-LHC)

# Sampling calorimeter

Simplified model of previous :  
Active medium : counts only charged particle produced in absorber shower development.  
 $N_{max} = E/E_c$  , 2/3 are charged particles

Pb :  $E_c = 7.4$  MeV  
For 1 GeV shower,  $N_{ch} \sim 90$   
 $\sigma(N_{ch})/N_{ch} = 1/\sqrt{N_{ch}} = 10\%$   
Typical best stochastic term of sampling EM calorimeter

Key parameters :

**Sampling frequency** : Number of times a high energy electron/ $\gamma$  is sampled. Linked to absorber thickness (t).

Thinner is t, higher is the sampling frequency, better is the resolution, but if too small correlated signals in two active layers

**Sampling fraction** : Fraction of energy deposited by a mip in active layer

$$f_{smp} = \frac{E_{mip}(active)}{E_{mip}(active) + E_{mip}(absorbeur)}$$

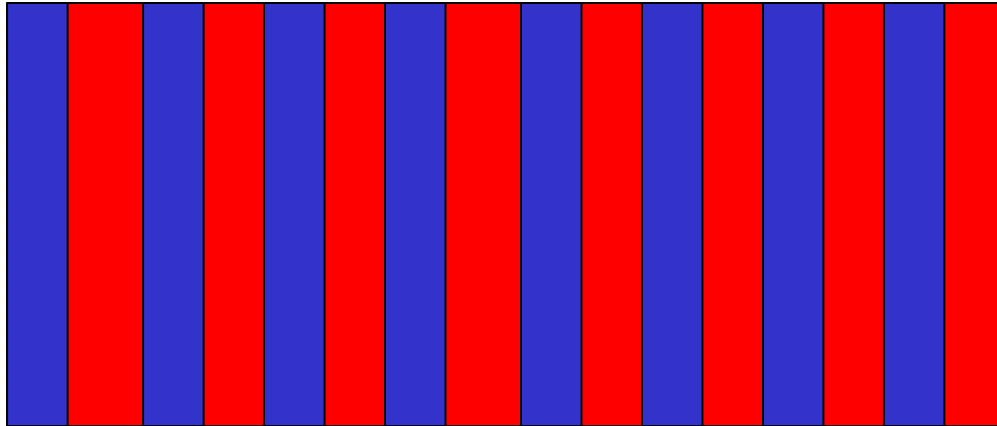
$E_{mip} = (dE/dx) * distance$   
t for passive material, s for active

**Fractional energy response**  $f_R = (E_{active}) / (E_{active} + E_{passive})$  (includes showering process)

# Sampling Calorimeter

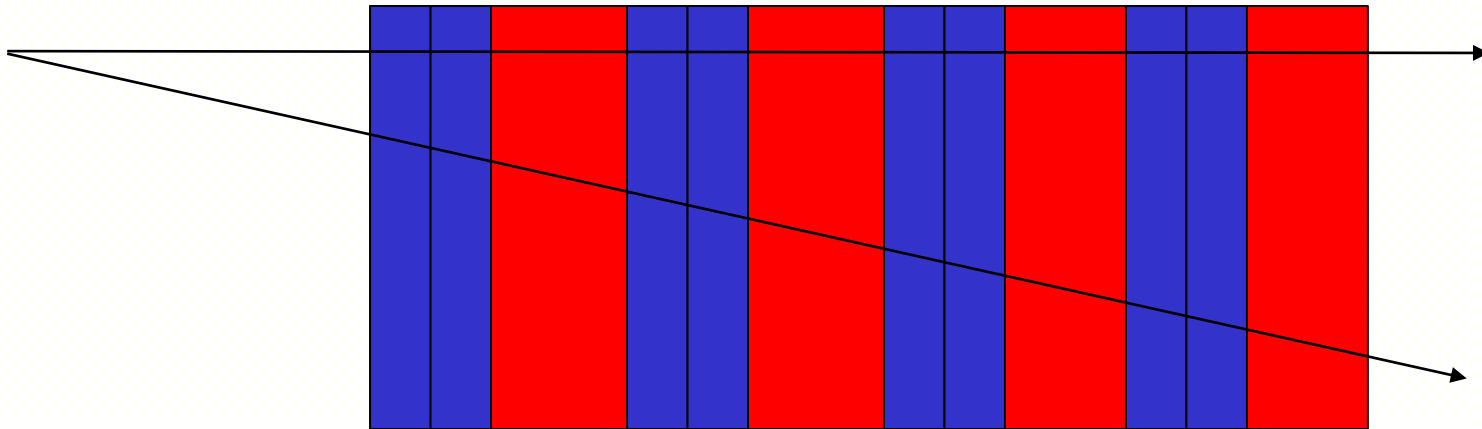
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Blue absorber (t) , red active medium (s)



# Sampling Calorimeter

Blue absorber (t) , red active medium (s)



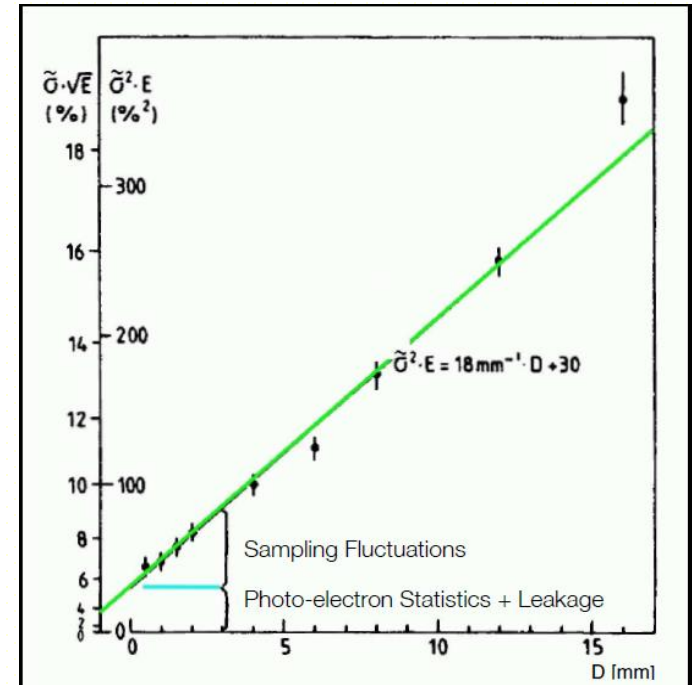
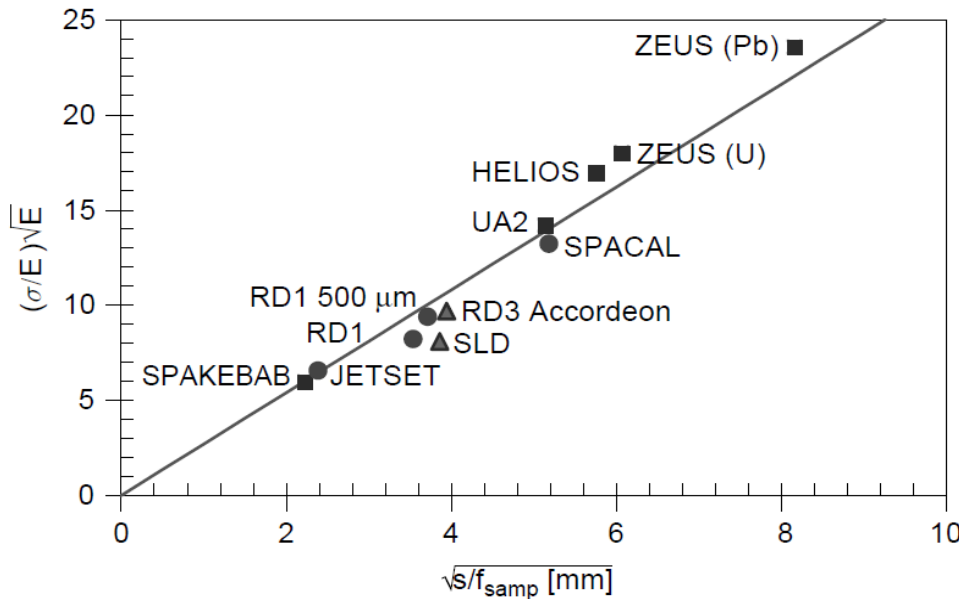
Same sampling fraction but smaller sampling frequency  
(4 / 8)  $\rightarrow$  worse stochastic term

Angular effect : constant sampling fraction but smaller frequency

# EM sampling calorimeter E resolution

Total track length detectable  $\rightarrow T_d = f_{\text{samp}} \cdot T$

$\rightarrow$  expect energy resolution as  $1/\sqrt{T_d} \sim 1/\sqrt{f_{\text{samp}}}$

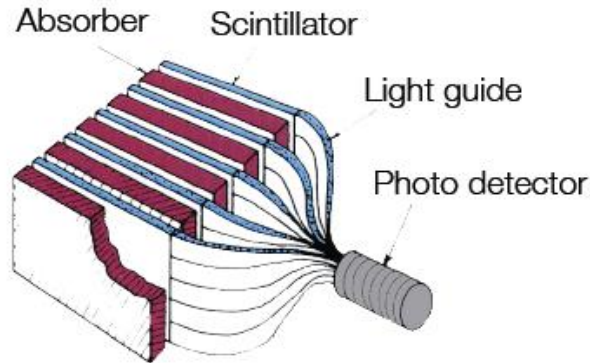


$$\frac{\sigma_{\text{samp}}}{E} = \frac{2.7\%}{\sqrt{E [\text{GeV}]}} \sqrt{\frac{s [\text{mm}]}{f_{\text{samp}}}}$$

$$\frac{\sigma E}{E} = 3.2\% \sqrt{\frac{E_c [\text{MeV}] \cdot t_{\text{abs}}}{F \cdot E [\text{GeV}]}}$$

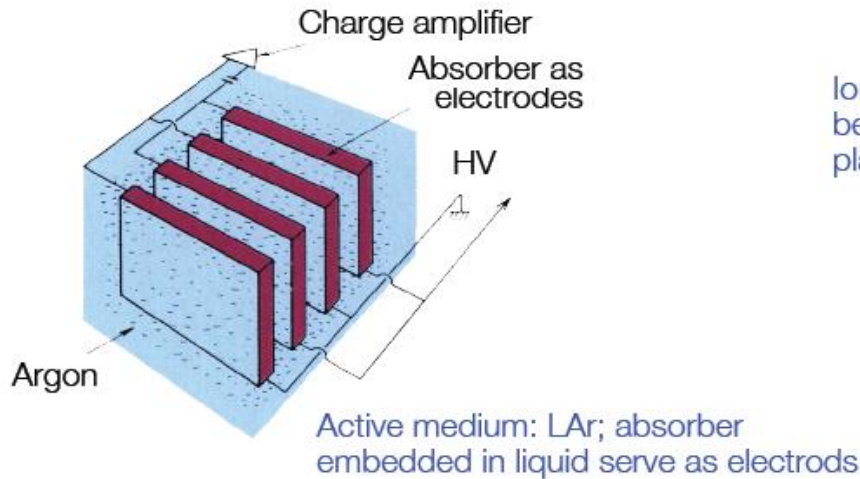
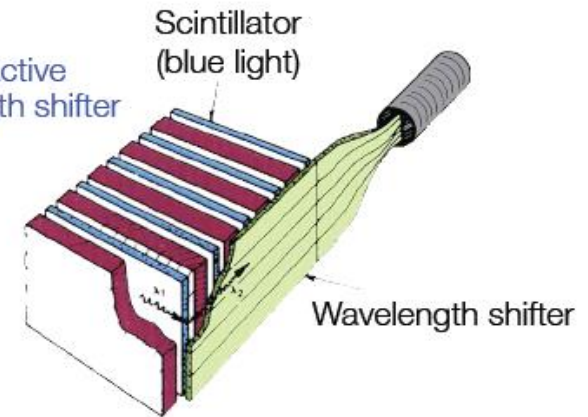
# EM sampling calorimeter examples

Scintillators as active layer;  
signal readout via photo multipliers

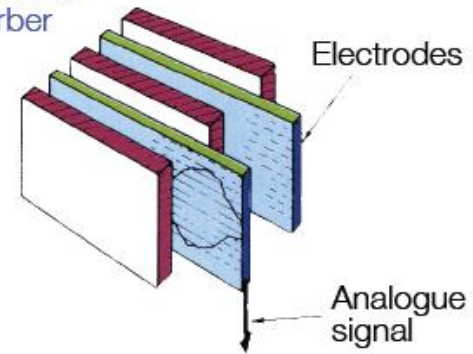


Possible setups

Scintillators as active layer; wave length shifter to convert light



Ionization chambers  
between absorber  
plates



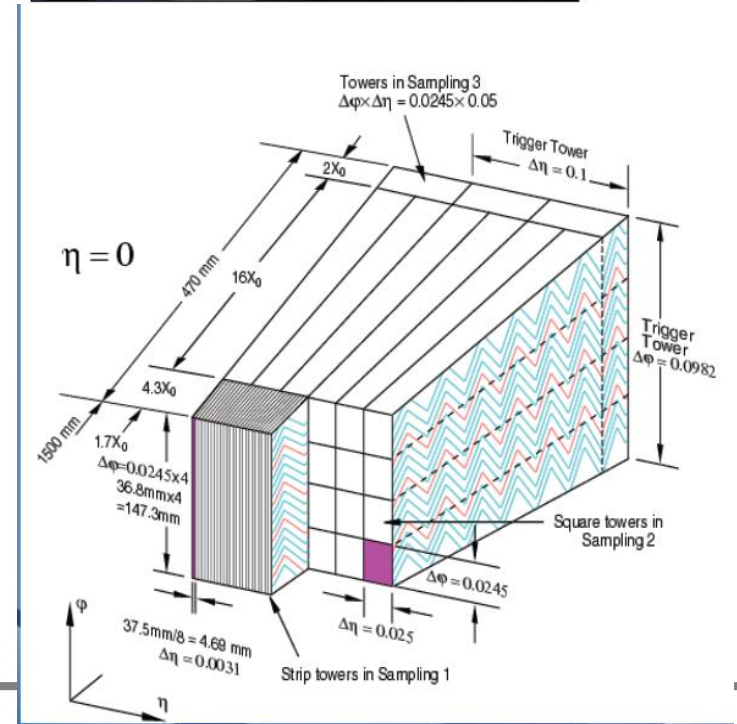
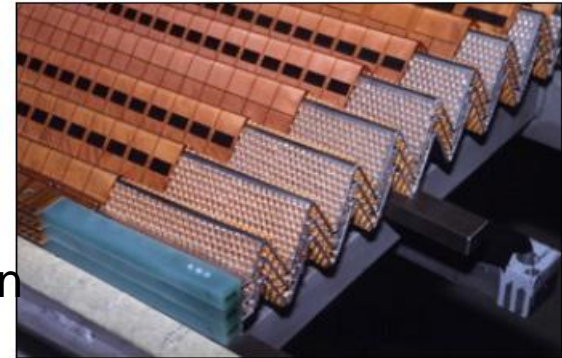
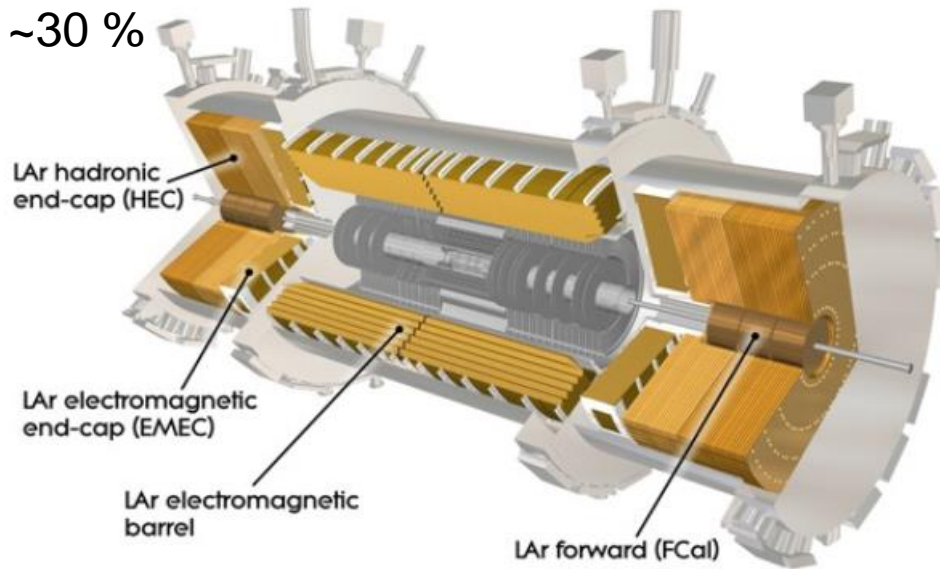


