

TD3

1) Stress energy tensor

$$1) \nabla_\mu T^{\mu\nu} = 0 = \partial_\mu T^{\mu\nu} + \Gamma^\mu_{\sigma\mu} T^{\sigma\nu} + \Gamma^\nu_{\sigma\mu} T^{\mu\sigma}$$

Have seen in TD2 (first box under eqⁿ (3)) that

$$\Gamma^\alpha_{\alpha\beta} = \frac{1}{\sqrt{-g}} \partial_\beta (\sqrt{-g})$$

$$\begin{aligned} \Rightarrow \nabla_\mu T^{\mu\nu} &= \partial_\mu T^{\mu\nu} + \frac{1}{\sqrt{-g}} \partial_\sigma (\sqrt{-g} T^{\sigma\nu}) + \Gamma^\nu_{\sigma\mu} T^{\mu\sigma} \\ &= \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} T^{\mu\nu}) + \Gamma^\nu_{\sigma\mu} T^{\mu\sigma} \end{aligned}$$

2) Perfect fluid of dust, $P=0 \Rightarrow T^{\mu\nu} = \rho u^\mu u^\nu$

$$\Rightarrow \nabla_\mu (\rho u^\mu u^\nu) = 0 = \nabla_\mu (\rho u^\mu) u^\nu + \rho u^\mu \nabla_\mu u^\nu = 0 \quad \text{--- ①}$$

Contract with u_ν (and use that $u^2 = -1$)

$$\Rightarrow - \nabla_\mu (\rho u^\mu) + \rho u^\mu u_\nu (\nabla_\mu u^\nu) = 0$$

Now since $u^\nu u_\nu = -1$
 $\Rightarrow (\nabla_\mu u^\nu) u_\nu = 0$ } hence $\nabla_\mu (\rho u^\mu) = 0$ --- ②

Substitute ② into ① $\Rightarrow u^\mu \nabla_\mu u^\nu = 0$ --- ③
which is nothing other than the geodesic eqⁿ.

So observe that $\nabla_\mu T^{\mu\nu} = 0$ is telling you about the dynamics of the fluid.

3) In Minkowski space

Want to show that reproduce usual Newtonian eq of m for a fluid (continuity & Euler eq^s)

a) In flat space $\nabla_\mu T^{\mu\nu} = 0 \rightarrow \partial_\mu T^{\mu\nu} = 0$

$$T^{\mu\nu} = (\rho + P/c^2) u^\mu u^\nu + p \eta^{\mu\nu}$$

$$\Rightarrow \left(\partial_\mu \left[(\rho + P/c^2) \right] \right) u^\mu u^\nu + (\rho + P/c^2) \left[(\partial_\mu u^\mu) u^\nu + u^\mu (\partial_\mu u^\nu) \right] + (\partial_\mu p) \eta^{\mu\nu} = 0 \quad \text{--- (4)}$$

Now, from $u^\nu u_\nu = -c^2 \Rightarrow (\partial_\mu u^\nu) u_\nu = 0$

So contract (4) with u_ν

$$\Rightarrow -c^2 \left(\partial_\mu \left[\rho + \frac{P}{c^2} \right] \right) u^\mu + c^2 \left(\rho + \frac{P}{c^2} \right) \left[-\partial_\mu u^\mu \right] + \cancel{(\partial_\mu p) u^\mu} = 0$$

$$\Rightarrow + (\partial_\mu p) u^\mu + \rho (\partial_\mu u^\mu) + \frac{P}{c^2} (\partial_\mu u^\mu) = 0$$

$$\Rightarrow \boxed{\partial_\mu (\rho u^\mu) + \frac{P}{c^2} (\partial_\mu u^\mu) = 0} \quad \text{(5) relativistic continuity eqⁿ}$$

$$= (\partial_\mu p) u^\mu + \rho (\partial_\mu u^\mu) + \frac{P}{c^2} (\partial_\mu u^\mu) = 0$$