# ATLAS Lar EM Calorimeter

Example:  
ATLAS Liquid Argon Calorimeter

Main optimisation : constant term and  $\gamma/\pi^0$  separation

$f_{\text{samp}} \sim 30\%$



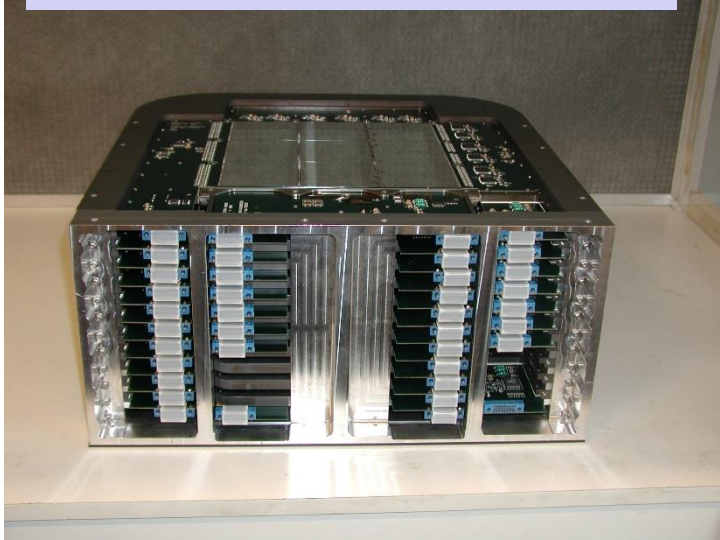
# Trigger parenthesis

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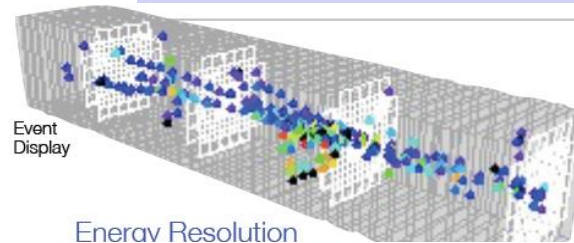
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# Si/W calorimeter : Calice & Pamela

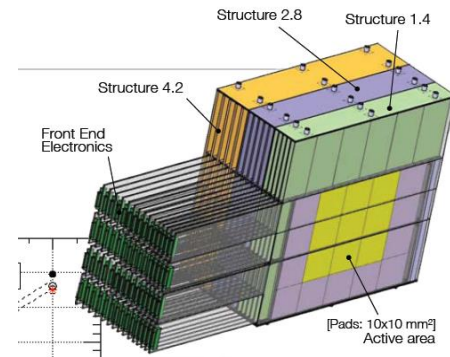
Pamela calorimeter (satellite cosmic rays experiment)



Calice : prototype for ILC detectors



ECAL 'Physics' Prototype  
[CALICE]



Compact calorimeter :

- 24x24cm transversal
- 16.3  $X_0$  depth, > 20layers

Topological reconstruction of shower  
 $e/\pi$  separation and position measurement,  
poor energy resolution

Optimized for shower separation (lateral and Longitudinal) especially for jets and Particle Flow Analysis

***e/γ***

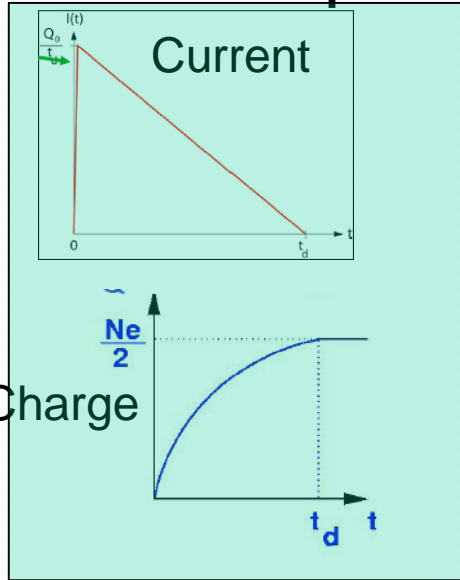
**Reconstruction**

**Calibration**

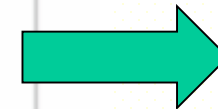
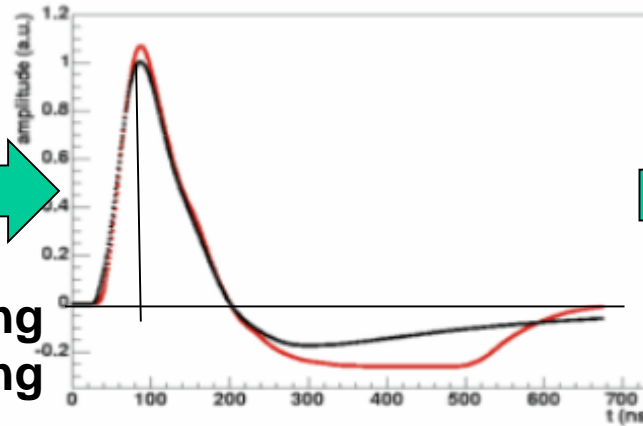
**Performance**

# From signal cell to cell energy (1)

## Detector output



**Filtering  
Shaping**



**Digitisation**

Amplitude  
digitised on  
 $n$  bits : adc counts  
12 bits : 0-4096  
(when no signal,  
measure pedestal)

$$E_{\text{cell}} = F(\mu\text{A} \rightarrow \text{MeV}) \times F(\text{adc} \rightarrow \mu\text{A}) \times (\text{adc-ped})$$

$F(\text{adc} \rightarrow \mu\text{A})$  : take into account electronics chain gain.

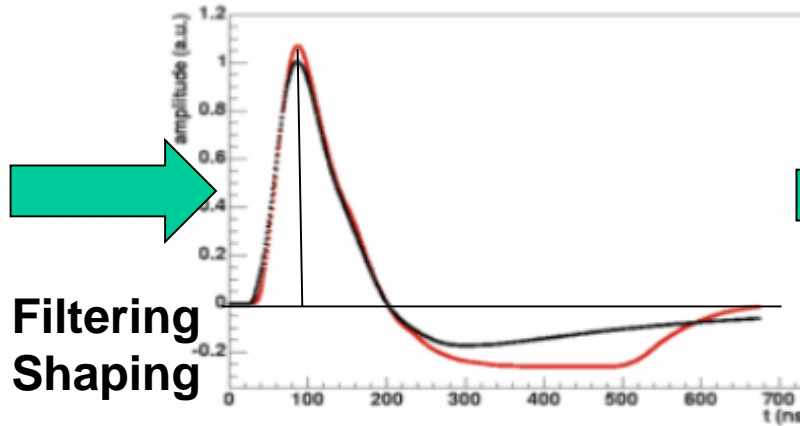
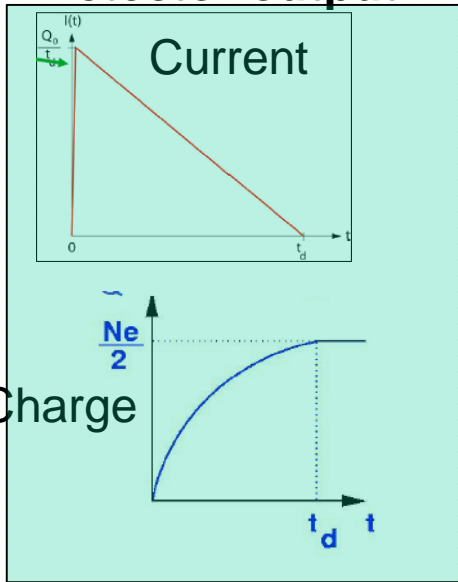
Calibration system can be laser signal in a crystal or inject charge at detector output as similar as possible to signal (but residual bias !)

Measure or correction for linearity. To be done for all channels !

Stability measurement with time / temperature .....

# From signal cell to cell energy (2)

## Detector output



Filtering  
Shaping

Digitisation

Amplitude  
Coding on  
n bits : adc  
(when no signal,  
Measure pedestal

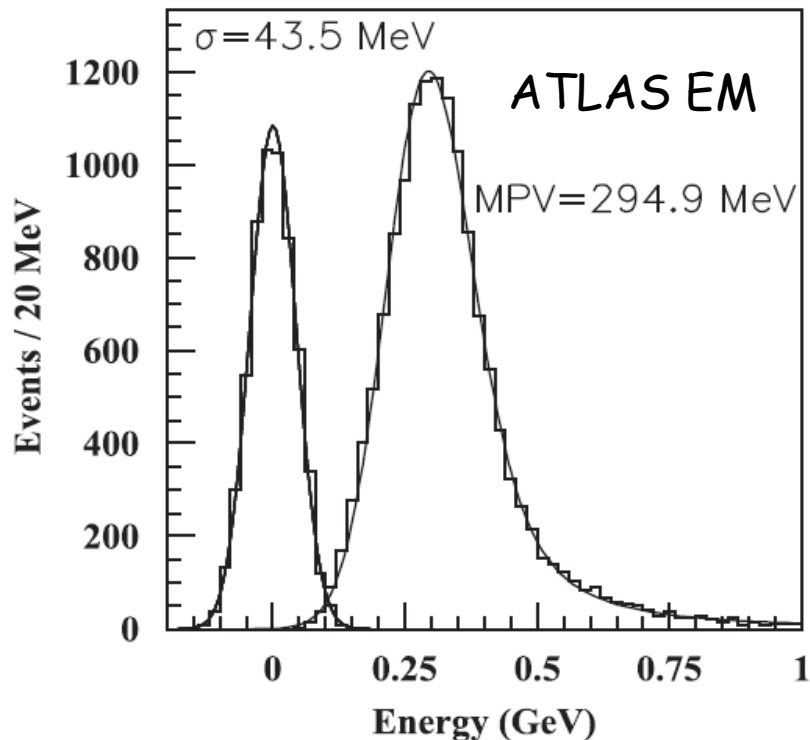
$$E_{\text{cell}} = F(\mu\text{A} \rightarrow \text{MeV}) \times F(\text{adc} \rightarrow \mu\text{A}) \times (\text{adc-ped})$$

$F(\mu\text{A} \rightarrow \text{MeV})$  : Can be computed from first principles to 5-10% but not enough accurate (for sampling calorimeters includes sampling fraction)  
Usually extracted from beam test with prototype by shooting and reconstructing particles of well know energy  
Still not accurate ultimately.....

# $\mu$ (mip) signal in calorimeter

Muons will not produce showers in calorimeter but deposit Minimum Ionizing Particle energy (dE/dx at minimum)

- Can be used for rough calibration / Inter calibration / time dependence
- Difficult to extract absolute EM energy scale as  $e/\mu$  for mip  $\neq 1$
- Landau spectrum with high energy tail, characterized by Most Probable Value
- Useful quantify is  $S/N = MPV/\sigma$  to qualify electronics readout/noise



# Shower energy reconstruction

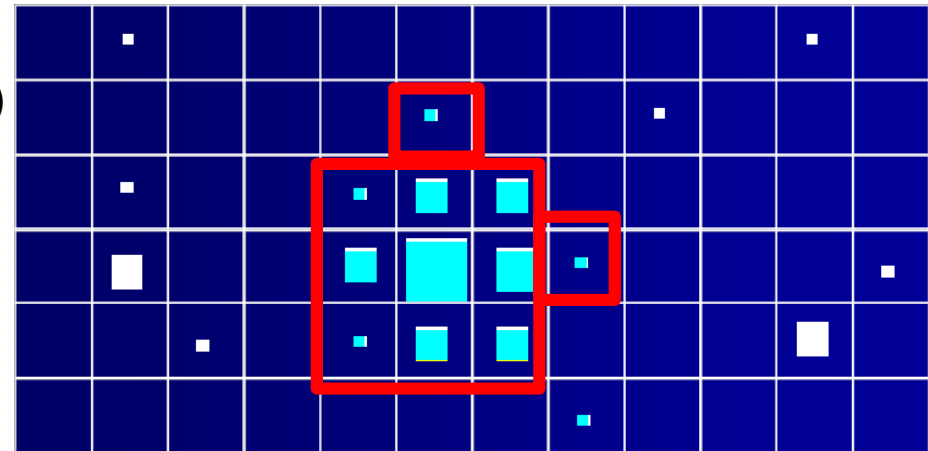
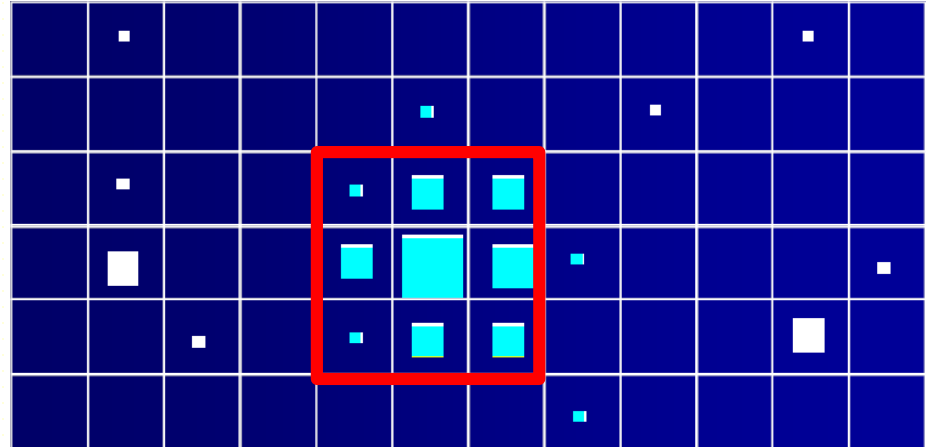
## Fixed cluster size :

- Large enough to contains >95 % of EM shower energy
- Small enough to minimize noise and shower separation

$$(\sigma_{\text{noise}} = N_{\text{cells}} \sigma_{\text{inco}} \oplus N_{\text{cells}}^2 \sigma_{\text{coh}})$$

Fast and easy algorithm

Em shower + noise (or other particle)



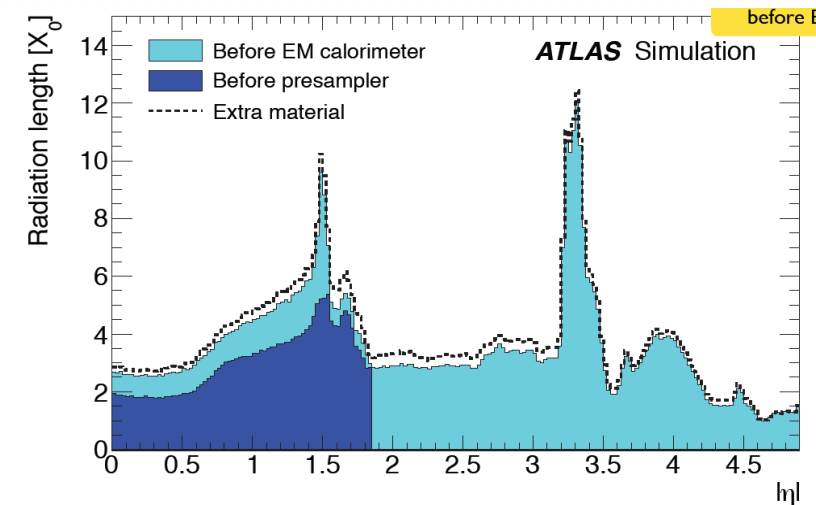
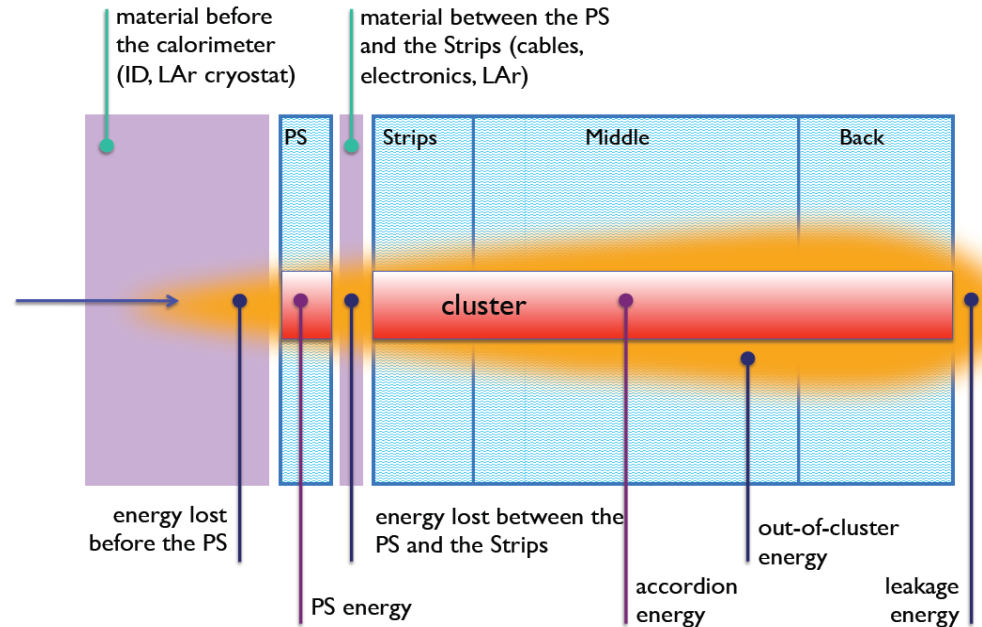
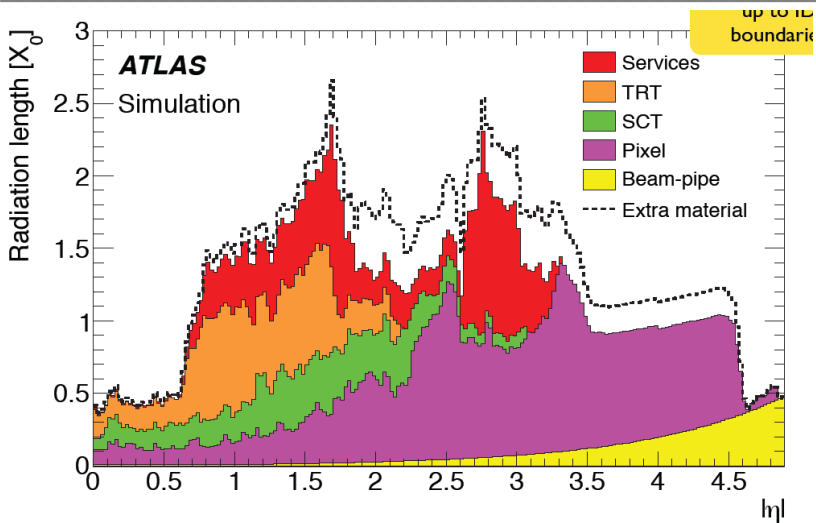
## Topological algorithm :

- Consider all cells with  $E > E_{\text{cut}}$  ( $3 \sigma$  noise)
- Start from a seed (max)
- Add neighbour cell if  $E > E_{\text{cut}}$

Iterative process. Can achieve same energy resolution but more difficult for linearity (calibration) and noise contribution (different from one shower to another)



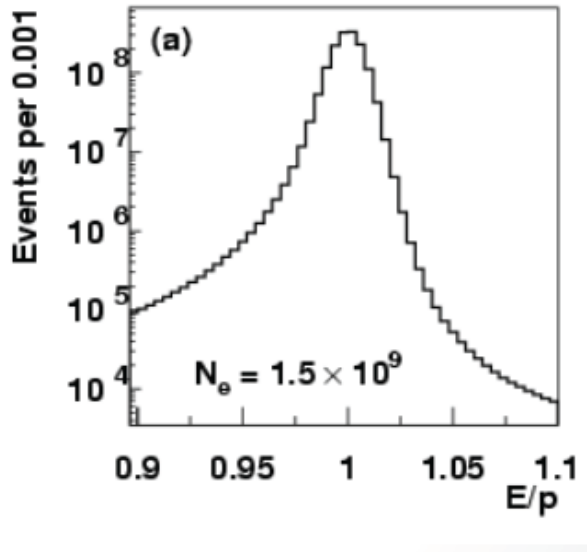
# From cluster to particle energy



Energy lost upstream, laterally and longitudinally for any calorimeter  
+presampler for ATLAS

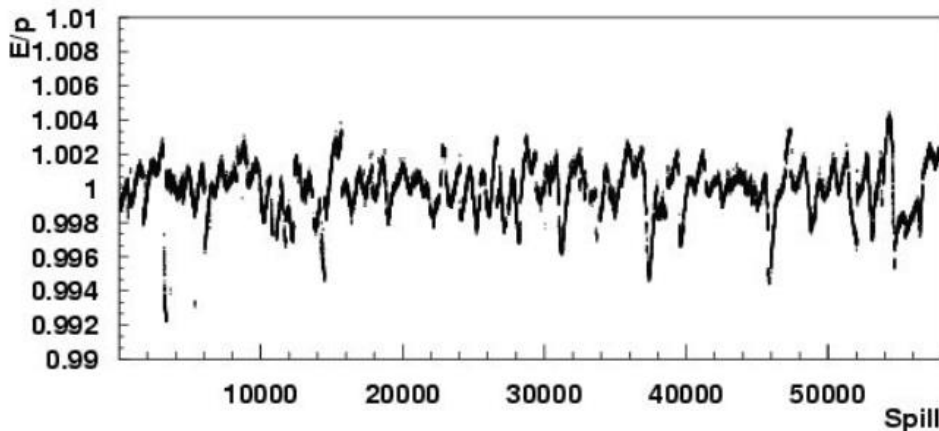
# In situ particle energy calibration

- Can use  $E(\text{cal})/p(\text{tracker})$  if material upstream uniform and not large



Example of KTeV CsI calorimeter :  
use electrons from  $K_L \rightarrow \pi e \nu$   
Set absolute energy scale  
Crystal to crystal calibration  
Time dependence of signals

Quite difficult at LHC with material variation along  $\eta \rightarrow E/p$  distribution with too many tails



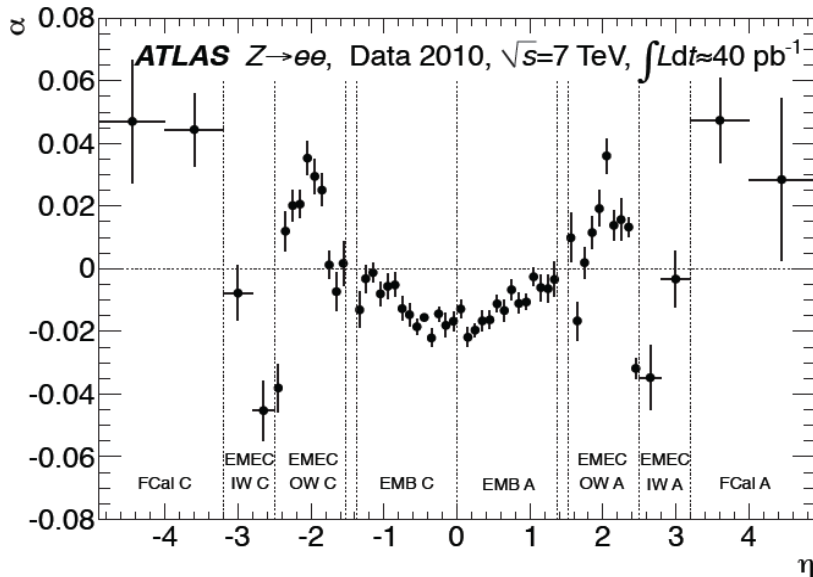
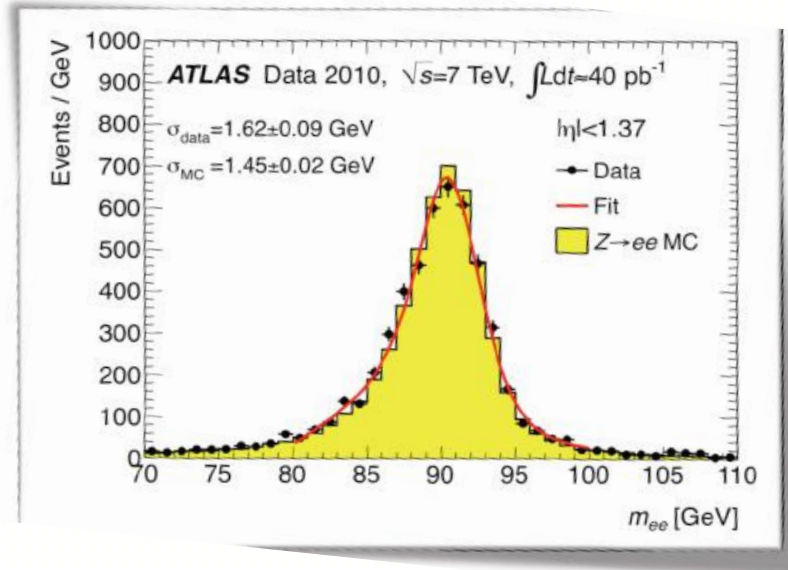
# In situ particle energy calibration (3)

- Use mass constraint on well known particle :  $Z \rightarrow e+e-$  @ LHC

$$m = \sqrt{2E_1 E_2 (1 - \cos(\theta_{12}))}$$

$$E^{\text{corr}} = E (1 + \alpha_i)$$

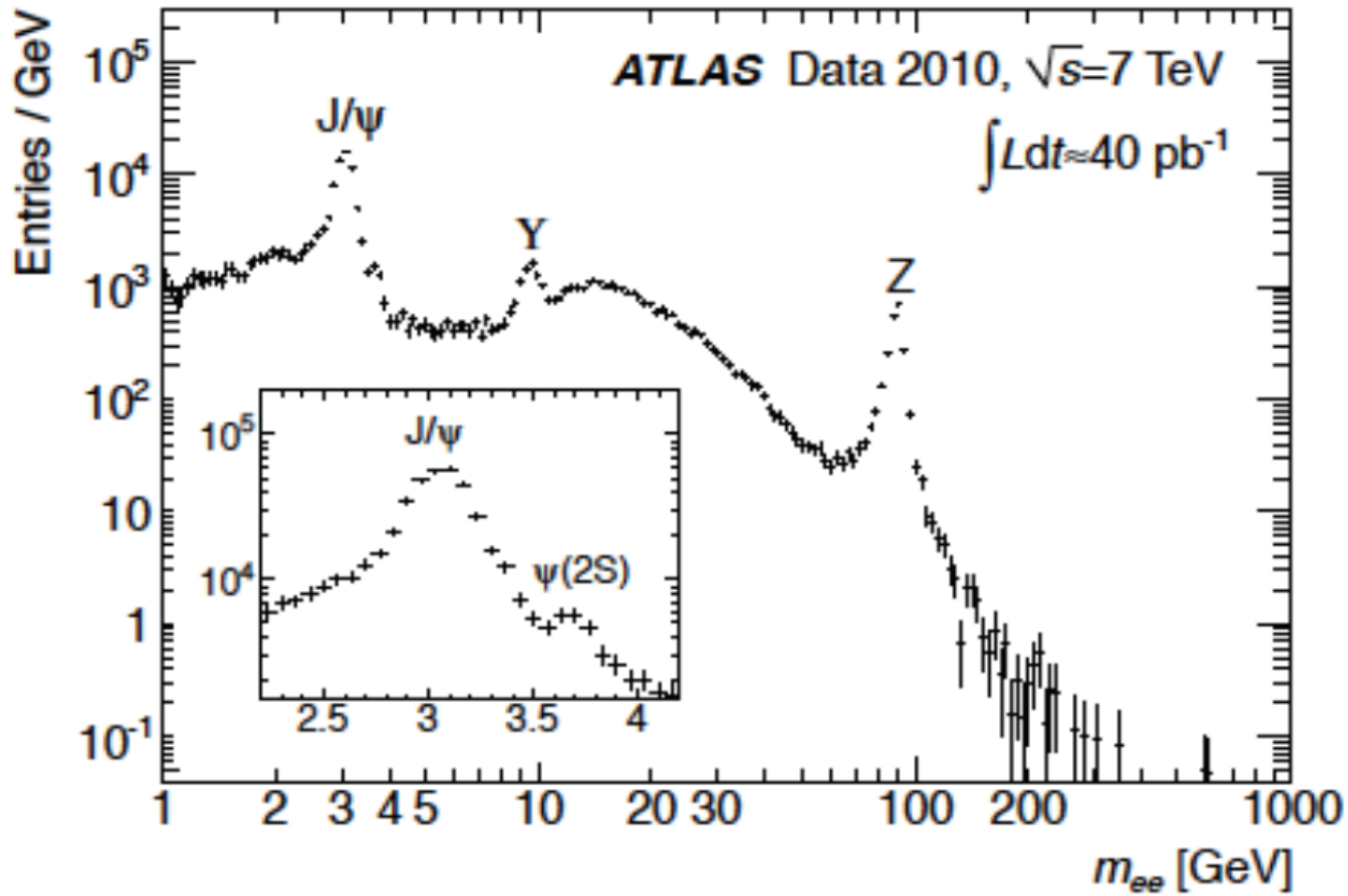
$$m_{ij}^{\text{corr}} \simeq m_{ij} \left( 1 + \frac{\alpha_i + \alpha_j}{2} \right)$$



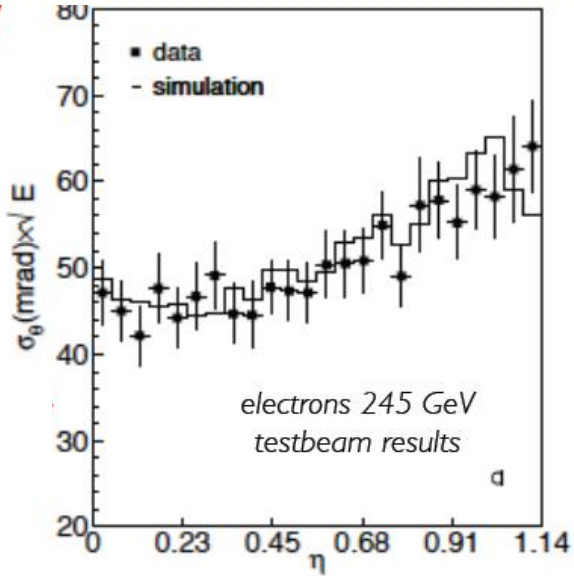
With more stat :

- Reduce region for each  $\alpha$  + along  $\varphi$
- Can use  $J/\psi \rightarrow e+e-$  for low energy (linearity)
- Needs to extrapolate  $\gamma$  from simulation or  $Z \rightarrow ee\gamma/\mu\mu\gamma$

# $e^+e^-$ resonances



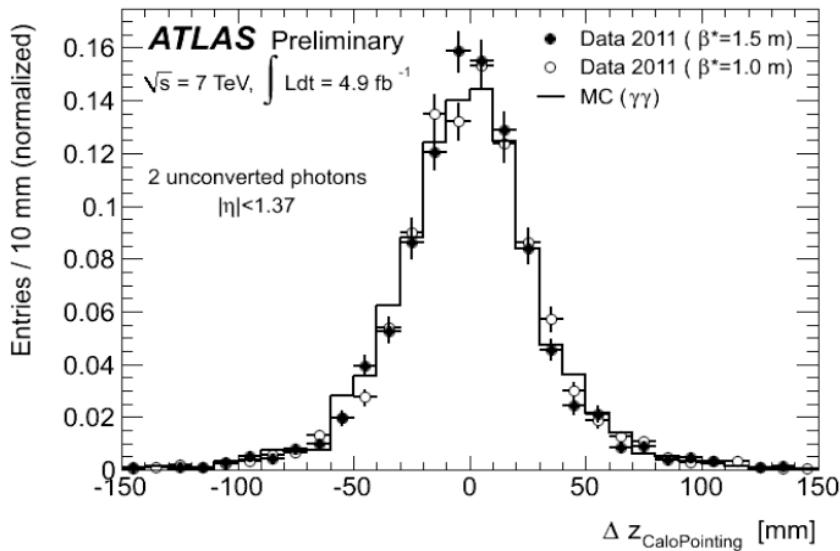
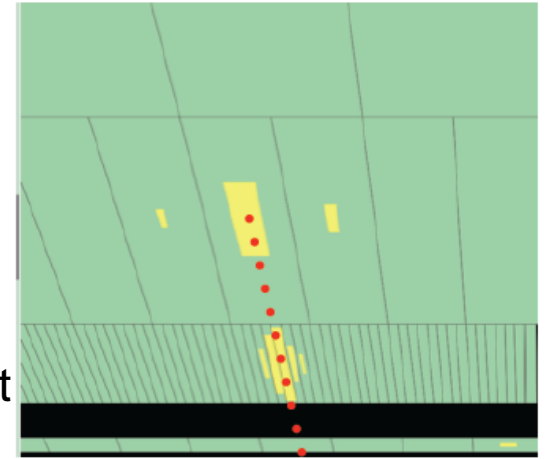
# Photon pointing in ATLAS