(b) - Substitute (5) into (4) : so (4) is

$$u^\nu \left[\underline{(\partial_\mu p) u^\mu} + \underline{(\partial_\mu u^\mu) (\rho + P/c^2)} \right] + \partial_\mu \left(\frac{P}{c^2} \right) u^\mu u^\nu + (\rho + P/c^2) u^\mu (\partial_\mu u^\nu) + (\partial_\mu p) \eta^{\mu\nu} = 0$$

use (5)

$$\Rightarrow \cancel{u^\nu (\partial_\mu u^\mu) \rho} + \partial_\mu \left(\frac{P}{c^2} \right) u^\mu u^\nu + (\rho + P/c^2) u^\mu (\partial_\mu u^\nu) - \cancel{\rho (\partial_\mu u^\mu) u^\nu} + (\partial_\mu p) \eta^{\mu\nu} = 0$$

Spherically symm metrics

- ds^2 can only depend on rotational invariants

$$\left(t, r, d\vec{x}^2 = dr^2 + r^2(d\theta^2 + \sin^2\theta d\phi^2) \quad \& \quad \vec{x} \cdot d\vec{x} = r dr \right)$$

$\Rightarrow ds^2$ takes form in ④:

$$ds^2 = -C(t,r)dt^2 + D(t,r)dr^2 + 2E(t,r)drdt + F(t,r)r^2 d\Omega^2$$

Aim: show that can set $E=0$ & $F=1$ by changes of coord.

1) let $r' = r F(t,r)$

which can invert $\rightarrow r(t,r')$

so that $C(t,r) \rightarrow C'(t,r')$

etc & $dr = \alpha(t,r')dt + \beta(t,r')dr'$

Hence ds^2 of form

$$\Rightarrow ds^2 = -\tilde{C}(t,r')dt^2 + \tilde{D}(t,r')dr'^2 + 2\tilde{E}(t,r')drdt + r'^2 d\Omega^2$$

for some functions \tilde{C}, \tilde{D} etc, which we don't need to write explicitly.

\Rightarrow let's drop the primes & tilda's \Rightarrow

$$ds^2 = -C(t,r)dt^2 + D(t,r)dr^2 + 2E(t,r)drdt + r^2 d\Omega^2. \quad \text{---} \textcircled{*}$$

2) let $dt' = \eta(t,r) [C(t,r)dt - E(r,t)dr]$

where choose η such that this is an exact differential

ie $dt' = \frac{\partial t'}{\partial t} dt + \frac{\partial t'}{\partial r} dr$

$$\text{So } \frac{\partial}{\partial r}(\eta C) = \frac{\partial}{\partial t}(-E\eta) \quad (2)$$

In principle can integrate this to find $\eta(t, r)$ given some $\eta(t_0, r) \forall r$.

$$\text{Then } dt'^2 = \eta^2 \left(C^2 dt^2 - 2CE dt dr + E^2 dr^2 \right)$$

$$\text{So } -\frac{dt'^2}{\eta^2} \frac{1}{C} = -C dt^2 + 2E dt dr - \frac{E^2}{C} dr^2$$

$$\Rightarrow -C dt^2 + 2E dt dr = -\frac{dt'^2}{\eta^2 C} + \frac{E^2}{C} dr^2$$

Substitute into $\otimes \Rightarrow$

$$ds^2 = -\frac{dt'^2}{\eta^2 C} + \left(D + \frac{E^2}{C} \right) dr^2 + r^2 d\Omega^2$$

$$\text{Let } B(t, r) \equiv \frac{1}{\eta^2(t, r) C(t, r)}$$

$$A(t, r) = D(t, r) + \frac{E^2(t, r)}{C(t, r)}$$

& relabel t' by $t \Rightarrow$

$$\boxed{ds^2 = -B(t, r) dt^2 + A(t, r) dr^2 + r^2 d\Omega^2}$$