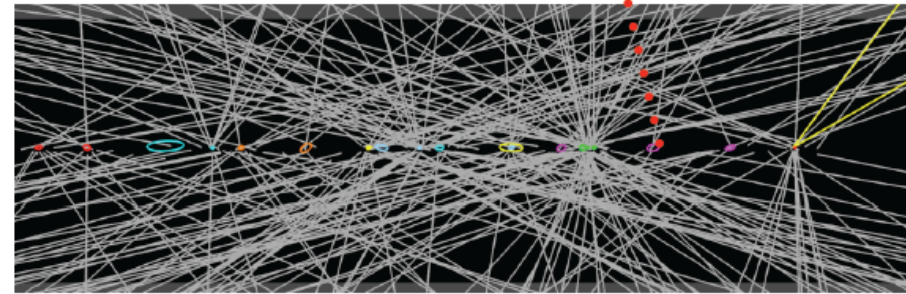
Barycentre in strips  
and middle section  
→ Angular measurement

$$\sigma_\theta = 60 \text{ mrad}/\sqrt{E} [\text{GeV}]$$

Z vertex position measurement



(not to scale)



Zoom on collision interaction

# Timing resolution

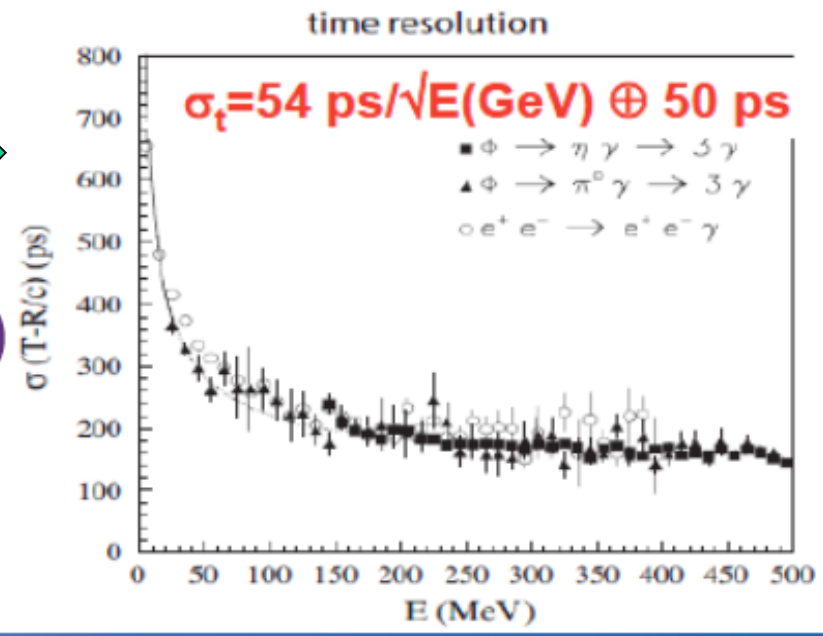
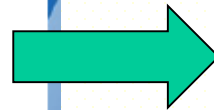
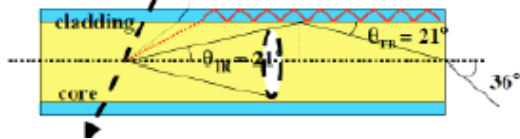
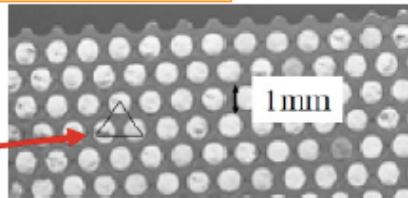
Shower (electron/photon) time measurement can also be achieved, can be useful to reject out of time events (accidentals) with respect to collision

KLOE calorimeter

## Fine sampling lead/scintillating fibers calorimeter

- Volume Ratio Fiber:Lead 50:50
- Energy sampling fraction 13 %
- $X_0 = 1.6 \text{ cm}$   $\rho = 5.3 \text{ g/cm}^3$

Triangular shape:  
high sampling  
frequency &  
flat response in  $\theta$



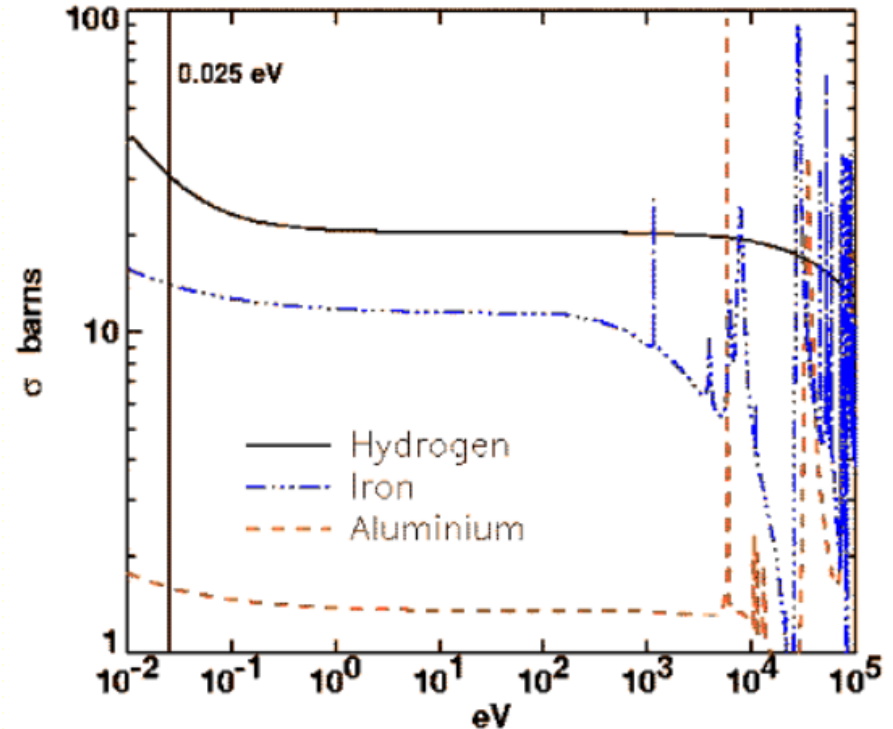
Time measurement in calorimeter, promising way to mitigate "in time" pile-up at LHC

# neutrons

Table 12-1. Average number of collisions required to reduce a neutron's energy from 2 MeV to 0.025 eV by elastic scattering

Element	Atomic Weight	Number of Collisions
Hydrogen	1	27
Deuterium	2	31
Helium	4	48
Beryllium	9	92
Carbon	12	119
Uranium	238	2175

## Neutron Cross Sections

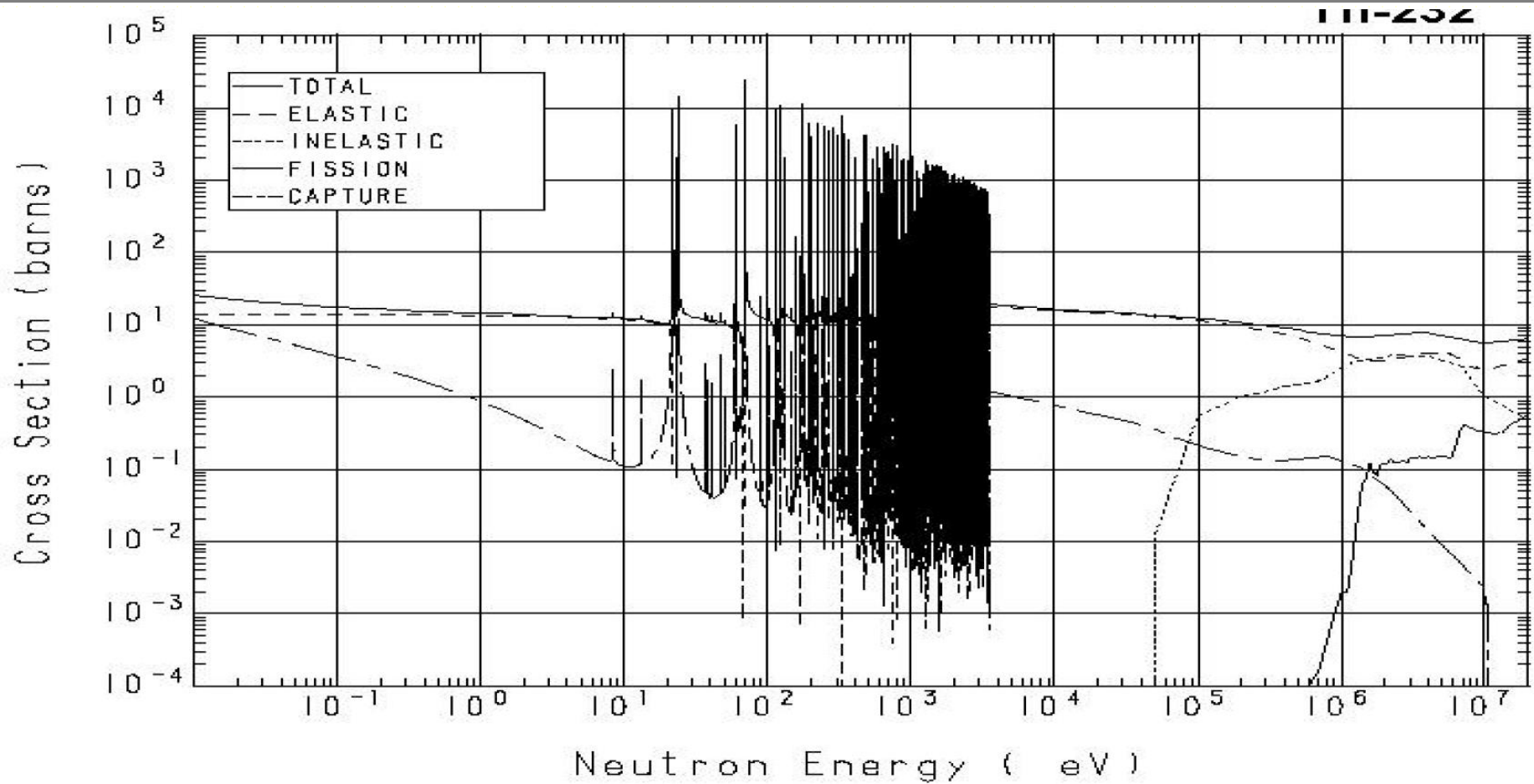


Capture may be increased at some energy due to resonance effect in the total cross section...

T (K)	$E_0$ (eV)	$v_0$ (m s <sup>-1</sup> )	$\bar{E}_{th}$ (eV)
300	0,0253	2200	0,038
400	0,034	2600	0,051
600	0,052	3100	0,075
800	0,069	3600	0,103
1000	0,086	4000	0,129



# Cross section on Th-232



# Consequences on LHC experiment hall

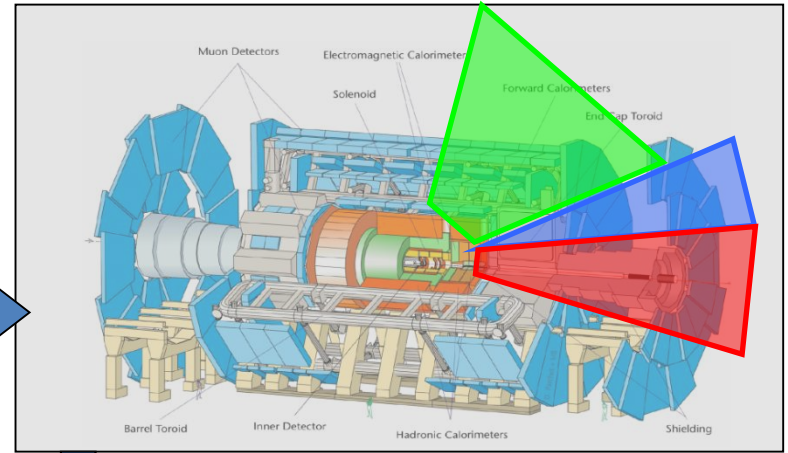
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# Background in Atlas cavern

Background comes from residues of p-p interactions (through spallation process):

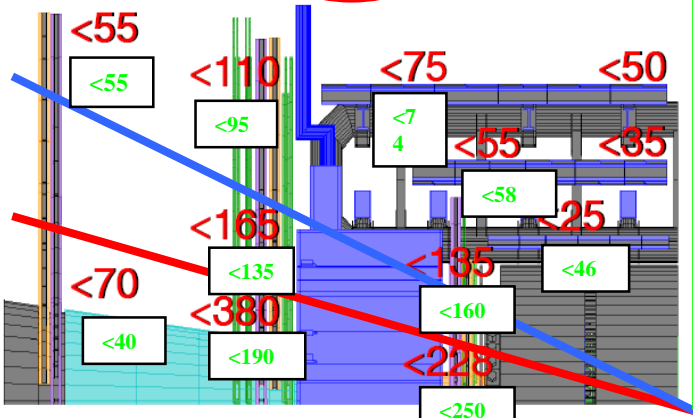
- Huge production of *neutrons*, thus creating  $\gamma$ , thus creating  $e$ , etc...
- Also at higher energy,  $n$  and  $\gamma$  create ionizing particle (mainly:  $p$ ,  $e^+$ ,  $e^-$ )
- Direct background:  $\mu$  and punchthrough (smaller)



Neutron "gas" in the cavern (-> therm. of neutrons)

$E_{cin}=10$  MeV :  $v_n \sim 15\% c$  (23 ns pour 1m)  
 10 keV :  $\sim 5 \cdot 10^{-3} c$  (0.7  $\mu s$  pour 1m), out of time

Background count rates (kHz/tube) at  $L = 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$



- (Numbers include a safety factor of 5.)
- pp x section for part. prod. (~1.2)
  - Rad. propagation in calo. and shielding (~2.9)
  - Efficiency in chambers (~1.4)

Atlas needs to measure cavern background using muon spectrometer) in order to reduce uncertainty on background for s-LHC.

$10 \text{ kHz/cm}^2$   
 $\epsilon_n \text{ hit} \sim 10^{-4} \text{ à } 10^{-3}$   
 $\epsilon_\gamma \text{ hit} \sim 4-8 \cdot 10^{-3}$

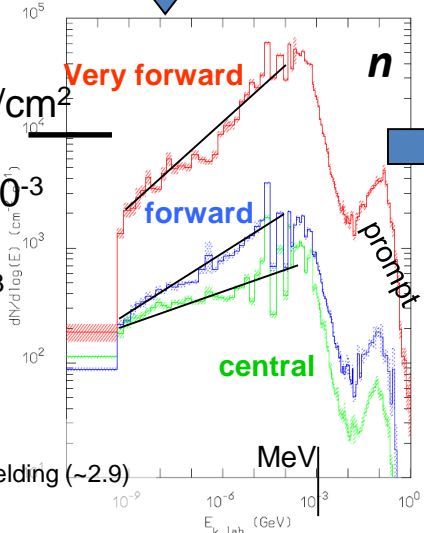


Figure 5-67 The expected neutron flux as a function of neutron energy in different rapidity regions of the muon spectrometer (top curve:  $2.3 < \eta < 2.7$ , middle curve:  $1.4 < \eta < 2.3$  and bottom curve:  $\eta < 1.4$ ).

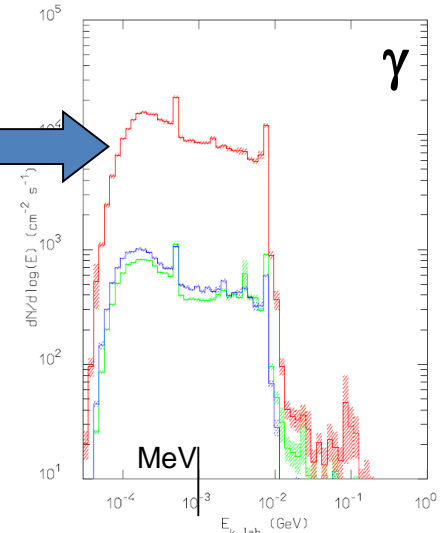
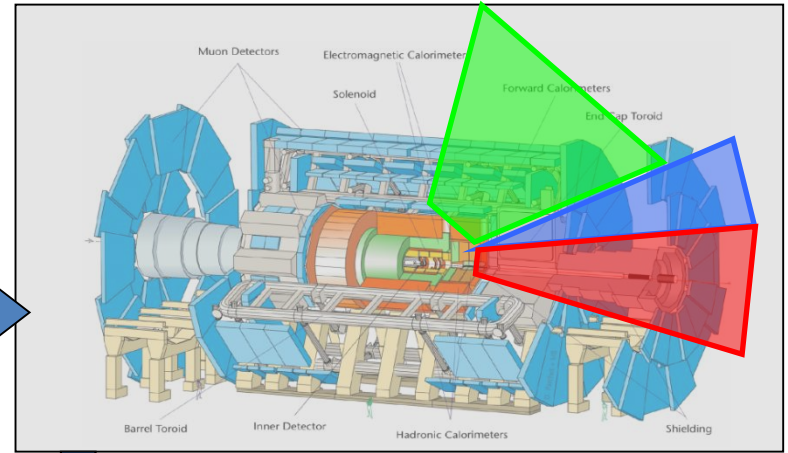


Figure 5-66 The expected photon flux as a function of photon energy in different rapidity regions of the muon spectrometer (top curve:  $2.3 < \eta < 2.7$ , middle curve:  $1.4 < \eta < 2.3$  and bottom curve:  $\eta < 1.4$ ).

# Background in Atlas cavern

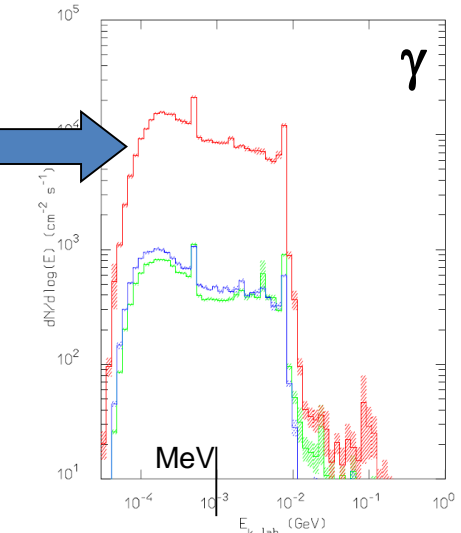
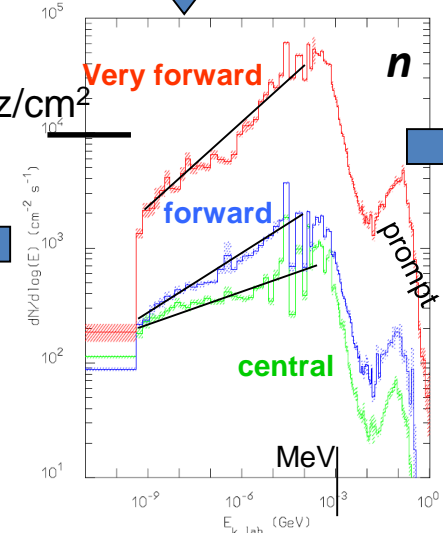
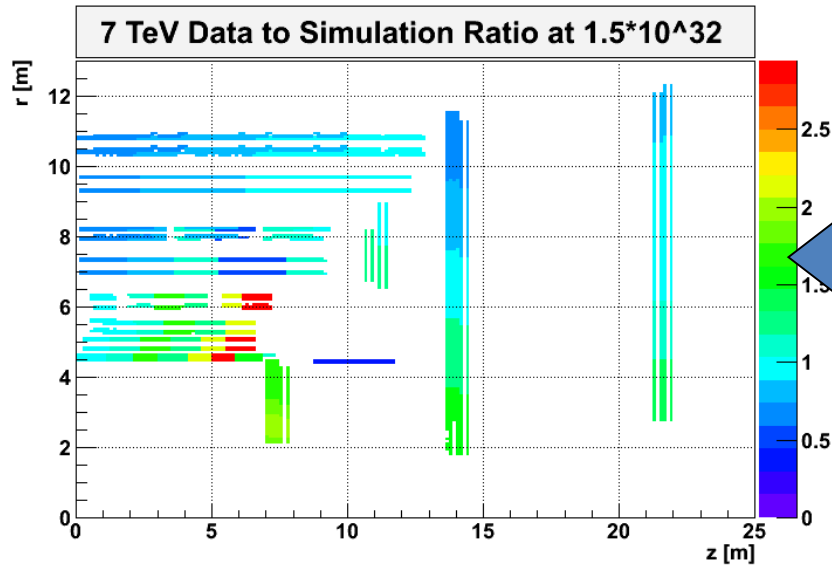
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(-> therm. of neutrons)

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**Figure 5-67** The expected neutron flux as a function of neutron energy in different rapidity regions of the muon spectrometer (top curve:  $2.3 < \eta < 2.7$ , middle curve:  $1.4 < \eta < 2.3$  and bottom curve:  $\eta < 1.4$ ).

**Figure 5-66** The expected photon flux as a function of photon energy in different rapidity regions of the muon spectrometer (top curve:  $2.3 < \eta < 2.7$ , middle curve:  $1.4 < \eta < 2.3$  and bottom curve:  $\eta < 1.4$ ).

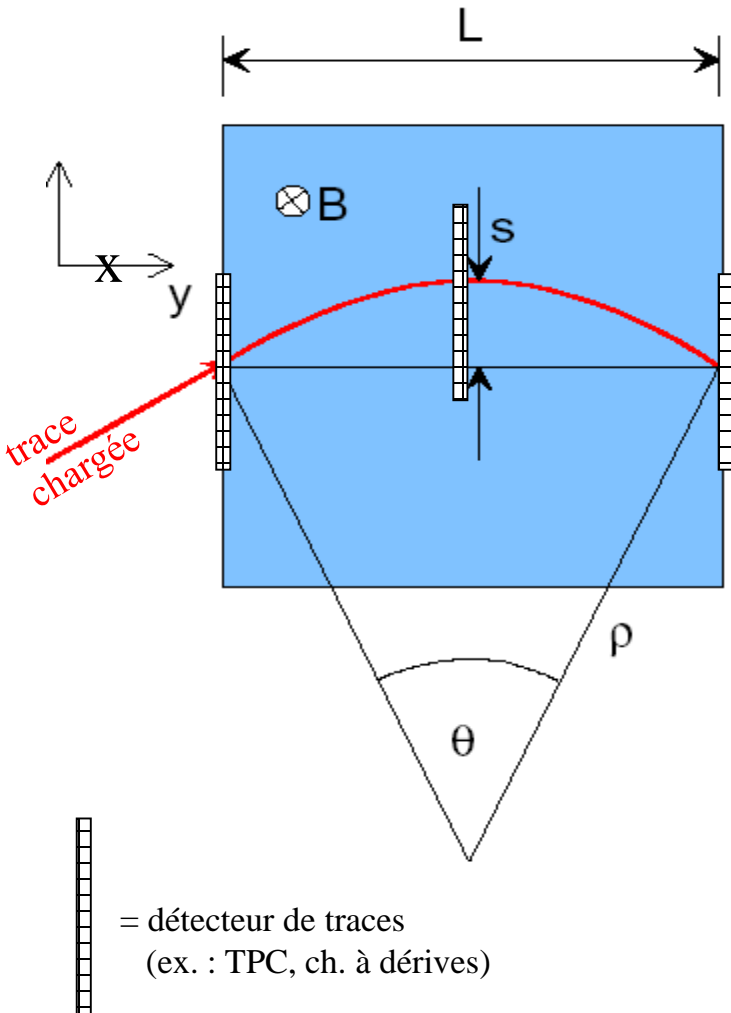
(dernière simulation : facteur de sécurité ~ x2)

# Magnets

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# Charge track momentum measurement in a magnetic field



$$p_T = qB\rho$$

$$p_T \text{ (GeV/c)} = 0.3B\rho \quad (\text{T} \cdot \text{m})$$

$$\frac{L}{2\rho} = \sin \theta/2 \approx \theta/2 \rightarrow \theta \approx \frac{0.3L \cdot B}{p_T}$$

$$\Delta p_T = p_T \sin \theta \approx 0.3L \cdot B$$

$$s = \rho(1 - \cos \theta/2) \approx \rho \frac{\theta^2}{8} \approx \frac{0.3}{8} \frac{L^2 B}{p_T}$$

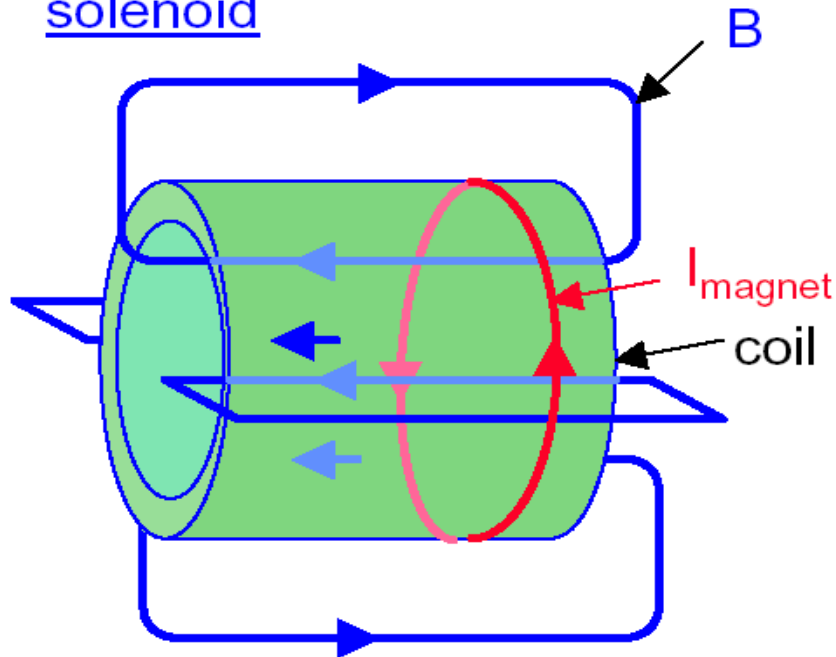
La résolution est dégradée par :

diffusion multiple (matière au milieu)

ET désalignement

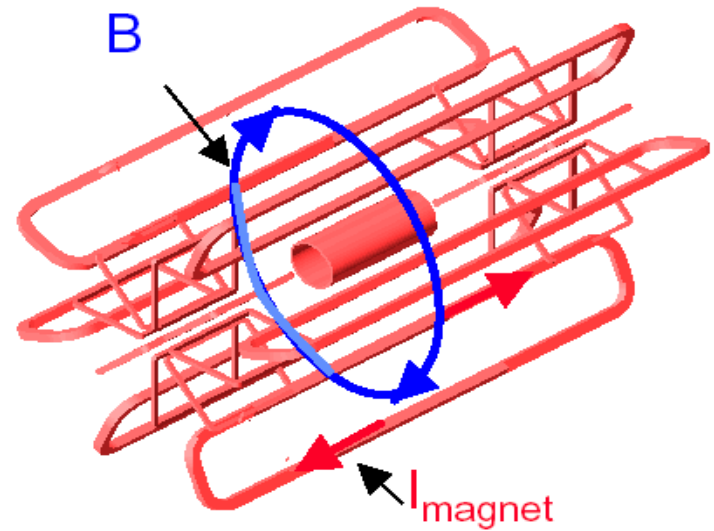
## Examples of magnetic field configuration

solenoid



- + Vertex information usefull
- + Large homogenous field inside coil
- weak opposite field in return yoke
- Size limited (cost)
- rel. high material budget

toroid

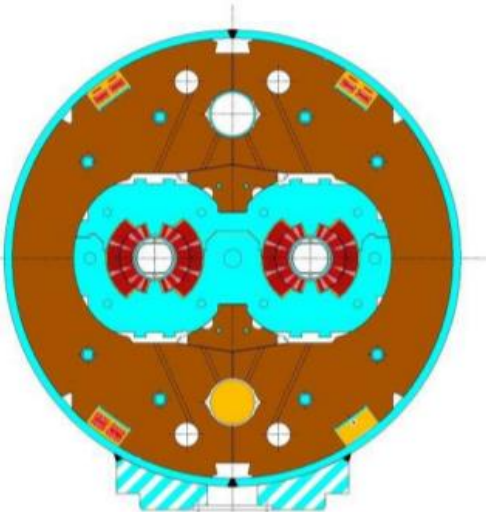


- + independant muon system (redundancy)
- + Rel. large fields over large volume
- + Rel. low material budget
- non-uniform field
- complex structure
- Vertex non-usable

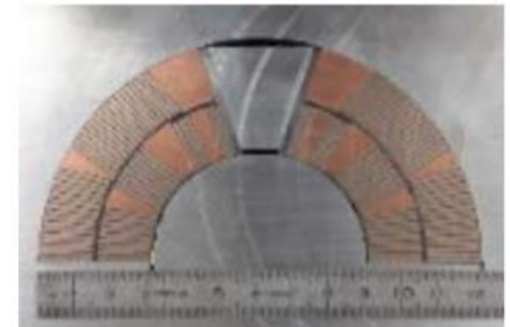
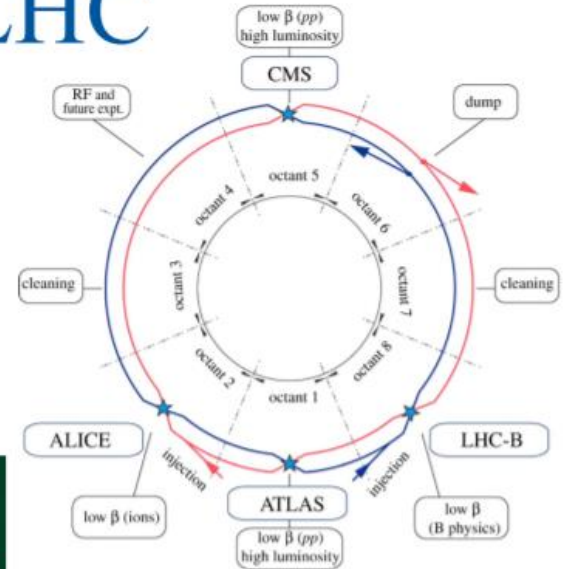
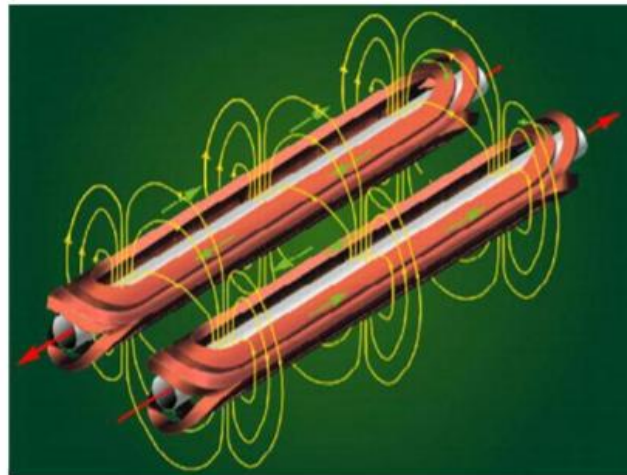
# Superconducting devices in LHC

## Magnets

- **LHC ring magnets (Nb-Ti): Rutherford cables**
  - 1232 main dipoles: 8.3 T x 15 m
  - 392 Main quadrupoles 223 T/m (7 T) x 4 m
  - 7600 other SC magnets (cable or wire)

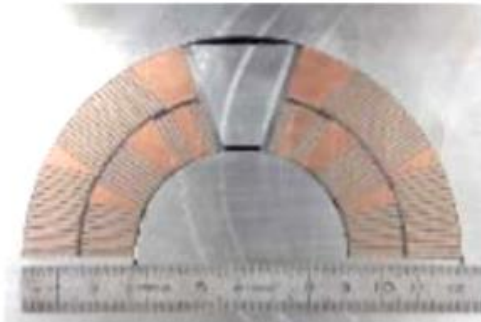


- RF cavities (Nb coating)



➤ **Rutherford Nb-Ti cable: a key technology for LHC**

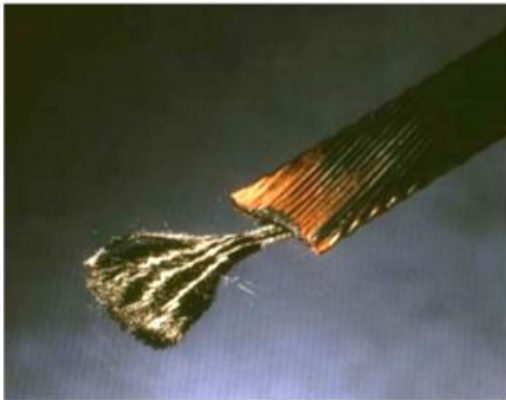




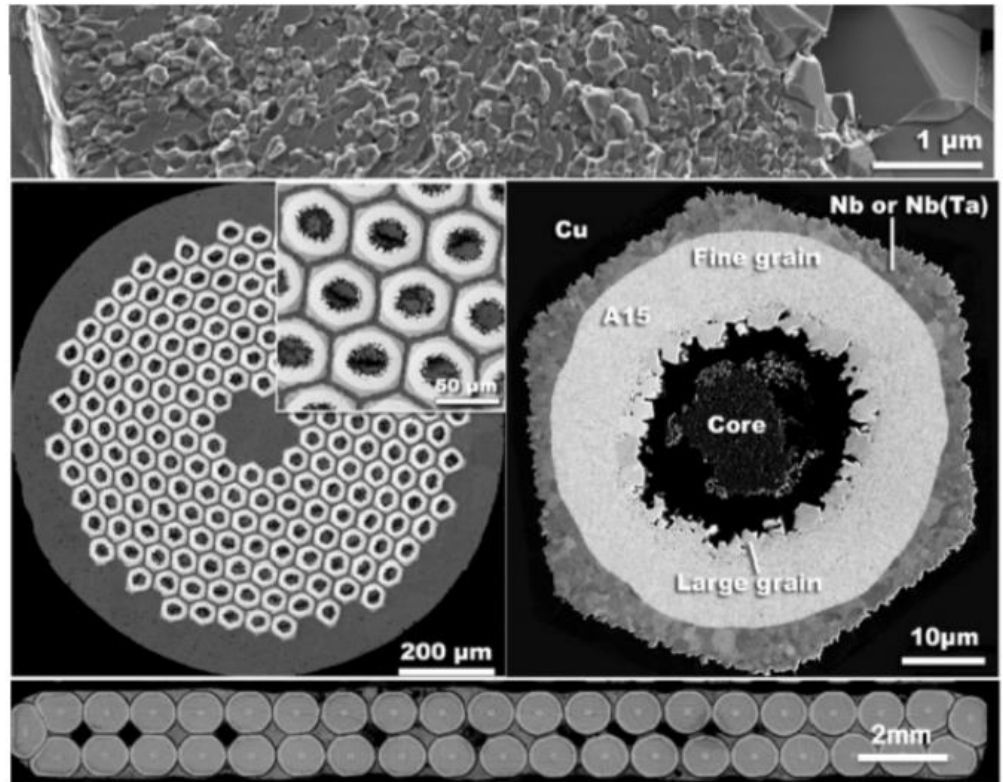
O

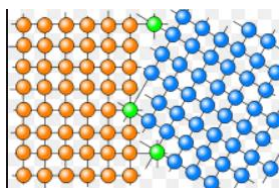
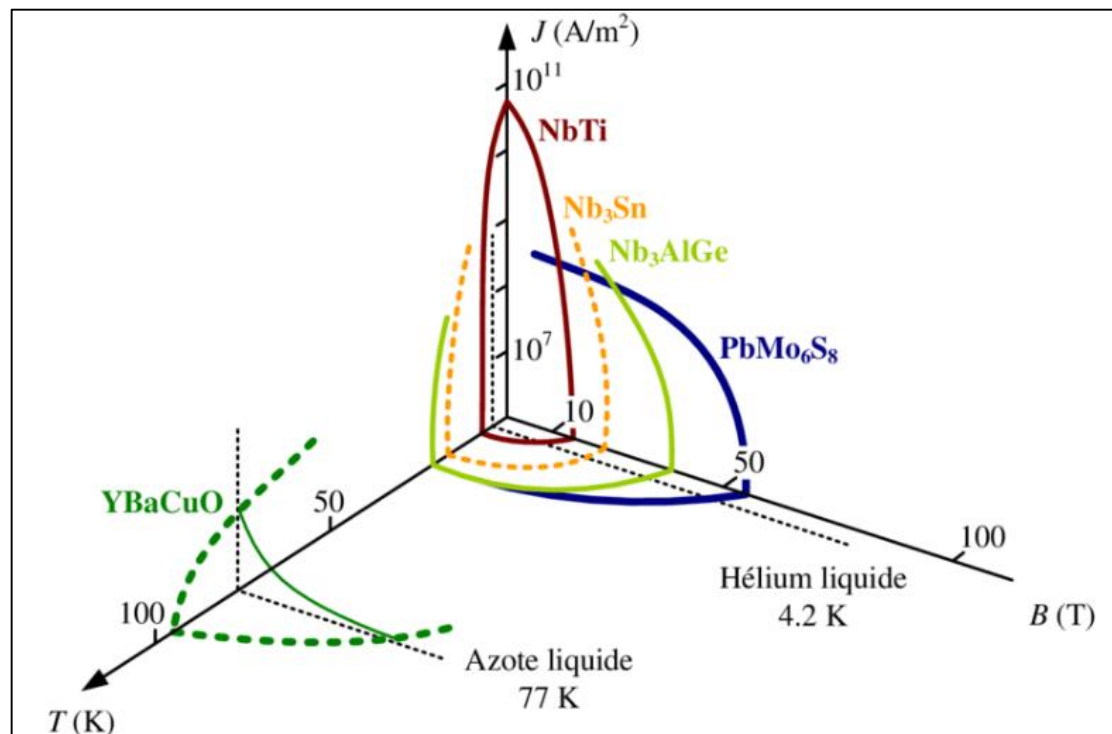
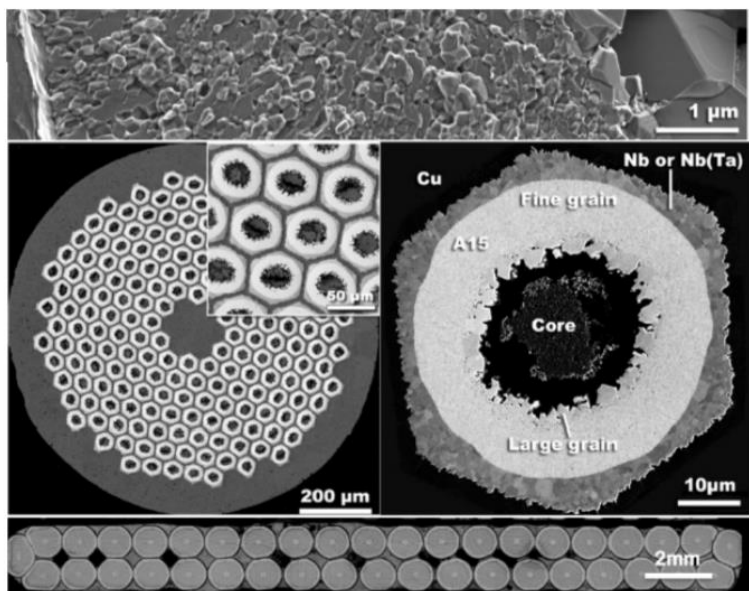
	Tevatron	HERA	RHIC	LHC
Dipole field	4.4 T	5.3 T	3.5 T	8.3 T
Number of strands	23	24	30	28-36
Cable current	4 kA	5.5 kA	5 kA	11.8 kA

cables



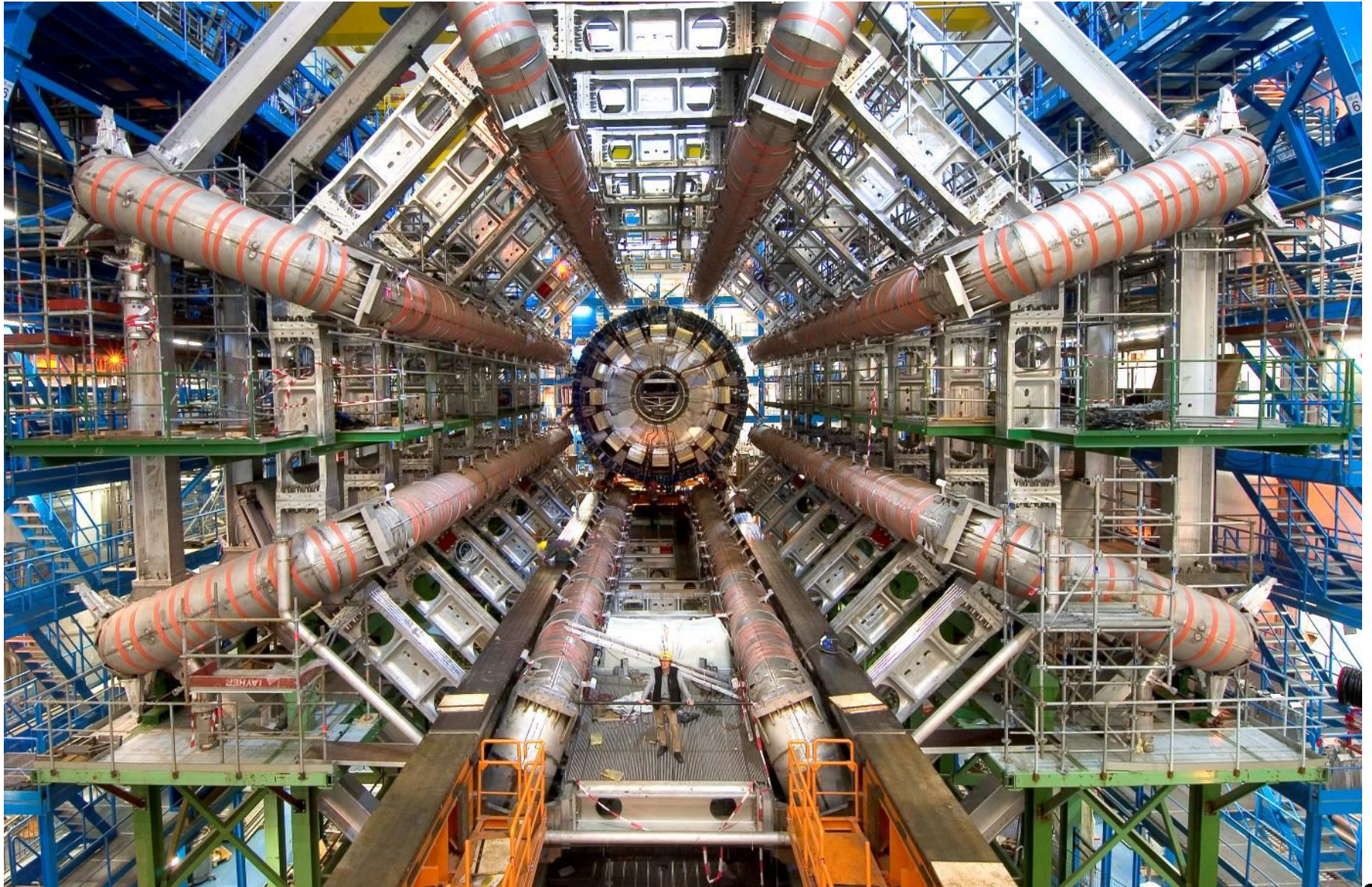
ess)

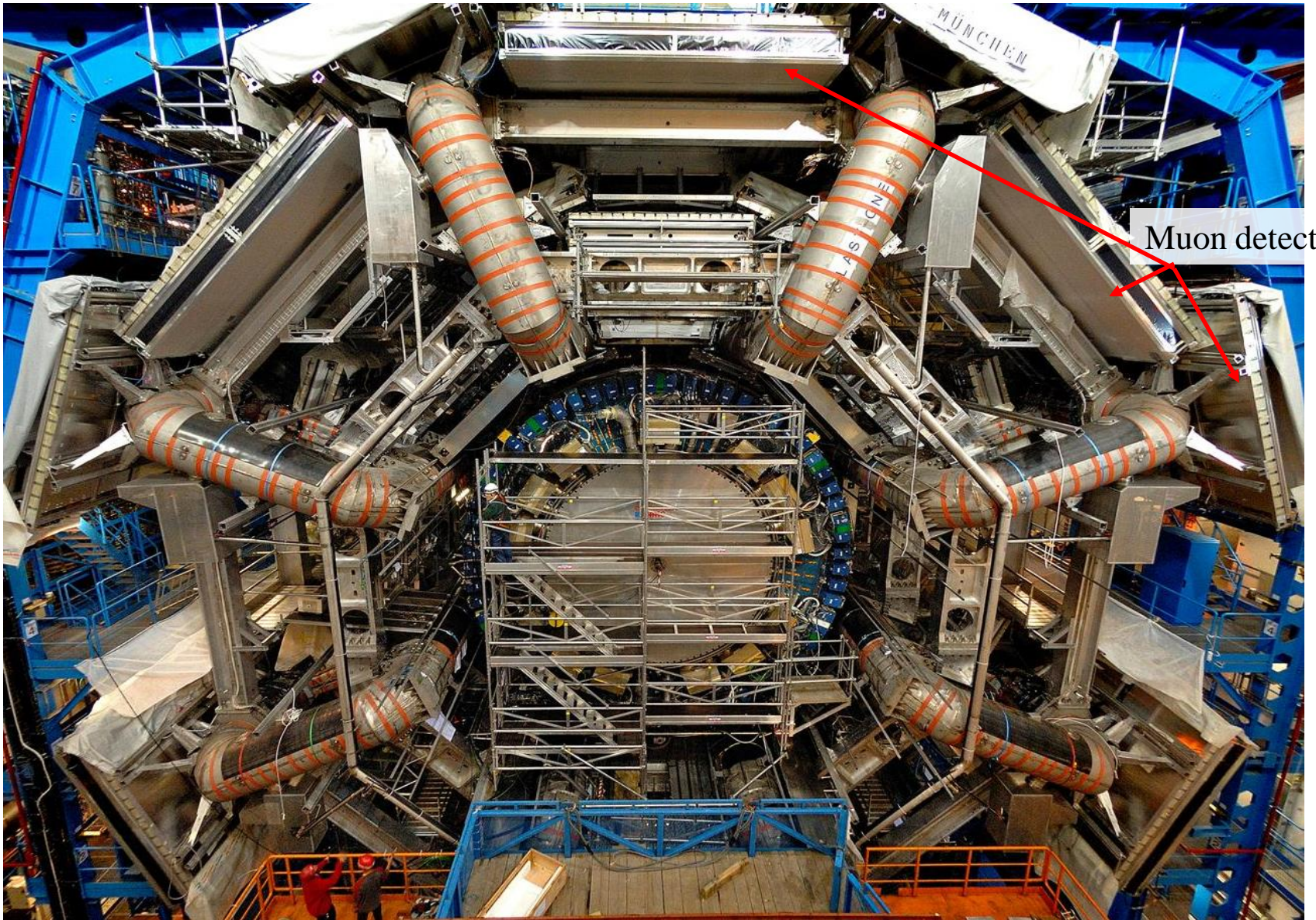




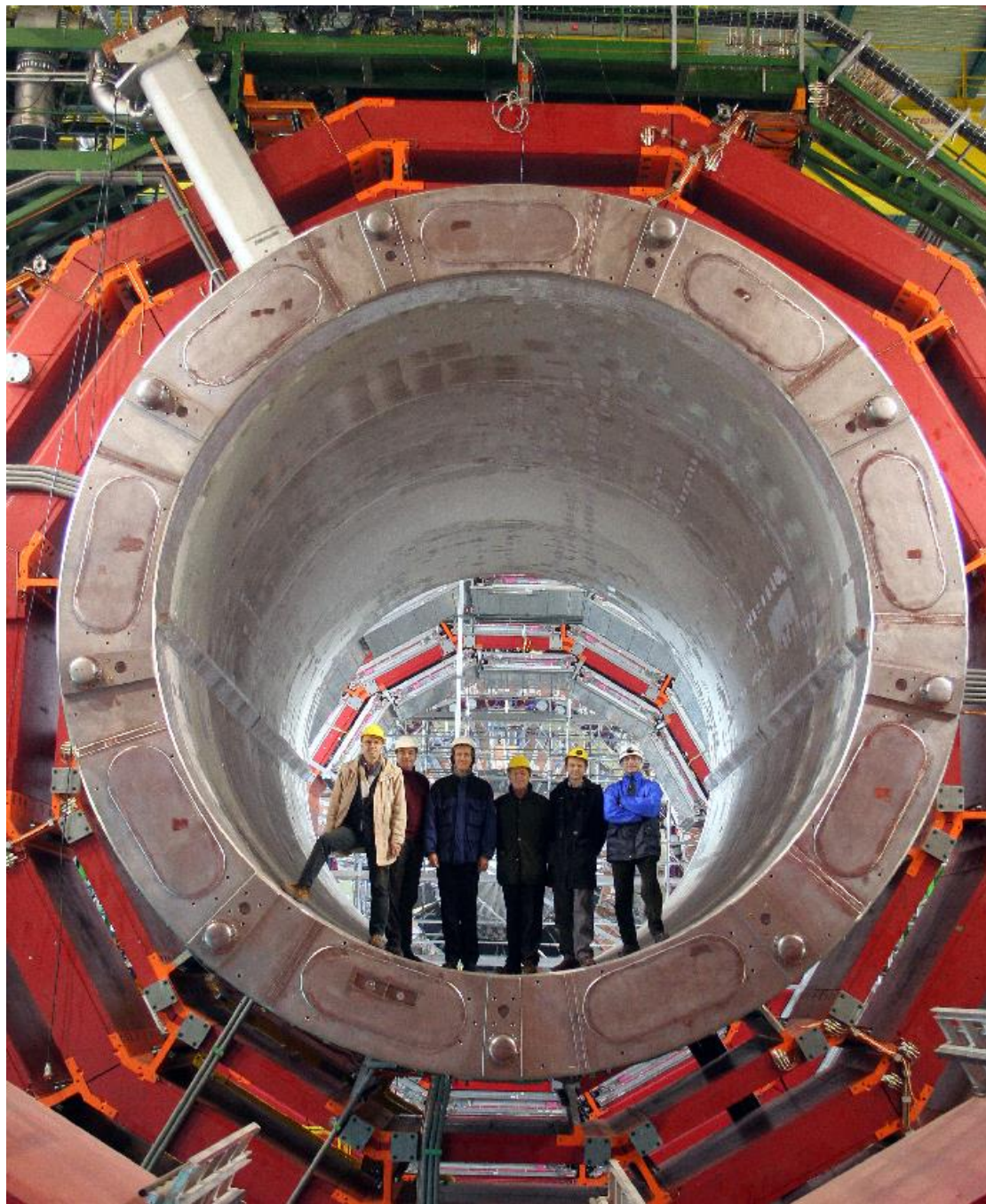
Critical current limited by density of cable defects  
(grain boundary / joint de grain)

# Magnetic fields : supraconducting magnets of ATLAS



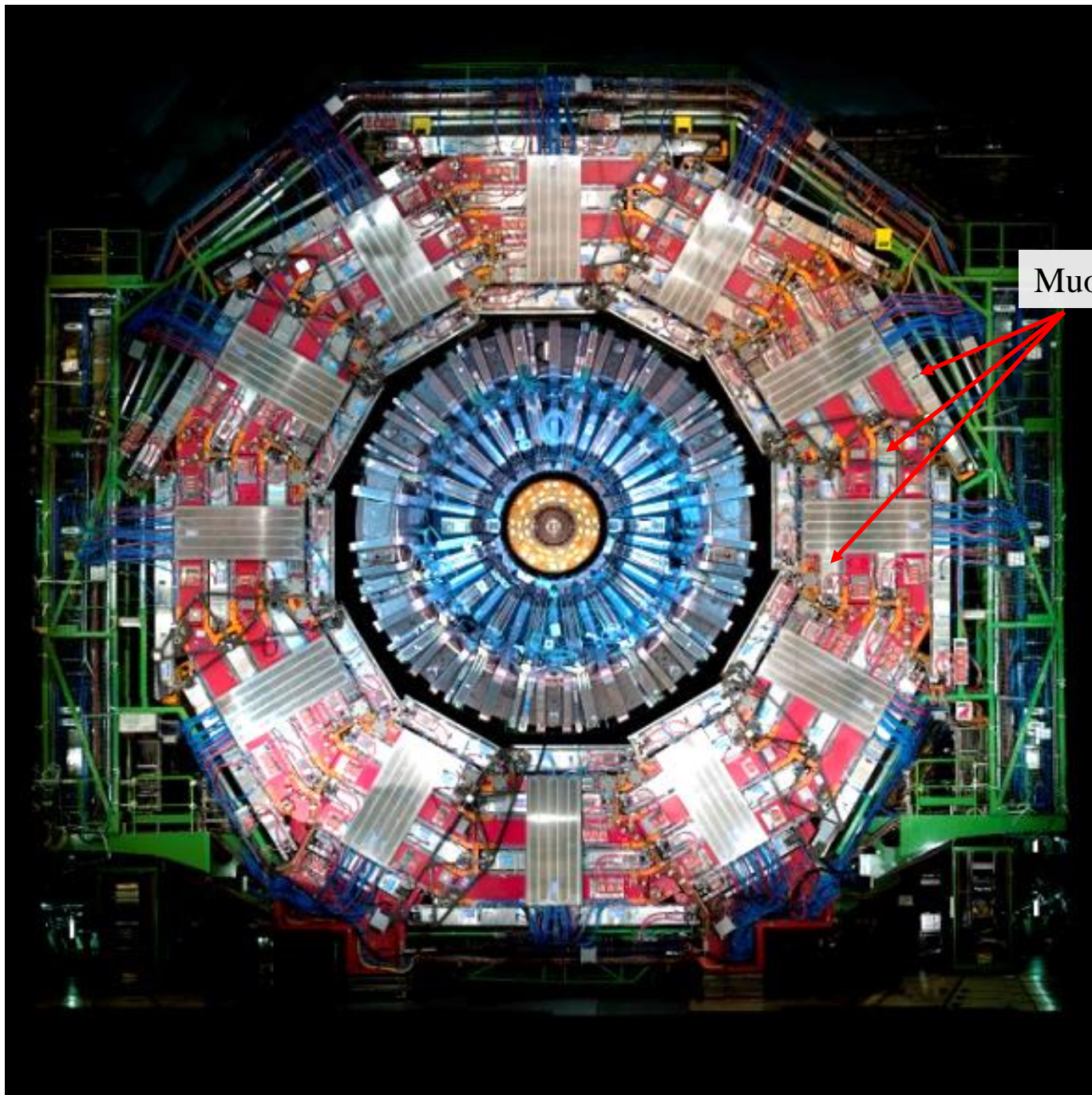


Champs magnétiques possibles :  
aimant solénoïdale de CMS



Muon detectors





Muon detectors

backup

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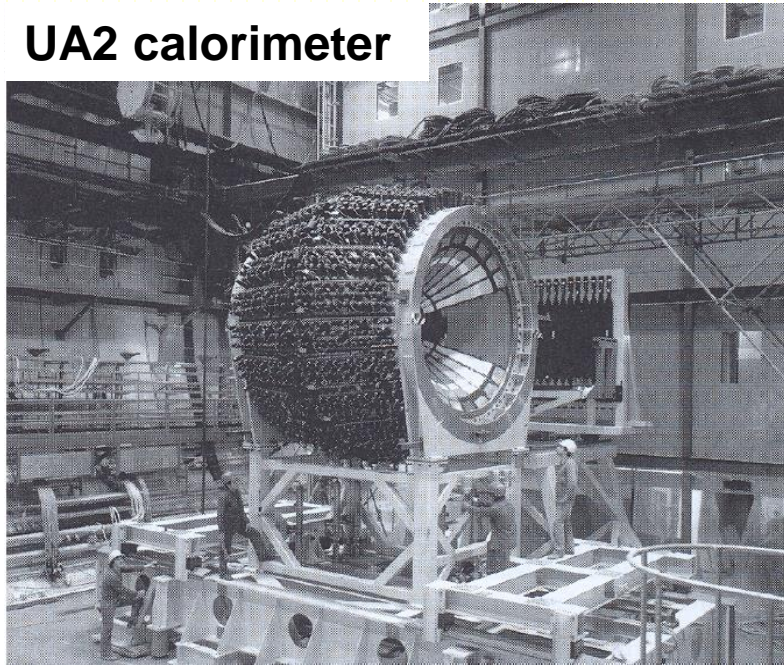
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# Example of calorimeters

**Collider calorimeters** : Geometry is usually more complex, need to cover almost  $4\pi$  solid angle (Missing energy) but also to extract signals. Usually central part with cylindrical geometry (barrel) and small angle part at each end (endcap/forward)

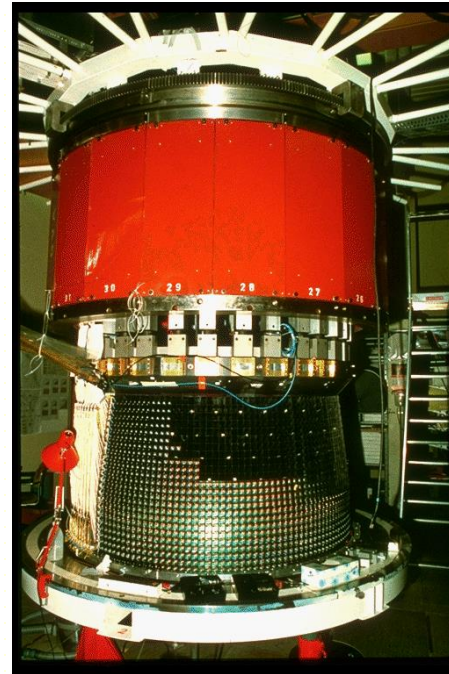
SPS experiments UA1 and UA2

**UA2 calorimeter**



Calorimeter had a crucial role in W/Z discoveries :next slides

LEP experiments :



L3 had a EM calo with excellent energy resolution ( $\gamma$ ) :  
11 000 BGO crystals

But no real impact on main physics topics at LEP

Other experiments (ALEPH, DELPHI and OPAL) put more emphasis on TPC, and Calorimeter granularity



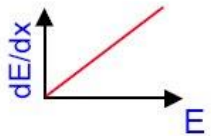
# Electromagnetic shower

$e^+ / e^-$

- Ionisation

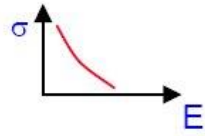


- Bremsstrahlung

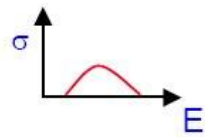


$\gamma$

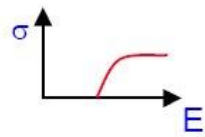
- Photoelectric effect



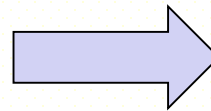
- Compton effect



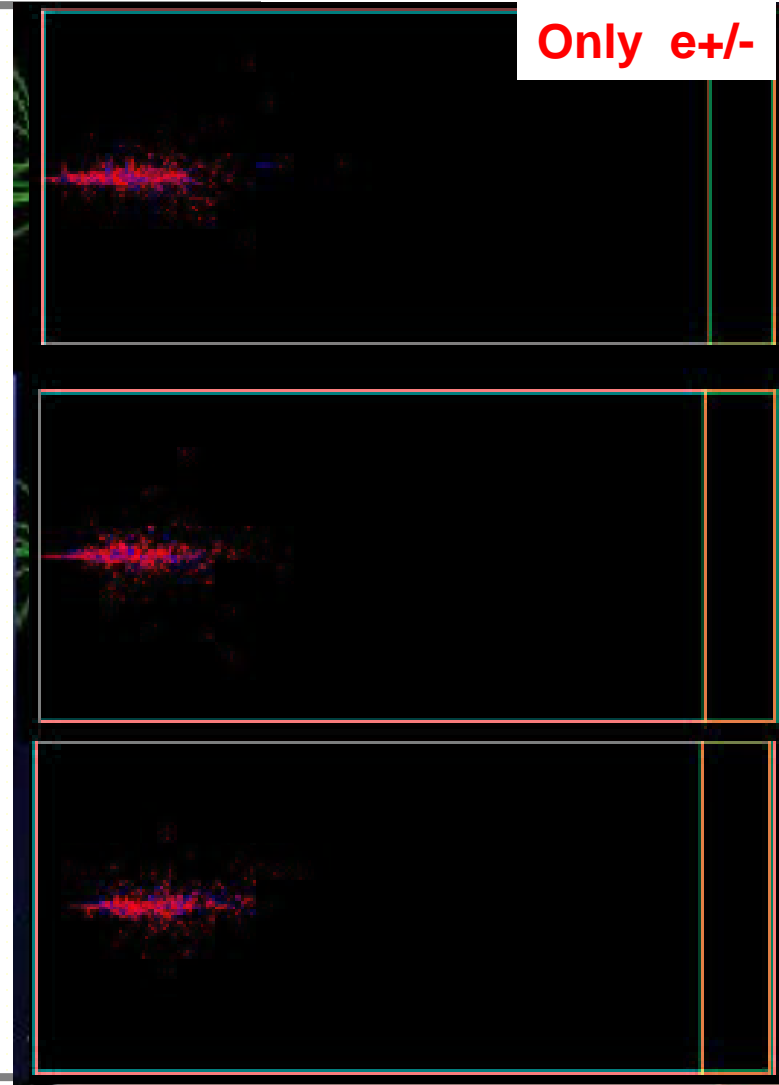
- Pair production



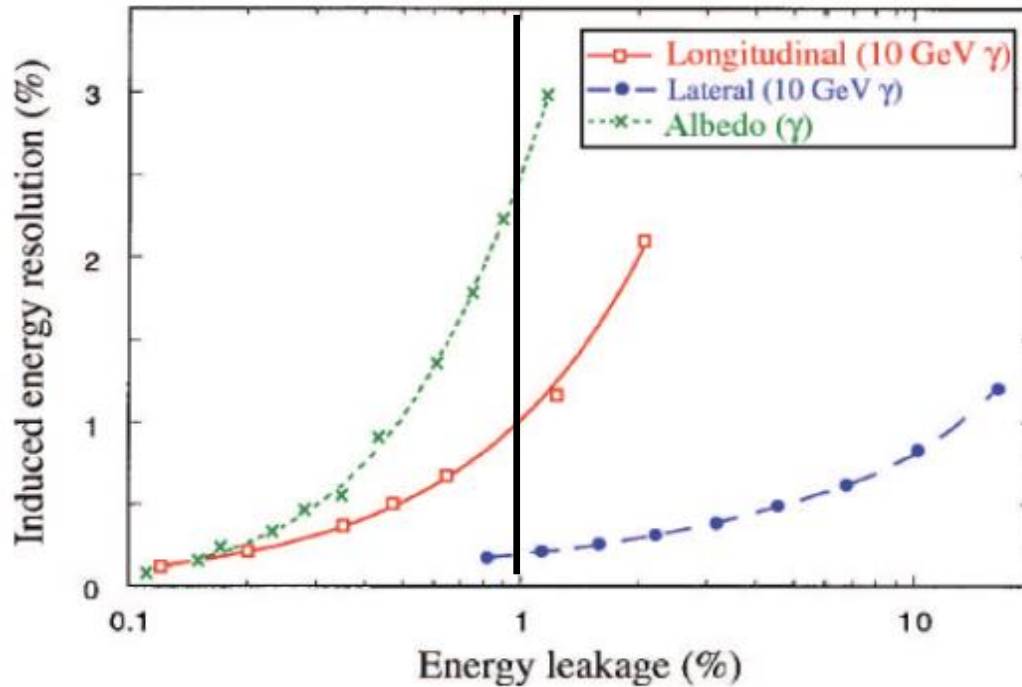
20 GeV  $e^-$



Only  $e^+/-$



# Impact of leakage



Leakage fluctuation usually not poissonian  $\rightarrow$  induces low energy tails  
Longitudinal leakage worsens more the resolution than lateral leakage at fixed value.  
Albedo (back scattering photon) usually dominated by dead material energy loss in front of calorimeter

# Sampling calorimeter with gas

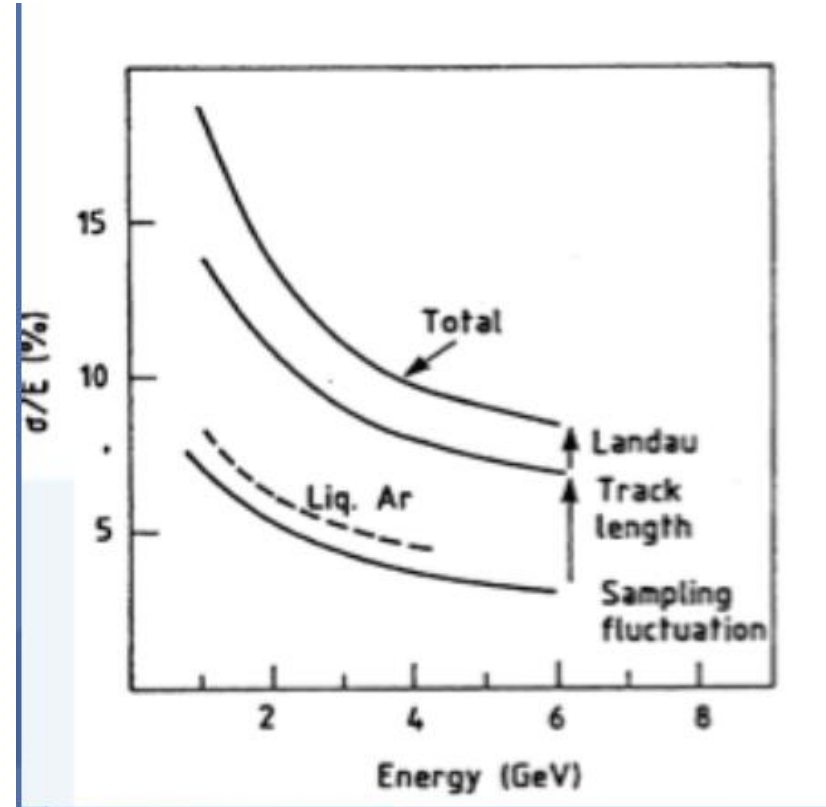
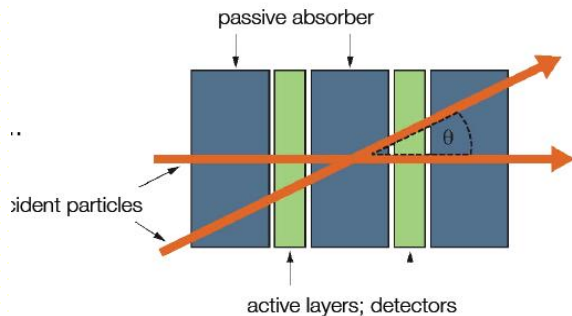
Gas low density medium

Usually poor energy resolution:

- Small sampling fraction (so need larger gap)

+ Track length fluctuation : low electron can travel much in gap

Resolution increases with  $\sqrt{s}$



+ Landau fluctuation

Asymmetric energy deposit in thin active layer (non Gaussian energy measurement)

Calorimeter with gas detector not optimal for good resolution

$$\left[ \frac{\sigma(E)}{E} \right]_{\text{Landau fluctuations}} \propto \frac{1}{\sqrt{N} \ln(k \cdot \delta)} \quad \delta \text{ proportional to density}$$

# From cluster to particle energy (2)

$$E_{reco}^{cal} = a(E_{reco}^{acc}, |\eta|) + b(E_{reco}^{acc}, |\eta|)E_{ps}^{cl LAr} + c(E_{reco}^{acc}, |\eta|)(E_{ps}^{cl LAr})^2 \quad (\text{upstream})$$

$$+ f_{acc}(X, |\eta|) \times (1 + f_{out}(X, |\eta|)) \times \left( \sum_{i=1}^3 E_i^{cl LAr} \right) \left( 1 + f_{leak}(X, |\eta|) \right) \times F(\eta, \varphi)$$

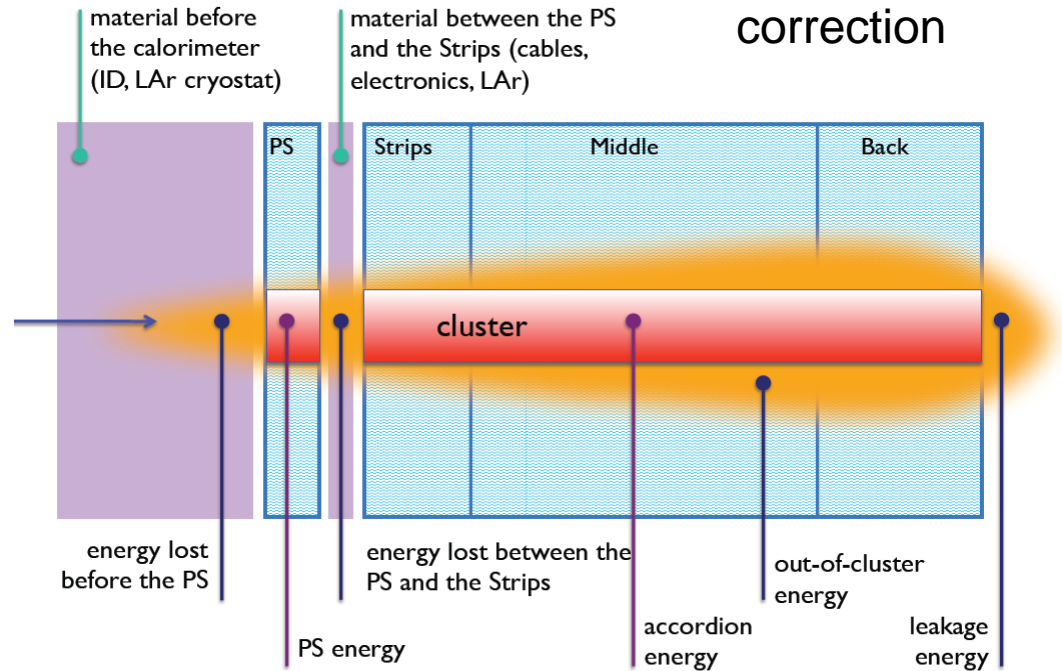
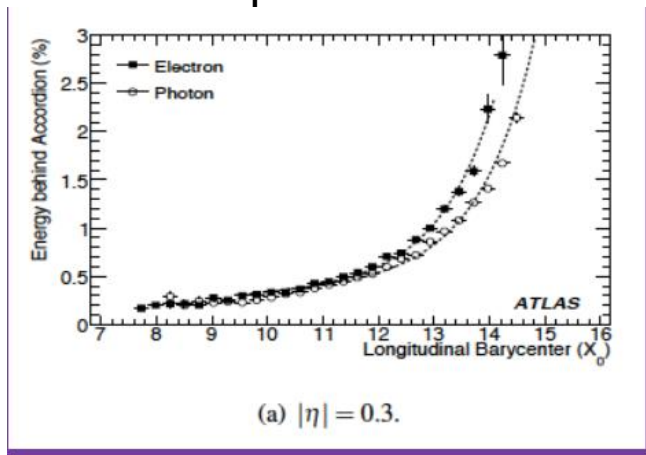
geometry

lateral

longitudinal

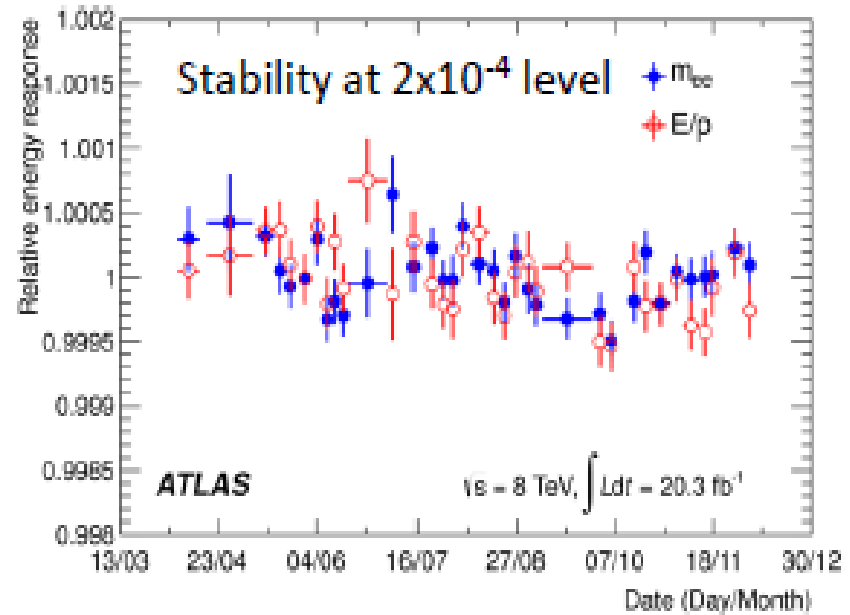
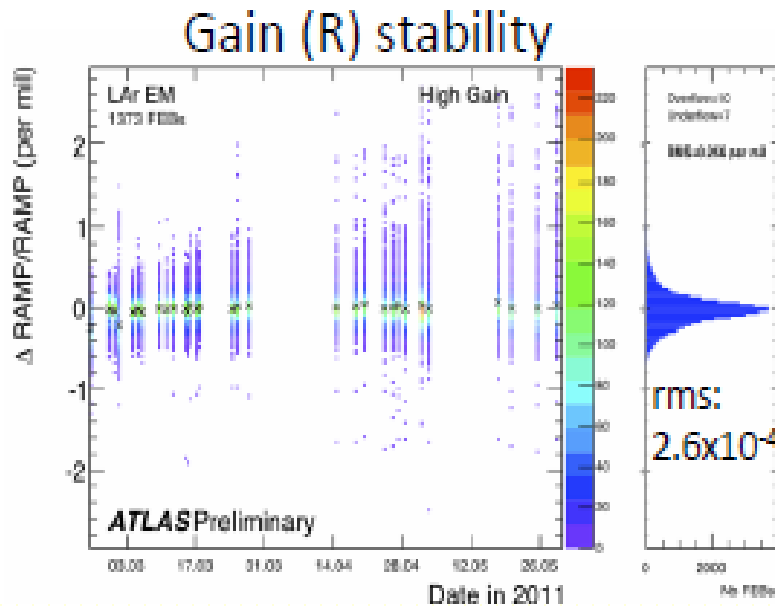
Energy position correction

Parameters/function determined on simulation events, different for electron/photon

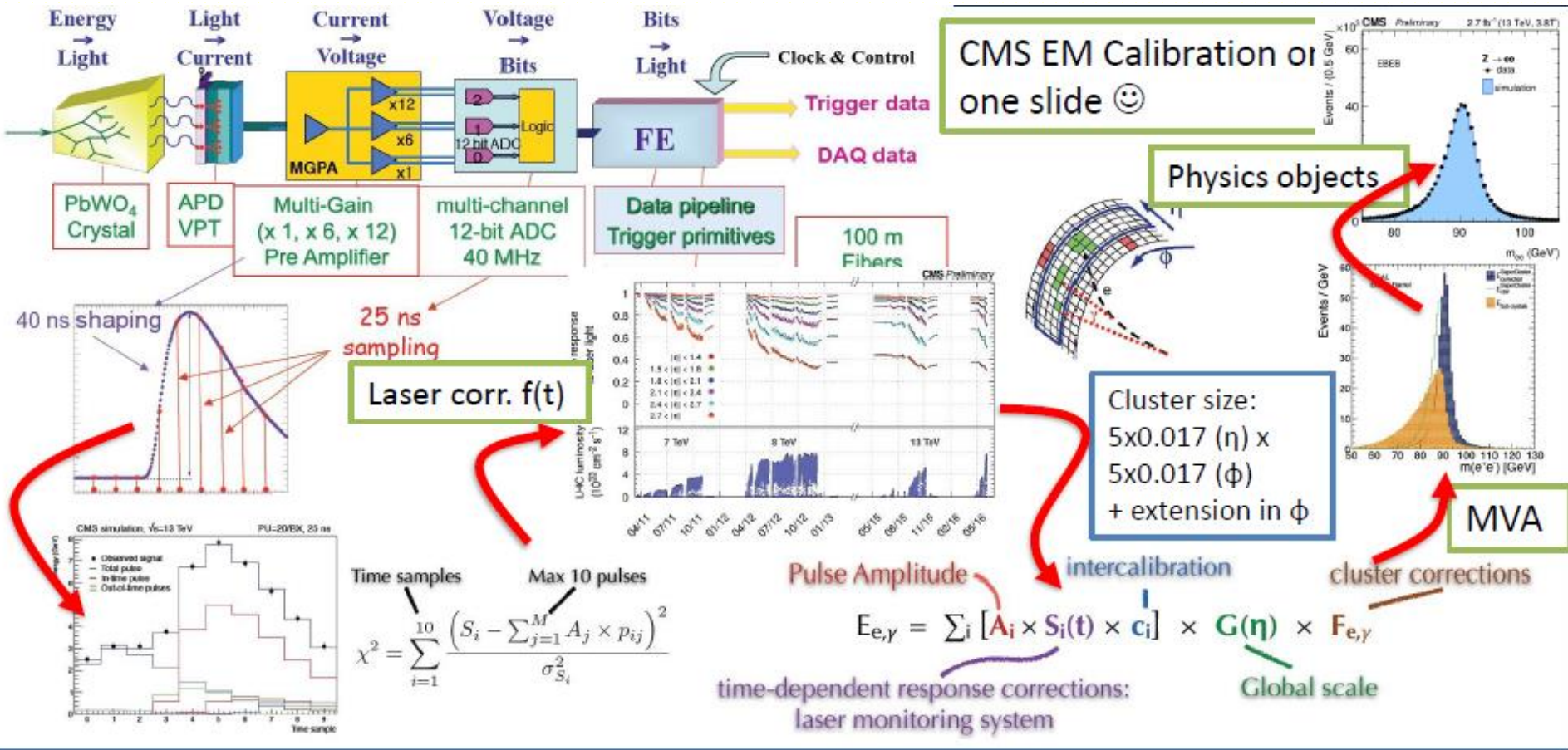


But still not ultimate correlation as detector description/simulation not perfect

# Time stability of ATLAS Calo



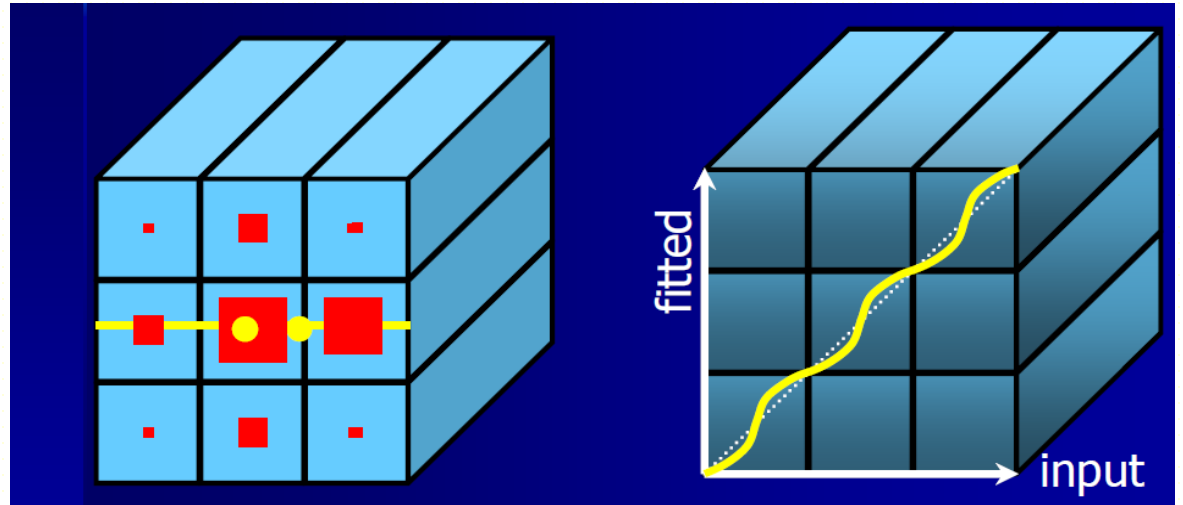
# Calibration not easier in CMS !



# Shower position reconstruction

Energy weighted barycentre

$$E_{rec} = \sum_{i,j} E_{ij}$$
$$x = \frac{1}{E_{rec}} \sum_i x_i \cdot E_i$$
$$y = \frac{1}{E_{rec}} \sum_j y_j \cdot E_j$$



Bias due to finite cell size : S shape  $\rightarrow$  correction to apply

If longitudinal segmentation can also estimate shower depth from

$$X = \sum X_i^0 E_i / E_{rec}$$

From barycentre per layer  $\rightarrow$  Shower direction